### <span id="page-0-0"></span>The Sommerfeld Theory of Metals

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### <span id="page-1-0"></span>Outline





3 [Thermal properties of the free electron gas](#page-26-0)



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<span id="page-3-0"></span>General remarks

#### Maxwell-Boltzmann velocity distribution

Drude assumed the validity of the Maxwell-Boltzmann velocity distribution (at thermal equilibrium)

$$
f_B(\mathbf{v}) = n\left(\frac{m}{2\pi k_B T}\right)^{\frac{3}{2}} e^{-\frac{mv^2}{2k_B T}}
$$

- normalized such that  $\int f_{\mathcal{B}}(\textbf{\textit{v}})d\textbf{\textit{v}} = n, \ n = \frac{N}{V}$ V
	- check yourself  $\left(\int_{-\infty}^{\infty} e^{-\alpha x^2} dx = \sqrt{\frac{\pi}{\alpha}}\right)$
- $f(\bm{v})d\bm{v}$ : number of  $e^-$  with velocities in the volume element  $d\bm{v}$  at  $\bm{v}$ • per unit volume
- Equipartition theorem follows ( $c_v = \frac{3}{2}$  $\frac{3}{2}k_B$ ;  $\bar{\varepsilon}=\frac{3}{2}$  $\frac{3}{2}k_B T$

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<span id="page-4-0"></span>General remarks

#### Fermi-Dirac velocity distribution

• Valid for Fermions, as a consequence of the Pauli exclusion principle

$$
f(\mathbf{v}) = \frac{(m/\hbar)^3}{4\pi^3} \frac{1}{e^{\frac{(\frac{1}{2}mv^2 - k_B T_0)}{k_B T}} + 1}
$$

 $T_0$  determined such that  $\int f(\mathbf{v})d\mathbf{v} = n$ 

- Sommerfeld theory replaces  $f_B(\mathbf{v})$  of Drude's theory with  $f(\mathbf{v})$ 
	- profound consequences on  $\bar{\varepsilon}$  and  $c_v$

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<span id="page-5-0"></span>General remarks

Maxwell-Boltzmann vs Fermi-Dirac velocity distribution



Plot of Maxwell-Boltzmann and Fermi-Dirac distributions for the same n given by  $T = 0.01T_0$ .

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[Daniele Toffoli](#page-0-0) January 11, 2017 6 / 48

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<span id="page-6-0"></span>Quantum mechanical solution

#### Mathematical treatment

- The free electron gas is confined in a cube of edge L  $(L=V^{\frac{1}{3}})$
- Assume the independent electron approximation
- Solve the TISE (separation of variables)

$$
-\frac{\hbar^2}{2m}\nabla^2\psi(\mathbf{r})=\varepsilon\psi(\mathbf{r})
$$

Apply Born-von Karman boundary conditions to the general solution

$$
\begin{cases}\n\psi(x, y, z + L) = \psi(x, y, z) \\
\psi(x, y + L, z) = \psi(x, y, z) \\
\psi(x + L, y, z) = \psi(x, y, z)\n\end{cases}
$$

Fill the energy levels by using the Pauli exc[lus](#page-5-0)i[o](#page-7-0)[n](#page-5-0) [pr](#page-6-0)[in](#page-7-0)[c](#page-1-0)[i](#page-2-0)[p](#page-19-0)[le](#page-20-0)<br>Angle Toffoli

<span id="page-7-0"></span>Quantum mechanical solution

#### Separation of variables

- Ansatz  $\psi(\mathbf{r}) = X(x)Y(y)Z(z)$
- Upon substitution on the TISE:

$$
\begin{cases}\n-\frac{\hbar^2}{2m}\frac{d^2}{dx^2}X(x) &= \varepsilon_x X(x) \\
-\frac{\hbar^2}{2m}\frac{d^2}{dy^2}Y(y) &= \varepsilon_y Y(y) \\
-\frac{\hbar^2}{2m}\frac{d^2}{dz^2}Z(z) &= \varepsilon_z Z(z)\n\end{cases}
$$

$$
\bullet \varepsilon = \varepsilon_x + \varepsilon_y + \varepsilon_z
$$

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<span id="page-8-0"></span>Quantum mechanical solution

Separation of variables

- Ansatz  $\psi(\mathbf{r}) = X(x)Y(y)Z(z)$
- The boundary conditions are:

$$
\begin{cases}\nZ(z+L) = Z(z) \\
Y(y+L) = Y(y) \\
X(x+L) = X(x)\n\end{cases}
$$

