

A MODERN PHYSICS

PRIMER

FOR ENGINEERING STUDENTS

L. Lanceri
Dept. of Physics
Trieste

Trieste, September - December 2010

A.A. 2010 - 2011

"Complementi di Fisica"

Foreword

I am writing these notes to document a course on "Modern Physics" for engineering students. The main initial motivation was given by the opportunity to establish a physical background on semiconductors, for students of the Electronics curriculum, before they attended a course on solid-state electronic devices. The present scope is somewhat less well-defined, including the interests of engineers approaching the curriculum in Materials Science. Having clarified this, I have to admit that this is not a general introduction to Solid State Physics, nor a full-fledged introduction to Quantum Mechanics, but rather an attempt to introduce a minimum set of modern physics concepts, necessary to understand the potentiality and limits of several present-day technology developments.

The course will touch the following subjects:

Introduction: Particles and Waves: the crisis of Classical Physics.

Part I: An introduction to Quantum Mechanics.

Part II: Some concepts in Solid State Physics.

Part III: Semiconductors at equilibrium.

Part IV: Transport phenomena in semiconductors.

Conclusion: Case study: PV cells.

(c) A shortcut: energy conservation (conservative forces only)

$$E = K + U = \frac{1}{2}mv^2 + U = \frac{p^2}{2m} + U$$

kinetic energy potential energy

Ex.1 :

remember: $\vec{E}$ = force/charge (numerically: force on unit charge)

V = potential energy/charge (---)

in one dimension:

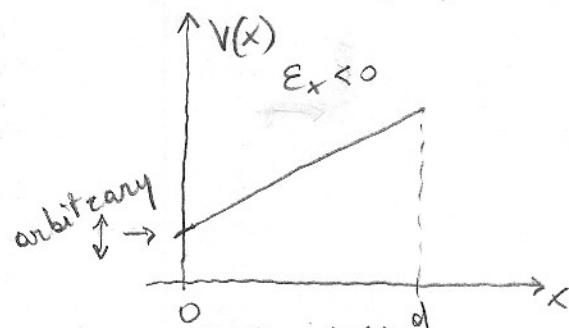
$$V(x) = V(0) - \int_0^x E_x dx \quad ; \quad E_x = - \frac{dV}{dx}$$

work on unit charge from 0 to x

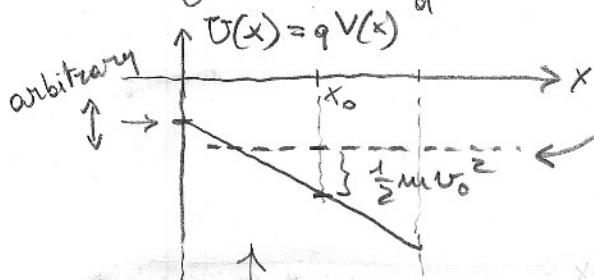
for the electron (charge $q = -|q| < 0$): potential energy

$$U(x) = -|q| V(x) = -|q| V(0) - \int_0^x \underbrace{-|q| E_x}_{\text{force work}} dx$$

in our example, constant $E_x < 0$:



(independent of the charge!
defined in general, for the unit charge)



$$E = U(x_0) + \frac{1}{2}mv_0^2 = \text{const.}$$

an invisible "potential wall"!

depending on the initial x_0 and v_0 , the turning point (if any) of the trajectory is readily found.

Lecture 1

Introduction, then:

Particles and Waves
classical point of view.

1.1. Motivations ; frontiers in physics and technology ;
the electron : particle or wave ? Practical issues.

(see slides)

1.2 Classical particles : mechanics (reminder)



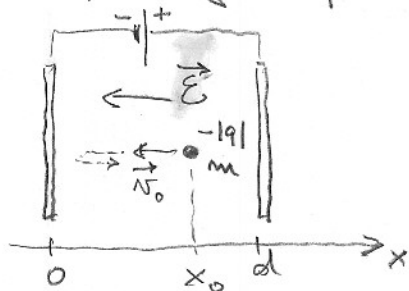
(a) Newton's laws : forces, initial conditions $\Rightarrow$ well defined trajectories by integration

$$\begin{cases} \text{(I)} \dots \\ \text{(II)} \sum \vec{F} = \frac{d\vec{p}}{dt} \\ \text{(III)} \dots \end{cases} \quad \vec{F} = m\vec{a} \quad d\vec{p} = \vec{F}_{net} dt$$

(example : electrostatic field acting on an electron $\Rightarrow$ see slides)

(b) Standard approach : from the differential equation of motion with known forces and initial conditions, to the trajectory, that in principle can be determined with arbitrary precision, depending only on the knowledge of the initial conditions.

Ex.1

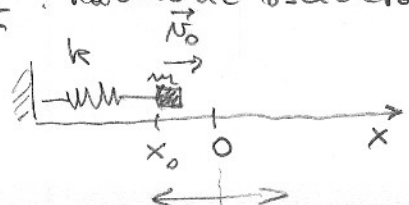


electron in an electrostatic field $\vec{E} = E_x \hat{i}$
electric charge : $q = -|q| < 0$ ($q < 0$)

$$\begin{cases} -|q|E_x = m \frac{d^2x}{dt^2} \\ x(0) = x_0, v_x(0) = v_0 \end{cases} \Rightarrow x(t) = x_0 + v_0 t + \frac{1}{2} a_x t^2$$

another example : mass m , elastic force (Spring : k)

Ex.2 : harmonic oscillator:



$$\begin{cases} -kx = m \frac{d^2x}{dt^2} \\ x(0) = x_0, v_x(0) = v_0 \end{cases}$$

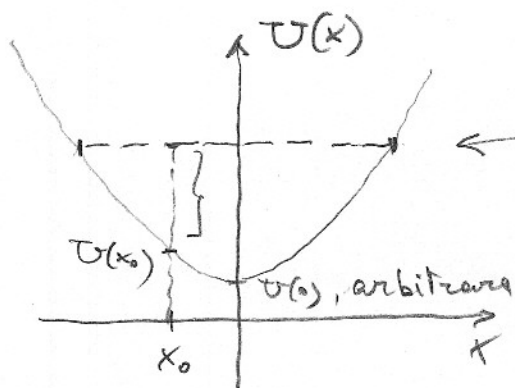
$$\Rightarrow x(t) = A \cos(\omega t + \varphi)$$

$\uparrow$ $A(x_0, v_0)$ $\uparrow$ $\sqrt{k/m}$ $\uparrow$ $\varphi(x_0, v_0)$

Similarly:

Ex. 2: the particle is confined within a "potential well"

Potential energy:
$$U(x) = U(0) - \underbrace{\int_0^x (-kx) dx}_{\text{work}} = U(0) + \frac{1}{2} kx^2$$



$$E = U(x_0) + \frac{1}{2} m v_0^2 = \text{const.}$$

↖ again: an invisible "potential well"!

(d) Note that, for a particle,

$$H \equiv \frac{p^2}{2m} + U \quad \stackrel{\text{def.}}{=} \text{"hamiltonian"}$$

is the basis of further developments in classical mechanics

for ex.: "Hamilton equations" re-formulation of mechanics

$$\left. \begin{array}{l} q : \text{generalized coordinates} \\ p : \text{"momenta"} \end{array} \right\} H = H(p, q)$$

if you are curious about that:

$$\left\{ \begin{array}{l} \frac{dq}{dt} = \dot{q} = \frac{\partial H}{\partial p} \\ \frac{dp}{dt} = \dot{p} = -\frac{\partial H}{\partial q} \end{array} \right.$$

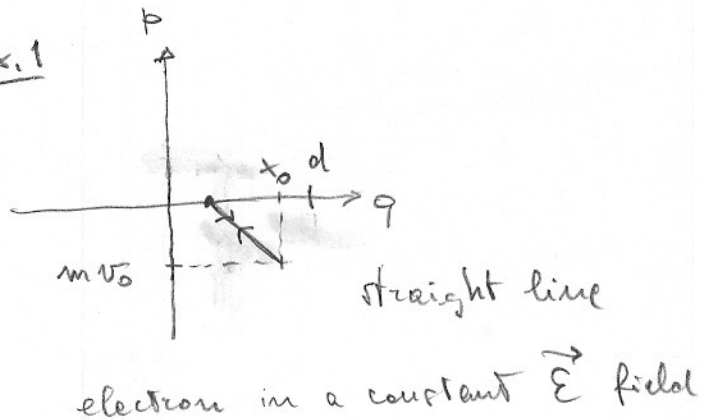
easy to check that this is an equivalent formulation in a specific example, like the harmonic oscillator.

and becomes a fundamental "operator" in quantum mechanics, as we will shortly discover!