Three similar homogeneous ODEs of the second order

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<span id="page-9-0"></span>Quantum mechanical solution

#### General and particular solutions

$$
\begin{cases} \frac{d^2}{dx^2}\phi(x) + k^2\phi(x) & = 0\\ \phi(x+L) & = \phi(x) \end{cases}
$$

#### General solution:  $\phi(x) = c_1 e^{ikx} + c_2 e^{-ikx}$  $k^2 = \frac{2m\varepsilon}{\hbar^2} \Longrightarrow \varepsilon = \frac{\hbar^2 k^2}{2m}$ 2m

Particular (normalized) solution:  $\phi(x) = \frac{1}{\sqrt{2}}$  $\frac{1}{L}e^{ikx}$ 

• 
$$
k = \frac{2n\pi}{L}
$$
,  $n = 0, \pm 1, \pm 2, \ldots$ 

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<span id="page-10-0"></span>Quantum mechanical solution

Eigenfunctions and eigenvalues of the problem

 $\psi_{\bm k}(\bm r) = \frac{1}{\sqrt{2}}$  $\frac{1}{V}e^{i\boldsymbol{k}\cdot \boldsymbol{r}}$ 

normalized inside the cube:  $\int |\psi_{\bm k}(\bm r)|^2 d\bm r = 1$ 

**allowed** wave vectors  $k$ :

• 
$$
k_x = \frac{2\pi n_x}{L}
$$
;  $k_y = \frac{2\pi n_y}{L}$ ;  $k_z = \frac{2\pi n_z}{L}$ 

• 
$$
n_x, n_y, n_z = 0, \pm 1, \pm 2, ...
$$

• 
$$
\varepsilon = \varepsilon(|\mathbf{k}|) = \frac{\hbar^2 k^2}{2m}
$$
 (depends only on  $k = |\mathbf{k}|$ )



allowed wave vectors in a 2D k space

<span id="page-11-0"></span>Quantum mechanical solution

#### Eigenfunctions and eigenvalues of the problem

 $\psi_{\bm k}(\bm r)$  is eigenfunction of the momentum operator  $\bm p = \frac{\hbar}{\tilde{l}} \nabla$ 

• 
$$
\frac{\hbar}{i}\nabla\psi_{k}(r)=p\psi_{k}(r); p=\hbar k
$$

• Its velocity is 
$$
\mathbf{v} = \frac{\mathbf{p}}{m} = \frac{\hbar \mathbf{k}}{m}
$$



allowed wave vectors in a  $2D$   $k$  space

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 $A \cup B \rightarrow A \oplus B \rightarrow A \oplus B \rightarrow A \oplus B \rightarrow B$ 

<span id="page-12-0"></span>Quantum mechanical solution

Counting the quantum mechanical solutions

- Volume per allowed wave vector in *k*-space:  $(\frac{2\pi}{L})^3 = \frac{8\pi^3}{V}$ V
- For a region  $Ω$ , the number is  $\frac{ΩV}{8π³}$ 
	- must be very large on the scale of  $\frac{2\pi}{L}$
	- not too irregularly shaped
- *k*-space density of levels:  $\frac{V}{8\pi^3}$



allowed wave vectors in a  $2D$   $k$  space

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<span id="page-13-0"></span>Quantum mechanical solution

#### Occupation of the ground-state ( $T = 0K$ )

- Place a maximum of two  $e^-$  on each level, starting with the lowest
	- $k = 0 \Longrightarrow \varepsilon_k = 0$
	- $\psi \equiv \psi_{\boldsymbol{k}\sigma}$ ,  $\sigma = \pm \frac{1}{2}$
- $\bullet$   $\varepsilon_k$  varies with the distance squared from O
- The occupied region is a sphere (Fermi sphere)
	- $\bullet$  for very large N
	- radius  $k_F$ : Fermi wave vector
	- volume:  $\Omega = \frac{4\pi}{3}k_F^3$

• 
$$
N = 2\frac{\Omega V}{8\pi^3} = 2(\frac{4\pi k_F^3}{3})(\frac{V}{8\pi^3}) = \frac{k_F^3}{3\pi^2}V
$$

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<span id="page-14-0"></span>Quantum mechanical solution

Occupation of the ground-state ( $T = 0K$ )