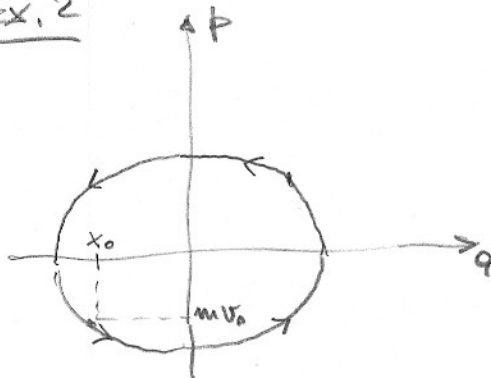
An aside : "phase space", useful representation of the evolution of the state of the system under study, that will reappear later in this course.

In our two simple examples : $q \equiv x$ coordinate
 $p = mv$ momentum

Ex. 1



Ex. 2



$$q \equiv x = A \cos(\omega t + \varphi)$$

$$p \equiv mv = -m\omega A \sin(\omega t + \varphi)$$

it becomes a multi-dimensional space for systems described by a larger number of "generalized coordinates".

1.3. Classical Waves (in mechanics)

Newton's laws can be applied to continuous media (simplest example: a string) to obtain the equations of motion.



(a) equilibrium stretched position.

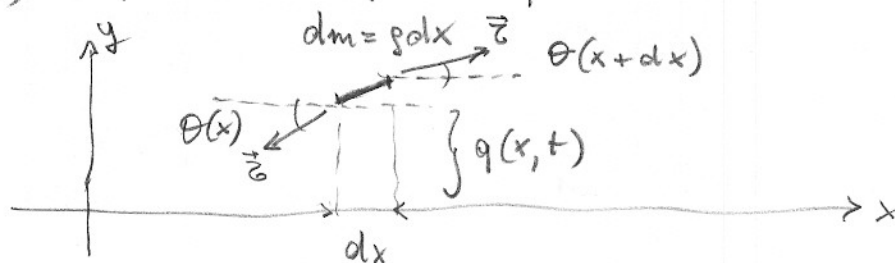
string characterized by:

density $\rho = \frac{dm}{dx}$

tension τ



(b) displacement from equilibrium: (x, y) plane



string element, mass:
 $dm = \rho dx$

$q(x, t)$: displacement in the y direction

Newton's law, projected along y, applied to the string element

$$\sum F_y = \rho dx \left[\frac{d^2 q}{dt^2} \right]_{x \text{ fixed}}$$

net force mass acceleration

$$\tau \cdot [\sin \theta(x+dx) - \sin \theta(x)]_{t \text{ fixed}}$$

linearized assuming, for small θ , $\sin \theta \approx \theta \approx \tan \theta$

$$\Rightarrow \sum F_y = \tau [\tan \theta(x+dx) - \tan \theta(x)]_{t \text{ fixed}} =$$

$$= \tau \left[\frac{dq}{dx} \Big|_{x+dx} - \frac{dq}{dx} \Big|_x \right]_{t \text{ fixed}} = \tau dx \frac{d^2 q}{dx^2} \Big|_{x, t \text{ fixed}}$$

⇒ substituting, we obtain:

$$\cancel{\tau} dx \left[\frac{d^2 q}{dx^2} \right]_{t \text{ fixed}} = \cancel{\rho} dx \left[\frac{d^2 q}{dt^2} \right]_{x \text{ fixed}}$$

in other words:

$$(c) \quad \frac{\partial^2 q(x,t)}{\partial x^2} = \frac{1}{u^2} \frac{\partial^2 q(x,t)}{\partial t^2}$$

with $u = \sqrt{\frac{\tau}{\rho}}$ is a velocity!

(dimensionless, easy to check)

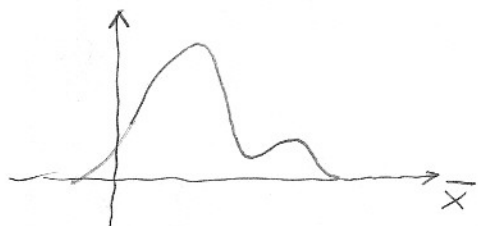
very important result! The transverse motion of the string is governed by a 2nd order diff. equation (linearized for small displacements) relating the second partial derivatives w.r. to x and t .

and: "velocity" ^{squared} appearing as a coefficient:

$$u^2 \propto \tau \text{ (tension)}, \quad u^2 \propto \frac{1}{\rho} \text{ (density)}$$

Let us look at the solutions of this equation!

(d) "D'Alembert solution"



$f(\bar{x})$ arbitrary function of $\bar{x}$,
twice differentiable

define $\bar{x} \equiv x - ut$

D'Alembert solution: $f(\bar{x}) = f(x - ut)$

easy to check:

$$\frac{\partial f}{\partial x} = \frac{df}{d\bar{x}} \frac{\partial \bar{x}}{\partial x} = \frac{df}{d\bar{x}} = f'(\bar{x})$$

$$\frac{\partial^2 f}{\partial x^2} = \frac{d^2 f}{d\bar{x}^2} \frac{\partial \bar{x}}{\partial x} = \dots = f''(\bar{x})$$

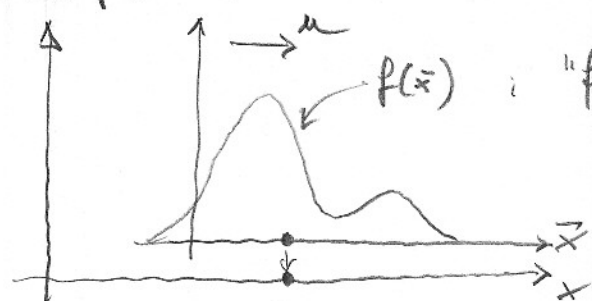
$$\frac{\partial f}{\partial t} = \frac{df}{d\bar{x}} \frac{\partial \bar{x}}{\partial t} = f'(\bar{x}) \cdot (-u)$$

$$\frac{\partial^2 f}{\partial t^2} = \dots = f''(\bar{x}) \cdot (-u)(-u) = u^2 f''(\bar{x})$$

⇒ OK.

⇒ any function $q(x, t) = f(x - ut)$ is a good solution!
 (differentiable twice)

Interpretation?



$\bar{x} = x - ut$
 moving with speed u
 with respect to x
 $(x = \bar{x} + ut)$

The "pulse" $f(\bar{x}) = f(x - ut)$ moves without change in shape with velocity u in the x direction.

Easy to check that changing $u \rightarrow -u$ simply reverses the direction of the motion
 $(x - ut) \rightarrow (x + ut)$

and: the equation is linear ⇒ accepts linear combinations of solutions

$$\Rightarrow \boxed{q(x, t) = f(x - ut) + g(x + ut)}$$

pulse → ← pulse

"D'Alembert solution"

LH 19.09.2010

LØ2/4

Good to see in practice what this means :
demo with a real string/rpe.

...

Next lecture :

let's get comfortable with sinusoidal waves and
their properties, and with superpositions of sinusoidal
waves; standing waves in particular.

Then:

more on to a reminder of

- Statistical mechanics (and thermodynamics ?)
 - electromagnetism.
-

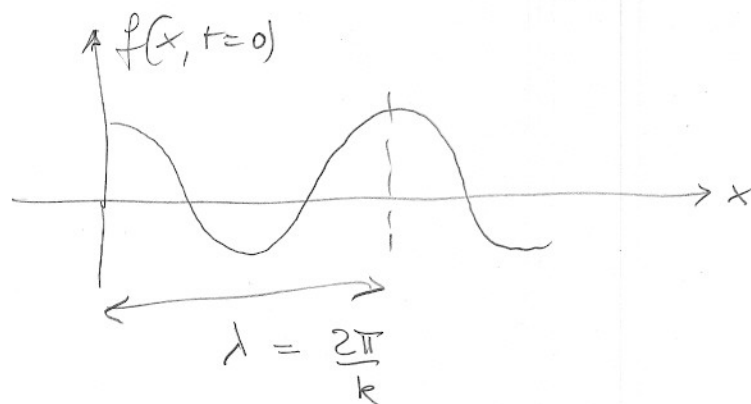
Waves: solutions of the wave equation.

sinusoidal waves { periodic solutions "excited by harmonic oscillator"
can reproduce arbitrary periodic solutions by superposition

$$f(\vec{x}) = \cos(k(x - ut)) = \cos\left(kx - \overbrace{kut}^{\omega}\right)$$

$\underbrace{[L]^{-1}}_{[L]} \underbrace{[L]}_{[L]} \quad \underbrace{[T]^{-1}}_{[T]} [T]$

(a) snapshot at fixed t (for instance $t=0$)

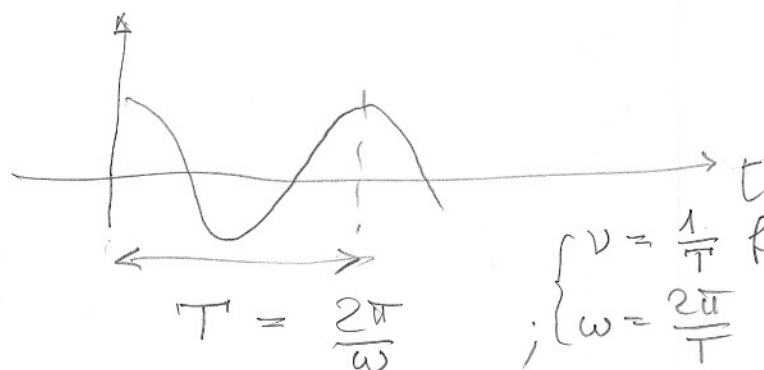


$$f(x, t=0) = \cos kx$$

"period" $\equiv$
 $\equiv$ "wavelength"

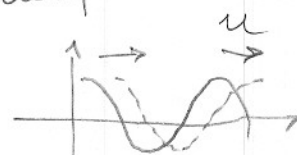
$$k = \frac{2\pi}{\lambda} \quad [m^{-1}] \quad \text{"wave number"}$$

(b) Time evolution at fixed x (for instance at $x=0$)



$$f(x=0, t) = \cos(\omega t)$$

$$\left\{ \begin{array}{l} \nu = \frac{1}{T} \text{ frequency} \\ \omega = \frac{2\pi}{T} \end{array} \right.$$



(c) u : "phase velocity" (for instance: displacement of max in unit time)

$$k u = \omega \Rightarrow u = \frac{\omega}{k} = \frac{2\pi}{T} \cdot \frac{\lambda}{2\pi} = \frac{\lambda}{T} = \nu \lambda$$

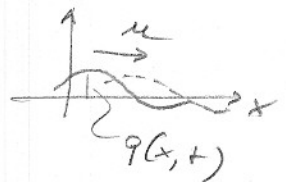
$\lambda = u T, \quad T = \frac{\lambda}{u}$

useful relations
easy to remember (dim.)

OK. we found that mechanical ^{on a string} waves propagate according to:

$$\frac{\partial^2 q}{\partial x^2} = \frac{1}{u^2} \frac{\partial^2 q}{\partial t^2} \Rightarrow f(x-ut) + g(x+ut) = q(x,t)$$

d'Alembert solution



sinusoidal waves: example

$$q(x,t) = A \cos(kx - \omega t)$$

[complex exponential form

$$A e^{i(kx - \omega t)}]$$

$$u = \frac{\omega}{k} \quad \left\{ \begin{array}{l} \lambda = \frac{2\pi}{k} \\ T = \frac{2\pi}{\omega} \\ = \frac{\lambda}{u} = \frac{1}{\nu} \end{array} \right.$$

numerical example for the string: u is fixed by $\left\{ \begin{array}{l} \rho \text{ density} \\ \tau \text{ tension} \end{array} \right.$

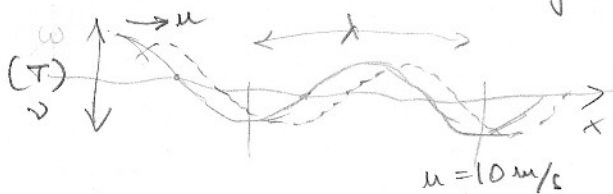
$$u = \sqrt{\frac{\tau}{\rho}}$$

$$[u] = \left[\cancel{M} \right]^{\frac{1}{2}} \left[L \right]^{\frac{1}{2}} \left[T \right]^{-\frac{1}{2}} \cdot \left[\cancel{M} \right]^{-\frac{1}{2}} \left[L \right]^{\frac{1}{2}} = [L][T]^{-1}$$

$$\left. \begin{array}{l} \rho = 0,10 \text{ kg/m} \\ \tau = 10 \text{ N} \end{array} \right\} \Rightarrow u = \sqrt{10^2} \approx 10 \text{ m/s}$$

in this model, u is constant (fixed) for fixed ρ and τ ; does not depend on ω or k .

$\Rightarrow$ if Δ shake one end of the string with a sinusoidal time dependence, Δ will generate sinusoidal waves, propagating at phase velocity $u \approx 10 \text{ m/s}$, and wavelength $\lambda = uT$



$$\begin{array}{lll} \Rightarrow T = 1 \text{ s} & \longrightarrow & \lambda = 10 \text{ m} \quad (\nu = \frac{1}{T} = 1 \text{ Hz}) \\ T = 0,1 \text{ s} & \longrightarrow & \lambda = 1 \text{ m} \quad \left\{ \begin{array}{l} 0,610 \text{ Hz} \\ 100 \text{ Hz} \end{array} \right. \\ T = 0,01 \text{ s} & \longrightarrow & \lambda = 0,1 \text{ m} \end{array}$$

NB u is different if Δ change the tension τ of the same rope (fixed ρ)

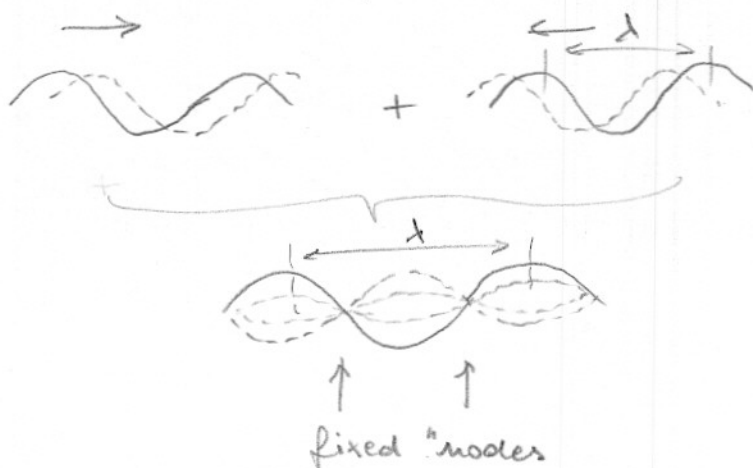
from Taylor - Zefiretos. Dubson, p. 205

$$q(x,t) = \vec{B} \sin(kx - \omega t) + \overset{\leftarrow}{B} \sin(kx + \omega t) =$$

↑ same amplitude ↑

Trigonometry: $\sin a + \sin b = 2 \sin \frac{a+b}{2} \cos \frac{a-b}{2}$

$$= \underbrace{(2B)}_{\text{"envelope" of amplitudes}} \sin kx \cos \omega t \quad \text{NB:}$$



two oppositely travelling waves
(same $k, \omega, \lambda, \dots$)

make a "standing wave"

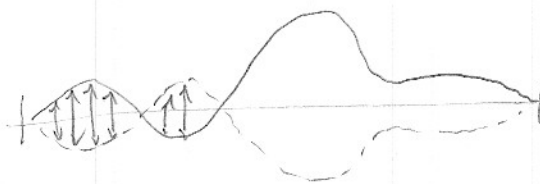
this is a typical configuration realized when a string of given length L is clamped at its ends.

Let's analyze this more quantitatively.