- Given a density  $n=\frac{N}{V}$  $\frac{N}{V}$ , the ground-state is formed by:
	- $\bullet$  occupying all levels with  $k < k_F$
	- all levels with  $k > k_F$  are empty

$$
\bullet \ \ k_F = (3\pi^2 n)^{\frac{1}{3}}
$$

- Some nomenclature:
	- region  $Ω$ : Fermi sphere
	- $\bullet$   $k_F$ : Fermi wave vector
	- surface of Ω: Fermi surface
	- $p_F = \hbar k_F$ : Fermi momentum

• 
$$
v_F = \frac{p_F}{m}
$$
: Fermi velocity

• 
$$
\varepsilon_F = \frac{\hbar^2 k_F^2}{2m}
$$
: Fermi energy

- $T_F = \frac{\varepsilon_F}{k_B}$ : Fermi temperature
- $\bullet$  The above quantities can be estimated from  $n$

<span id="page-15-0"></span>Quantum mechanical solution

#### Some numbers

\n- \n
$$
k_F = \frac{(\frac{9\pi}{4})^{\frac{1}{3}}}{r_s} = \frac{1.92}{r_s} = \frac{3.63}{\frac{r_s}{s_0}} \text{\AA}^{-1}
$$
\n
\n- \n $r_s \sim 2 - 6 \text{ Å} \implies k_F \sim \text{Å}^{-1} (\lambda \sim \text{\AA})$ \n
\n- \n $v_F = \frac{\hbar}{m} k_F = \frac{4.20}{\frac{r_s}{s_0}} \times 10^8 \text{cm/s}$ \n
\n- \n $1\%$  of *c*, (classical estimate at room temperature  $v \sim 10^7 \text{cm/s}$ )\n
\n- \n $\varepsilon_F = \frac{\hbar^2 k_F^2}{2m} = \left(\frac{e^2}{2a_0}\right) (k_F a_0)^2 = \frac{50.1}{(\frac{r_s}{a_0})^2} \text{ eV}$ \n
\n- \n $\varepsilon_F \in 1.5 - 15 \text{ eV}$ \n
\n- \n $T_F = \frac{\varepsilon_F}{k_B} = \frac{58.2}{(\frac{r_s}{a_0})^2} \times 10^4 \text{K}$ \n
\n- \n energy per electron of a classical ideal gas vanishes at  $T = 0 \text{K}$ \n
\n- \n $(\bar{\varepsilon} = \frac{3}{2} k_B T)$ \n
\n

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<span id="page-16-0"></span>Quantum mechanical solution

#### Some numbers



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<span id="page-17-0"></span>Quantum mechanical solution

Total energy of the ground-state ( $T = 0K$ )

$$
\bullet \ \ E = 2 \times \sum_{k < k_F} \varepsilon_k = \sum_{k < k_F} 2 \times \frac{\hbar^2 k^2}{2m}
$$

• Standard way of treating summations:

$$
\sum_{\mathbf{k}} F(\mathbf{k}) = \frac{V}{8\pi^3} \sum_{\mathbf{k}} F(\mathbf{k}) \Delta \mathbf{k}
$$
  

$$
\lim_{V \to \infty} \frac{1}{V} \sum_{\mathbf{k}} F(\mathbf{k}) = \int \frac{F(\mathbf{k})}{8\pi^3} d\mathbf{k}
$$

**o** Therefore:

$$
\frac{E}{V} = 2 \int_{k < k_F} \frac{d\mathbf{k}}{8\pi^3} \frac{\hbar^2 k^2}{2m} = \frac{1}{\pi^2} \frac{\hbar^2 k_F^5}{10m}
$$

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<span id="page-18-0"></span>Quantum mechanical solution

Total energy of the ground-state ( $T = 0K$ )

\n- \n
$$
\frac{E}{V} = \frac{1}{\pi^2} \frac{\hbar^2 k_F^5}{10m}
$$
\n
\n- \n energy density of the electron gas\n
\n- \n
$$
\frac{E}{N} = \frac{3}{5} k_B T_F
$$
\n
\n- \n for a classical particle\n
\n- \n
$$
\frac{E}{N} = \frac{3}{2} k_B T \implies T = \frac{2}{5} T_F \sim 10^4 \text{K}
$$
\n
\n

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<span id="page-19-0"></span>Quantum mechanical solution