Let's discuss now the "normal modes"

("standing waves")
see L03/2 bis

- ① $\Rightarrow$ special case : oscillation with all points oscillating with the same frequency
(but possibly with a different phase)



Solutions with factorized x and t dependence

(practical examples : "stationary, standing waves on a string")

- ② $q(x, t) = q(x) \cos(\omega t + \delta)$
- Annotations:
- $q(x)$: amplitude is different for different points
 - $\cos(\omega t + \delta)$: initial phase could be different for different points
 - ωt : all points oscillate "in time"

- ③ substituting:

$$\frac{\partial^2 q}{\partial x^2} = \frac{1}{u^2} \frac{\partial^2 q}{\partial t^2}$$

$$\frac{d^2 q(x)}{dx^2} \cos(\omega t + \varphi) = \frac{1}{u^2} q(x) (-\omega^2) \cos(\omega t + \varphi)$$

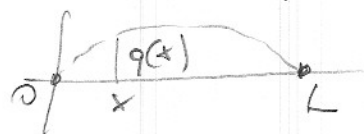
④

$$\boxed{\frac{d^2 q}{dx^2} = - \underbrace{\left(\frac{\omega^2}{u^2} \right)}_{k^2} q(x)}$$

$$q = q(x) !$$

"Helmholtz equation"

① Interesting case: fixed endpoints
 $q(0) = q(L) = 0$



general solutions:

$$q(x) = A \sin kx + B \cos kx$$

$$q(0) = 0 \Rightarrow B = 0$$

$$q(L) = 0 \Rightarrow A \sin kL = 0$$

$$\Leftrightarrow k_n L = n\pi$$

$$k_n = n \frac{\pi}{L}$$

$$n = 1, 2, \dots, \infty$$

"eigenvalues"

"eigenfunction"

(of the operator $\frac{d^2}{dx^2}$)

② $q_n(x) = A \sin k_n x$

to obtain
 $\int_0^L q_n(x) dx = 1$

$$\Rightarrow q_n(x) = \sqrt{\frac{2}{L}} \sin\left(\frac{n\pi}{L} x\right)$$

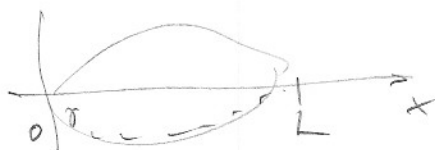
③ $\Rightarrow q_n(x, t) = \sqrt{\frac{2}{L}} \sin\left(\frac{n\pi}{L} x\right) \cos(\omega_n t + \delta_n)$

"normal modes"

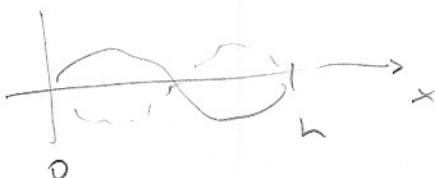
$$\omega_n = n k_n = n \frac{\pi}{L} v$$

phase velocity

examples: $n=1$



$n=2$



$$\left[\begin{array}{c} (\delta_n = 0) \\ \uparrow \\ \text{NB! ?} \end{array} \right]$$

④ These "factorized" solutions are also called "normal modes" of the string with fixed endpoints.

(ω_n, δ_n are constants to be determined by the INITIAL CONDITIONS of the motion)

① generic solution: linear superposition of "normal modes"

$$q(x,t) = \sum_1^\infty A_n \sqrt{\frac{2}{L}} \sin\left(\underbrace{n\frac{\pi}{L}x}_{k_n}\right) \cos\left(\underbrace{\omega_n t + \delta_n}_{\left(n\frac{\pi}{L}\right)u = k_n u}\right)$$

(satisfies Huygens eq. with the given boundary conditions, fixed endpoints)

② specific solution: initial conditions for

$$\begin{cases} q(x, t=0) \\ \frac{\partial q}{\partial t} \Big|_{t=0} \end{cases} \Rightarrow \text{fixes } A_n, \delta_n$$

initial "shape" of the string
initial transverse velocities

that is:

$$q(x, 0) = \sum_1^\infty \underbrace{A_n \cos \delta_n}_{a_n} \underbrace{\sin\left(n\frac{\pi}{L}x\right) \cdot \sqrt{\frac{2}{L}}}_{\phi_n(x)} = f(x)$$

$$\frac{\partial q}{\partial t} \Big|_{t=0} = - \sum_1^\infty \underbrace{\omega_n A_n \sin \delta_n}_{\omega_n a_n} \underbrace{\sin\left(n\frac{\pi}{L}x\right) \cdot \sqrt{\frac{2}{L}}}_{\phi_n(x)}$$

Fourier series !!!

give functions
"initial conditions"

it can be shown that the coeff. A_n, δ_n are fixed by these conditions.

③ To solve for A_n, δ_n one can use the following property:

$$\int_0^L \phi_m(x) \phi_n(x) dx = \delta_{mn} \begin{cases} 0 & n \neq m \\ 1 & n = m \end{cases}$$

$$\Rightarrow \int_0^L \phi_m(x) \cdot f(x) dx = \sum_1^\infty a_n \underbrace{\int_0^L \phi_m(x) \phi_n(x) dx}_{\delta_{mn}} = a_m$$

④ in practice, this means that the recipe to compute A_m and δ_m is:

$$\forall m \begin{cases} a_m = A_m \cos \delta_m = \int_0^L \underbrace{q(x, 0)}_{f(x)} \cdot \underbrace{\sqrt{\frac{2}{L}} \sin\left(n\frac{\pi}{L}x\right)}_{\phi_m(x)} dx = \dots \\ -\omega_m a_m = -\omega_m A_m \cos \delta_m = \int_0^L \underbrace{\frac{\partial q}{\partial t} \Big|_{t=0}}_{\dots} \cdot \sqrt{\frac{2}{L}} \sin\left(n\frac{\pi}{L}x\right) dx = \dots \end{cases}$$

explicit solution!

number that can be computed for the given $q(x, 0)$

① In summary:

every specific solution can be written as an (infinite) sum of "normal modes", with coefficients that can be computed from the initial configuration ($q(x, t=0)$ and $\frac{\partial q}{\partial t}|_{t=0}$).

② for a long string: $L \rightarrow \infty$

$$\Rightarrow k_n = \frac{\pi}{L} n : \Delta k = k_n - k_{n-1} = \frac{\pi}{L} \rightarrow 0$$

the spacing of wave numbers becomes vanishing small.
the "Fourier sum" becomes a "Fourier integral"
 $\left(\sum_{n=-\infty}^{\infty} \right) \quad \frac{1}{\pi} \int_0^{\infty} dk$

③ why?

for any function $f(k_n)$

$$\begin{aligned} \sum_{n=-\infty}^{\infty} f(k_n) &= \frac{L}{\pi} \sum_{n=-\infty}^{\infty} \cdot \frac{\pi}{L} f\left(\frac{\pi n}{L}\right) \\ &\quad \underbrace{\qquad\qquad\qquad}_{\substack{\parallel \\ \sum_{n=-\infty}^{\infty} \Delta k f(k_n)}} \\ &\quad \downarrow L \rightarrow \infty \\ &\int_0^{\infty} dk f(k) \end{aligned}$$

④

in other terms:

$$\sum_{n=-\infty}^{\infty} \xrightarrow{L \rightarrow \infty} \frac{L}{\pi} \int_0^{\infty} dk$$

and if $f(k) = f(-k)$, even function, for $L \rightarrow \infty$

$$\sum_{n=-\infty}^{\infty} \xrightarrow{L \rightarrow \infty} \frac{L}{2\pi} \int_{-\infty}^{+\infty} dk$$

"Fourier integrals"

A quick reminder of electrodynamics

① We use mainly S.I. units.

$$\left\{ \begin{array}{l} \text{mass: kg} \\ \text{length: m} \\ \text{time: s} \\ \text{Force: } 1 \text{ N} = 1 \text{ kg} \cdot \text{m/s}^2 \end{array} \right.$$

$$\left\{ \begin{array}{l} \text{charge: C} \\ \text{voltage: V} \\ \text{current: A} \quad 1 \text{ A} = 1 \text{ C/s} \end{array} \right.$$

② Basic principles:

$$q_1 \bullet \xrightarrow{\vec{r}} \bullet q_2$$

$$\vec{r} = \vec{r}_2 - \vec{r}_1$$

$$\vec{F}_{21} = \frac{1}{4\pi\epsilon_0} \frac{q_1 q_2}{r^2} \left(\frac{\vec{r}}{r} \right)$$

on particle 2 by particle 1

Coulomb's law

$$8.988 \times 10^9 \frac{\text{N m}^2}{\text{C}^2}$$

$$\left(\vec{E}_1 = \frac{\vec{F}_{21}}{q_2} = \frac{1}{4\pi\epsilon_0} \frac{q_1}{r^2} \hat{r} \right) \leftarrow \frac{\text{N}}{\text{C}} = \frac{\text{V}}{\text{m}}$$

③

magnetostatic field
 $\vec{B}$

$$\left[\text{diagram of solenoid with current } i \text{ and magnetic field } \vec{B} \text{ inside} \right] \Rightarrow$$

simpler expression for the field generated by a solenoid:

$$\vec{B} = \mu_0 n i$$

$$\left[\text{diagram of infinite solenoid with magnetic field } \vec{B} \text{ inside} \right]$$

"infinite solenoid" $\left\{ \begin{array}{l} n \text{ turns/m} \\ \text{current } i \text{ (A)} \end{array} \right.$

defined as

$$\frac{\mu_0}{2\pi} \approx 1,000 \times 10^{-7} \frac{\text{N s}^2}{\text{C}^2}$$

$$= 1,000 \times 10^{-7} \frac{\text{T m}}{\text{A}}$$

B measured in Tesla (T)

$$\left(1 \text{ T} = \frac{1 \text{ kg}}{\text{C s}} = 10^4 \text{ gauss} \right)$$

④

Force on a charged particle

"Lorentz"

$$\vec{F} = q(\vec{E} + \vec{v} \times \vec{B})$$

(a) An example of how a ^{charged} particle interacts with $\vec{E}$ and $\vec{B}$ fields in an experimental apparatus:

①



$\vec{B} \perp$ drawing (entering) ; $\vec{E} = 0$
 uniform
 $q > 0$ (for instance: proton)

$$\vec{F} = q(\vec{v} \times \vec{B}) = \frac{d\vec{p}}{dt} = m\vec{a} \quad (\text{non relativistic})$$

$$qvB = m \frac{v^2}{r}$$

⊗ relativistic:
 $\vec{p} = \gamma m_0 \vec{u}$

→ circular motion, radius: $r = \frac{mv}{qB} \Rightarrow r \propto p$

②

orders of magnitude (non-relativistic)

proton: $\begin{cases} m_p = 1.6 \times 10^{-27} \text{ kg} \\ v = 3.0 \times 10^7 \text{ m/s} \quad (1/10 \text{ speed of light}) \\ q = 1.6 \times 10^{-19} \text{ C} \end{cases}$

mag. field: $B = 0.1 \text{ T} = 10^3 \text{ gauss}$

$$r \approx \frac{1.6 \times 10^{-27} \times 3 \times 10^7}{1.6 \times 10^{-19} \times 10^{-1}} = 3 \times \frac{10^{-20}}{10^{-20}} \approx 3 \text{ m}$$

($r \approx 0.3 \text{ m}$ if $B \approx 1 \text{ T}$)

NB: in this example

$$v = 0.1 c$$

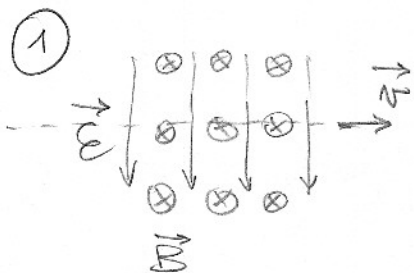
↑
speed of light
in vacuum

$$\Rightarrow p = \gamma m v = \frac{m}{\sqrt{1 - (0.1)^2}} v \approx m v$$

↑
 $\gamma \approx 1$

for relativistic particles ($v \approx c$) $\vec{F} = \frac{d\vec{p}}{dt}$ remains the same,
 but $\vec{p} = \gamma m \vec{v}$, $\gamma = \frac{1}{\sqrt{1 - v^2/c^2}}$

(b) Another example: a "velocity selector" with static fields $\vec{E} \perp \vec{B}$



a particle (m, q) entering this space region is subject to a force

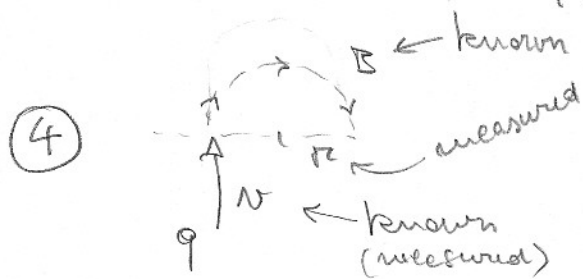
$$\vec{F} = q (\vec{E} + \vec{v} \times \vec{B})$$

(2) in this configuration, the two contributions can be opposite
 $(q\vec{E} - q\vec{v} \times \vec{B}) \Rightarrow \vec{F} = 0 \Leftrightarrow E = vB$
 $v = \frac{E}{B}$

for instance: $\left\{ \begin{array}{l} E = 10^3 \text{ V/cm} = 10^5 \text{ V/m} \\ B = 1 \text{ T} \end{array} \right\}$ selects $v = 10^5 \text{ m/s}$

("small" with respect to $c = 3 \times 10^8 \text{ m/s}$)

(c) a combination of (b) velocity selector and (a) magnetic field.. makes a "mass spectrograph"



$$m = \left(\frac{qB}{v} \right) r$$

known
measured

← experimental mass determination.

(d) classic circular motion in the Coulomb el. field of a nucleus:

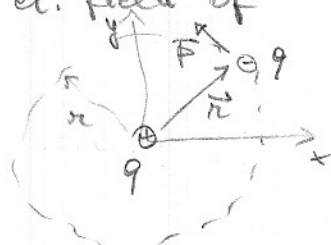
(1) Newton's law $\rightarrow m\vec{a} = \vec{F}$

kinematics $\rightarrow m \frac{v^2}{r} = \frac{q^2}{4\pi\epsilon_0} \frac{1}{r^2}$

(2) Angular momentum $\vec{l} = \vec{r} \times \vec{p} = (r m v) \hat{z}$

$$\frac{m^2 v^2 r^2}{m} = \frac{l^2}{m} = \left(\frac{q^2}{4\pi\epsilon_0} \right) r$$

$r \propto l^2$



Maxwell's equations

The description of the interactions between charged particles (static or in motion $\rightarrow$ currents) is much simpler if we think in terms of:

"Sources" $\rightarrow$ fields $\rightarrow$ forces on particles/currents

ch. density $\rho(\vec{r}, t)$

$\vec{E}, \vec{B}$

curr. density $\vec{J}(\vec{r}, t)$

[divergence]

$$\vec{\nabla} \cdot \vec{E} = \frac{\rho}{\epsilon_0}$$

$$\vec{\nabla} \cdot \vec{B} = 0$$

[curl]

$$\vec{\nabla} \times \vec{E} = -\frac{\partial \vec{B}}{\partial t}$$

$$\vec{\nabla} \times \vec{B} = \mu_0 \vec{J} + \mu_0 \epsilon_0 \frac{\partial \vec{E}}{\partial t}$$

+ "charge conservation"

$$\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot \vec{J} = 0$$

in cartesian coordinates:

$$\vec{v} = (v_x \hat{i} + v_y \hat{j} + v_z \hat{k}) \text{ any vector}$$

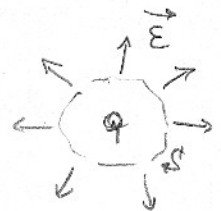
$$\vec{\nabla} = \left(\frac{\partial}{\partial x} \hat{i} + \frac{\partial}{\partial y} \hat{j} + \frac{\partial}{\partial z} \hat{k} \right)$$

$$\vec{\nabla} \cdot \vec{v} = \frac{\partial v_x}{\partial x} + \frac{\partial v_y}{\partial y} + \frac{\partial v_z}{\partial z} \quad \text{"divergence" (scalar)}$$

$$\vec{\nabla} \times \vec{v} = \left(\frac{\partial v_y}{\partial z} - \frac{\partial v_z}{\partial y} \right) \hat{i} + \left(\frac{\partial v_z}{\partial x} - \frac{\partial v_x}{\partial z} \right) \hat{j} + \left(\frac{\partial v_x}{\partial y} - \frac{\partial v_y}{\partial x} \right) \hat{k} \quad \text{"curl" (vector)}$$

integral form

$$\oint_S \vec{E} \cdot \hat{n} dS = \frac{Q}{\epsilon_0}$$

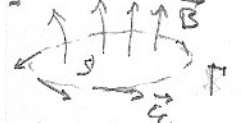


$$\oint_S \vec{B} \cdot \hat{n} dS = 0$$

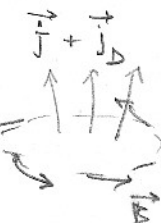


[line integral, closed path]

$$\oint_C \vec{E} \cdot d\vec{l} = -\frac{d}{dt} \int_S \vec{B} \cdot \hat{n} dS$$



$$\oint_C \vec{B} \cdot d\vec{l} = \mu_0 \int_S \vec{J} \cdot \hat{n} dS + \mu_0 \epsilon_0 \int_S \frac{\partial \vec{E}}{\partial t} \cdot \hat{n} dS$$



$$\oint_S \vec{J} \cdot \hat{n} dS = -\frac{dQ}{dt} = -\frac{d}{dt} \int_V \rho dV$$

locally $\neq 0$ if:

L 27.09.2010

LPH/2

e.m. waves

: from the Maxwell equations:

[NB: can explain in reverse order]



it can be shown that in an empty region (no charges, no currents)

② e.m. fields generated outside can propagate for instance in the x direction, obeying the following wave equation, derived from Maxwell's equations that connect $\vec{E}$ and $\vec{B}$ (time-varying $\vec{B}$ field is a source of $\vec{E}$ field and vice-versa)

③ wave equations for E_y, B_z

$$\begin{aligned} \vec{B} &= B_z \hat{k} \\ \vec{E} &= E_y \hat{j} \end{aligned} \quad \begin{array}{c} \nearrow x \rightarrow \hat{i} \\ \nwarrow \end{array}$$

$$\left\{ \begin{aligned} \frac{\partial^2 E_y}{\partial x^2} &= \mu_0 \epsilon_0 \frac{\partial^2 E_y}{\partial t^2} \\ \frac{\partial^2 B_z}{\partial x^2} &= \mu_0 \epsilon_0 \frac{\partial^2 B_z}{\partial t^2} \end{aligned} \right.$$

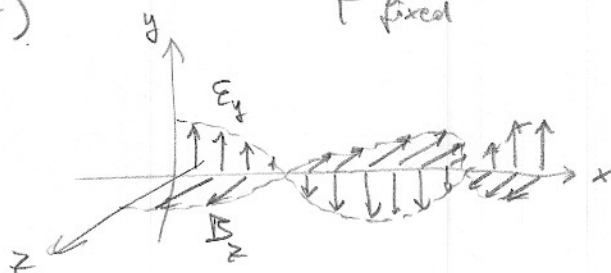
$$\frac{1}{c^2} = \mu_0 \epsilon_0$$

$$c = \frac{1}{\sqrt{\mu_0 \epsilon_0}} = 2.998 \times 10^8 \text{ m/s}$$

④ solutions ("plane wave" of Maxwell's and wave equation)

$$\begin{cases} E_y(x, t) = A \cos k(x - ct) \\ B_z(x, t) = \frac{1}{c} E_y(x, t) \end{cases}$$

t fixed



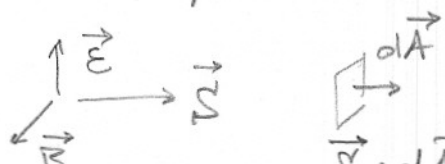
NB emphasize physical meaning and speed $c = \frac{1}{\sqrt{\mu_0 \epsilon_0}}$

- ① E.m. waves carry energy and momentum.

(explain in intuitive physical terms...) → static fields
→ propagating fields

② energy: $\vec{S} = \frac{1}{\mu_0} \vec{E} \times \vec{B}$ units: $\frac{J}{m^2 s}$

"energy flux"

 $\vec{S} \cdot d\vec{A}$: energy flowing through $d\vec{A}$ in unit time (s)

③ momentum: $\vec{P} = \frac{\vec{S}}{c}$

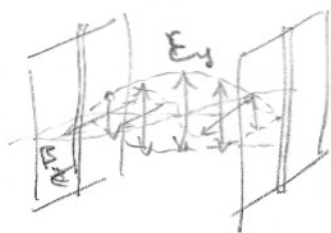
"momentum flux" : momentum through $d\vec{A}$ in unit time

$\vec{P} \cdot d\vec{A}$



- ④ "Normal modes" for an e.m. field

it is an interesting exercise to verify that there are "standing wave" solutions for $\vec{E}$ and $\vec{B}$ between perfectly conducting parallel plates, in a vacuum



$$E_y = A \sin\left(n \frac{\pi}{L} x\right) \cos\left(n \frac{\pi}{L} c t\right) \quad n = 1, 2, \dots, \infty$$

$$-c B_z = A \cos\left(n \frac{\pi}{L} x\right) \sin\left(n \frac{\pi}{L} c t\right)$$

- ⑤ Radiation sources

↑ "normal modes" → $\omega_n = k_n c = n \frac{\pi}{L} c$

See next.

What are the possible sources of classical "radiation" (e.m. waves)?

Accelerated charges!

amplitude of e.m. radiation

(intensity (energy) $\propto A^2$)

$$A \propto q \left(\frac{d^2 \vec{r}}{dt^2} \right)_{\perp}$$

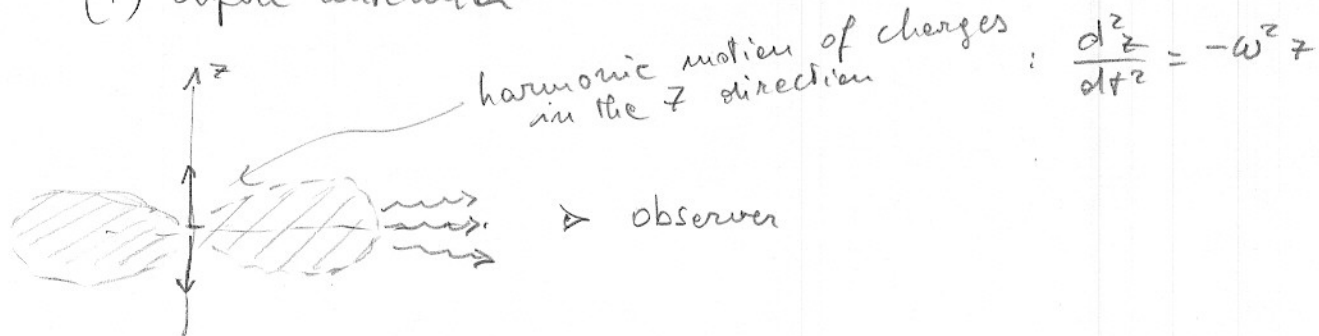
proportional

charge

component of the acceleration $\perp$ direction of observation.

examples:

(1) dipole antenna



(2) storage ring, synchrotron radiation



see also: J.R. Taylor, C.D. Zafiratos, M.A. Dubson
Modern Physics for Scientists and Engineers.

chapter 11, Atomic Transitions and Radiation

LØ3/12 bis

in more detail: (Taylor-Zafiratos-Dubson, p. 335)

power radiated by a single charge moving non-relativistically

$$P = \frac{1}{4\pi\epsilon_0} \cdot 2 \cdot \frac{q^2 a^2}{3c^2}$$

$$\frac{1}{4\pi\epsilon_0} \equiv k$$

useful when discussing Bohr's atom: ($q = 1.6 \times 10^{-19} \text{ C}$)

one can compute $P = 2.9 \times 10^{11} \text{ eV/s}$

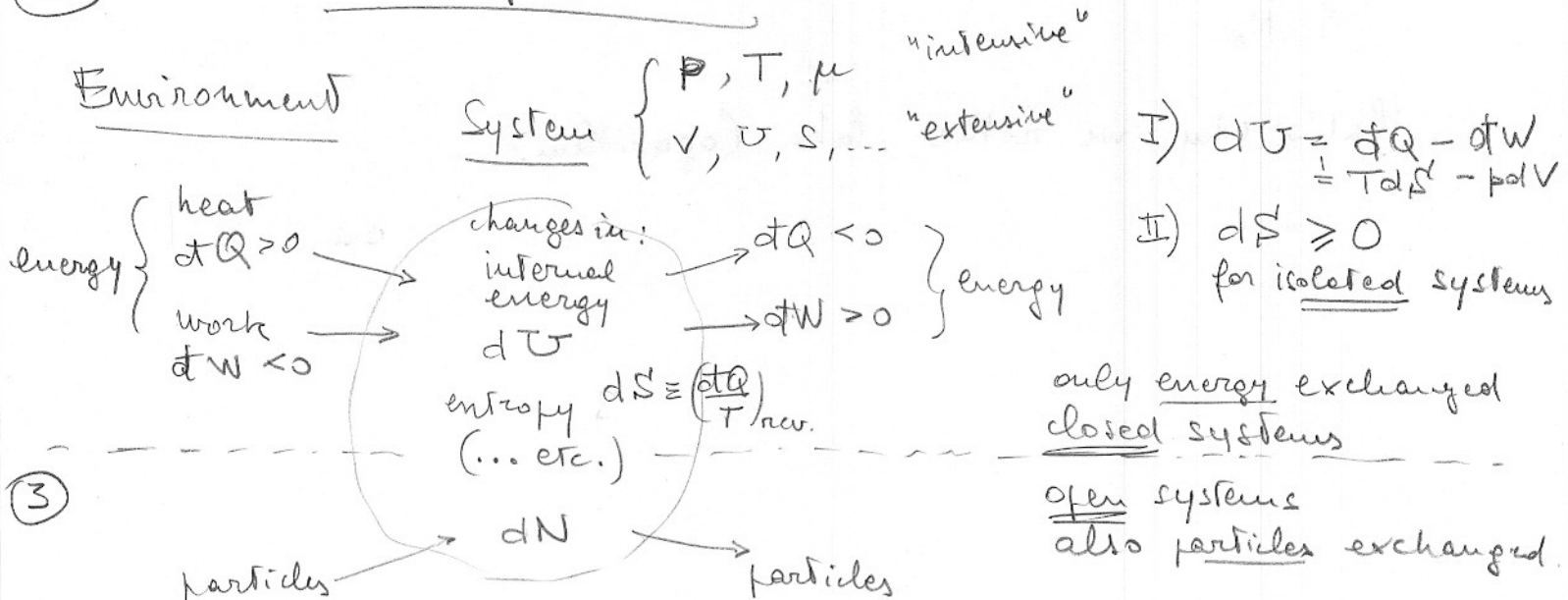
$$a = \frac{F}{m} = \frac{1}{4\pi\epsilon_0} \frac{q^2}{r^2} \cdot \frac{1}{m}$$

→ classically: quick collapse!

$$r \approx 0.5 \times 10^{-10} \text{ m}$$

① Let's recall now some results from Thermodynamics and Statistical Mechanics (also called sometimes "Thermal Physics")

② Thermodynamics ("macroscopic")



③

④ Statistical Mechanics

"macrostate": "ensemble" of "microstates" { very large number Ω , equiprobable, all explored, given enough time }

(typical example: $\circ \circ \circ \circ \circ \circ \circ \circ \circ \circ$
 $\Omega = 4$ quanta in $n = 10$ atoms: $\Omega = \binom{n}{n} = \frac{n!}{(n-n)! \cdot 1!} = 210$) * [see back]

equilibrium state of a large system: { entropy $S = k_B \ln \Omega$, etc... (average properties ... probabilities) }

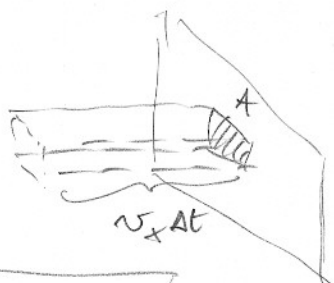
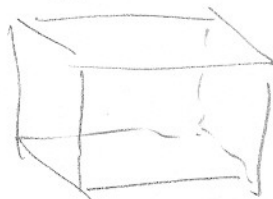
$dS = 0$: maximum entropy!

⑤ $\Rightarrow$ just recall a few results:

- meanings of T ($k_B T$), P , μ
- equipartition theorem: energy and degrees of freedom.

Meaning of T , the Boltzmann constant k_B , and the equipartition theorem.

Ideal gas



experimentally:

$$PV = \bar{n} R T \quad (1)$$

pressure

volume

n. of moles

$$\bar{n} N_A = N$$

$R \approx 8.3 \text{ J/mol} \cdot \text{K}$
universal constant for rarefied gases at high temperatures

P : force/area

$$\bar{n} = \frac{N}{V}$$

number of particles
Avogadro

$\bar{n} = \frac{N}{V}$ density of particles.

$$\Delta N = (v_x \Delta t) \cdot A \cdot \bar{n} \times \frac{1}{2}$$

number of particles hitting the wall in a time Δt

$2mv_x$

momentum transferred to the wall per particle (elastic scattering)

$$\Delta p = \Delta N \cdot 2mv_x = 2mv_x \cdot v_x \Delta t \cdot A \cdot \frac{1}{2} \bar{n}$$

average force on A

$F \Delta t$ impulse

$$\Rightarrow F = mv_x^2 \cdot A \cdot \bar{n}$$

$$\Rightarrow \langle P \rangle = \frac{F}{A} = \frac{N}{V} \cdot \frac{1}{2} mv^2$$

$$\frac{1}{3} mv^2 = \frac{2}{3} \cdot \frac{1}{2} mv^2 = \frac{2}{3} \langle \text{kinetic energy} \rangle$$

$$\Rightarrow PV = N \cdot \frac{2}{3} \langle \frac{1}{2} mv^2 \rangle = \frac{2}{3} U$$

internal energy

$$U = N \langle \frac{1}{2} mv^2 \rangle$$

$$PV = \bar{n} N_A \cdot \frac{2}{3} \langle \frac{1}{2} mv^2 \rangle \quad (2)$$

kinetic model

① Comparing the exp. tal results with the kinetic model.

$$pV = \begin{cases} \bar{n} R T & \text{(experiment)} \\ \bar{n} N_A \cdot \frac{2}{3} \langle \frac{1}{2} m v^2 \rangle & \text{(theory)} \end{cases} = \bar{n} \cdot N_A \cdot \left(\frac{R}{N_A} \right) T = \bar{n} N_A \cdot k_B T$$

(Note: $\frac{R}{N_A}$ and $\frac{2}{3} \langle \frac{1}{2} m v^2 \rangle$ are circled and labeled as k_B and "can be identified".)

② $\Rightarrow$ "internal energy" $\langle \frac{1}{2} m v^2 \rangle = 3 \times \frac{1}{2} k_B T$ $\leftarrow$ one particle

$$U = \bar{n} N_A \langle \frac{1}{2} m v^2 \rangle = 3 \bar{n} N_A \cdot \frac{1}{2} k_B T \leftarrow N \text{ particles.}$$

Several important conclusions:

③ T : is proportional to a measure of the "average kinetic energy" (at least for ideal gases)

④ k_B : Boltzmann constant :
Conversion factor from temperature (K) to energy (J)

$$k_B = \frac{R}{N_A} = \frac{8.3 \text{ J/mol} \cdot \text{K}}{6.0 \times 10^{23} \text{ mol}^{-1}} = 1.4 \times 10^{-23} \text{ J/K}$$

1.38

⑤ Special case of the "equipartition theorem" of statistical mechanics :

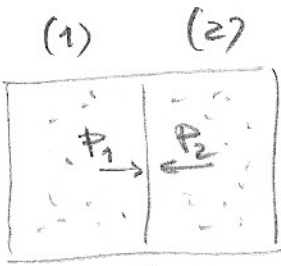
There is a contribution of $\frac{1}{2} k_B T$ to the mean energy of the system for every "quadratic degree of freedom" in the Hamiltonian.

in this case : ideal monatomic gas

$$U = 3 \bar{n} N_A \cdot \frac{1}{2} k_B T = \dots = \frac{3}{2} \bar{n} R T$$

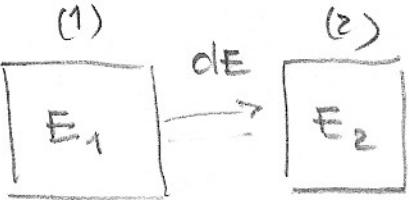
Next: meaning of $\begin{cases} P & \text{pressure} \\ T & \text{temperature} \\ \mu & \text{chemical potential} \end{cases}$

as parameters indicating "equilibrium" conditions.

(a)  Systems (1) and (2) can exchange work

$P_1 = P_2$: mechanical equilibrium

↑
mobile wall

(b)  Systems (1) and (2) can only exchange energy among themselves (heat)

$\Omega_1(E_1)$ $\Omega_2(E_2)$ microstates: number $\Omega(E)$

↑
depends on energy

most probable division of energy between (1) and (2):

max. $\Omega_1(E_1) \cdot \Omega_2(E_2)$ total number of microstates

$$\Leftrightarrow \frac{d}{dE_1} (\quad " \quad) =$$

$$\Omega_2(E_2) \frac{d\Omega_1}{dE_1} + \Omega_1(E_1) \frac{d\Omega_2}{dE_2} \frac{dE_2}{dE_1} = 0 \quad / \quad \frac{1}{\Omega_1 \Omega_2}$$

↑ = -1

$$\frac{1}{\Omega_1} \frac{d\Omega_1}{dE_1} = \frac{1}{\Omega_2} \frac{d\Omega_2}{dE_2} = 0$$

($dE_2 = -dE_1$: total energy unchanged.)

$$\frac{d \ln \Omega_1}{dE_1} = \frac{d \ln \Omega_2}{dE_2}$$

equilibrium :
"equal temperature"

↑

↑

we identify these expression with the def. of Temp.

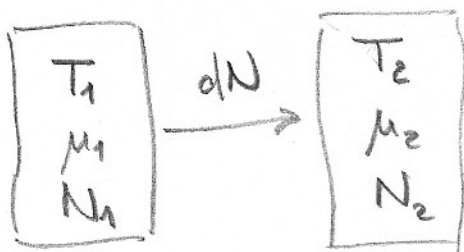
$$\frac{1}{k_B T} = \frac{d \ln \Omega}{dE}$$

$$\frac{1}{T} = \frac{d(k_B \ln \Omega)}{dE} = \frac{dS}{dE}$$

$$\frac{1}{T_1} = \frac{1}{T_2} \quad \text{or} \quad T_1 = T_2 : \quad \underline{\text{thermal equilibrium}}$$

$$(k_B = 1.38 \times 10^{-23} \text{ J K}^{-1}) \quad \text{L45/4}$$

(c)



Systems (1) and (2)
can exchange particles.

what happens? see next.

Blundell - Blundell, p. 244

1. The chemical potential

Generalization of thermodynamic potentials when the number of particles may change:

$$dU = TdS - PdV + \mu dN$$

← NB: for instance
 $N \sim 10^{23}$
 $dN \sim 10^6$ or 10^9
 is a "small" change

$$\mu = \left(\frac{\partial U}{\partial N} \right)_{S, V}$$

S, V not easy to fix in practice

⇒ use rather:

$$F = U - TS \Rightarrow dF = \cancel{TdS} - PdV - \cancel{TdS} - SdT + \mu dN$$

$$= -PdV - SdT + \mu dN$$

$$\Rightarrow \mu = \left(\frac{\partial F}{\partial N} \right)_{V, T}$$

$$G = H - TS =$$

$$= U + PV - TS \Rightarrow dG = \cancel{TdS} - \cancel{PdV} + \cancel{PdV} + VdP - \cancel{TdS} - SdT + \mu dN$$

$$= VdP - SdT + \mu dN$$

$$\Rightarrow \mu = \left(\frac{\partial G}{\partial N} \right)_{P, T}$$

⇒ chemical potential

"Gibbs free energy per particle"

↳ constant P, T

2. Meaning of the chemical potential ?

different values of the chemical potential drives particles to move from one place to another, much in the same way as a difference in temperature drives the transfer of energy (heat).

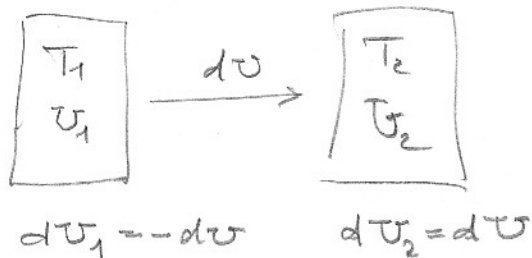
→ equilibrium ? max. entropy

$$dU = TdS - PdV + \mu dN$$

$$dS = \frac{1}{T} (dU + PdV - \mu dN)$$

$$\Rightarrow \left(\frac{\partial S}{\partial U} \right)_{V,N} = \frac{1}{T} \quad \left(\frac{\partial S}{\partial V} \right)_{U,N} = \frac{P}{T} \quad \left(\frac{\partial S}{\partial N} \right)_{U,V} = -\frac{\mu}{T}$$

(a) heat flow: (constant N) from system 1 to system 2



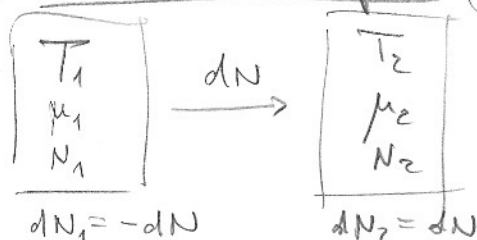
$$dS' = \underbrace{\left(\frac{\partial S_1}{\partial U_1} \right)_{N,V}}_{\frac{1}{T_1}} (-dU) + \underbrace{\left(\frac{\partial S_2}{\partial U_2} \right)_{N,V}}_{\frac{1}{T_2}} dU$$

$$dS' = \left(-\frac{1}{T_1} + \frac{1}{T_2} \right) dU \geq 0$$

$dS' = 0$ if $T_1 = T_2$ (thermal equilibrium : no entropy increase)
 : reversible heat exchange at $T_1 \approx T_2 \approx T$

$$\left. \begin{array}{l} dS' > 0 \\ dU > 0 \end{array} \right\} \Leftrightarrow -\frac{1}{T_1} + \frac{1}{T_2} > 0 \Leftrightarrow \frac{1}{T_2} > \frac{1}{T_1} \Leftrightarrow T_1 > T_2$$

(b) particle exchange (dN) from system 1 to system 2



$$dS' = \left(\frac{\partial S_1}{\partial N_1} \right)_{U,V} (-dN) + \left(\frac{\partial S_2}{\partial N_2} \right)_{U,V} dN = \left(\frac{\mu_1}{T_1} - \frac{\mu_2}{T_2} \right) dN \geq 0$$

For $T_1 = T_2 = T$

$$dS' = 0 \Leftrightarrow \mu_1 = \mu_2$$

$dN > 0$

$$dS' \geq 0 \Leftrightarrow \frac{\mu_1 - \mu_2}{T} dN \geq 0 \Leftrightarrow \mu_1 > \mu_2$$

3. "Total" chemical potential

if the particles have interactions with conservative forces, with a potential energy (for instance:

$$\begin{cases} U(z) = mgz & : \text{gravitational energy, or:} \\ U = qV & : \text{electrostatic energy} \end{cases}$$

etc.
total internal energy:

$$dU = TdS - PdV + \underbrace{\mu dN + qV}_{\text{Total electro-chemical potential}}$$

$\swarrow$ chemical potential μ $\swarrow$ electrical potential

Gibbs function:

$$dG = \underbrace{SdT}_{0} + \underbrace{VdP}_{0} + \underbrace{\mu dN + qdV}_{\downarrow} = 0$$

$T = \text{const}$
 $P = \text{const.}$

$$\mu dN = -q dV$$

NB: equilibrium requires the "total chemical potential", in this case the "electro-chemical potential", to be a constant (we will see this property is very useful in semiconductors) (discussing)

from here: Powerpoint slides.