Bulk properties of the ground-state

- Electronic pressure:  $P = \frac{2}{3}$ 3 E V
	- exerted by the electron gas
- Compressibility:  $K = -\frac{1}{V}$  $\frac{1}{V}(\frac{\partial V}{\partial P}$  $\frac{\partial V}{\partial P}$
- Bulk modulus:  $B = \frac{1}{K} = \frac{5}{3}$  $\frac{5}{3}P = \frac{10}{9}$ 9  $\frac{E}{V}=\frac{2}{3}$  $rac{2}{3}$ n $\varepsilon$ <sub>F</sub>
- Numerically:  $B = (\frac{6.13}{r_{\rm s}/a_0})^5 \times 10^{10}$ dynes/cm $^2$

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#### 2 [The Fermi-Dirac distribution](#page-20-0)

[Thermal properties of the free electron gas](#page-26-0)

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<span id="page-21-0"></span>The Fermi-Dirac distribution

#### The partition function

- $\bullet$  If T $\neq$  0, N-electron excited states become populated
	- thermal equilibrium is assumed
- **o** Boltzmann distribution law

$$
P_N(E) = \frac{e^{-\frac{E}{k_B T}}}{\sum_{\alpha} e^{-\frac{E_N^N}{k_B T}}}
$$

 $\bullet$   $P_N(E)$ : probability of finding the system in the state of energy E  $Z=\sum_\alpha {\rm e}^{-\frac{E_\alpha^N}{k_B T}}$  is the partition function

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[The Fermi-Dirac distribution](#page-22-0)

### <span id="page-22-0"></span>Thermal properties of the free-electron gas

The Fermi-Dirac distribution

The partition function

• 
$$
F_N = U - TS = -k_B T \ln(Z) \Longrightarrow Z = e^{-\frac{F_N}{k_B T}}
$$

• Helmholtz free energy

$$
\bullet \ \ P_N(E) = e^{-\frac{E - F_N}{k_B T}}
$$

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 $\left\{ \begin{array}{ccc} 1 & 0 & 0 \\ 0 & 1 & 0 \end{array} \right.$ 

<span id="page-23-0"></span>The Fermi-Dirac distribution

**Derivation** 

- The N-electron state is specified by a list of the one-electron levels occupied
	- states  $\psi_{\mathbf{k}\sigma}(\mathbf{r})$
- Define  $f_i^N = \sum P_N(E_\alpha^N)$ 
	- **•** probability that the one-electron level  $i$  is occupied in the N-electron state
	- Pauli exclusion principle requires  $0 \le f_i^{\textit{N}} \le 1$
	- $\bullet$  mean occupation of the level  $i$
- We will find an explicit expression for  $f^N_i$

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<span id="page-24-0"></span>The Fermi-Dirac distribution

#### Derivation

\n- \n
$$
f_i^N = 1 - \sum P_N(E_j^N)
$$
\n
\n- \n $\gamma$  labels *N*-electron states where *i* is not occupied\n
\n- \n $f_i^N = 1 - \sum P_N(E_\alpha^{N+1} - \varepsilon_i)$ \n
\n- \n $E_\gamma^N = E_\alpha^{N+1} - \varepsilon_i$ \n
\n- \n $N + 1$ -electron states obtained from  $\gamma$ 's by placing an electron in level *i*\n
\n- \n Defining  $\mu = F_{N+1} - F_N \Longrightarrow f_i^N = 1 - e^{\frac{\varepsilon_i - \mu}{k_B T}} \sum P_{N+1}(E_\alpha^{N+1})$ \n
\n- \n $f_i^N = 1 - e^{\frac{\varepsilon_i - \mu}{k_B T}} f_i^{N+1}$ \n
\n

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<span id="page-25-0"></span>The Fermi-Dirac distribution

#### Derivation

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• Assuming that 
$$
f_i^N = f_i^{N+1}
$$
 for  $N \sim 10^{22}$ 

$$
f^{\mathcal{N}}_i = \frac{1}{e^{\frac{(\varepsilon_i - \mu)}{k_{\mathcal{B}} T}}+1}
$$

\n- \n
$$
N = \sum_{i} f_i^N = \sum_{i} \frac{1}{\frac{(\varepsilon_i - \mu)}{\varepsilon^k \varepsilon^T + 1}}
$$
\n
\n- \n
$$
N \text{ (or } n = \frac{N}{V} \text{) as a function of } T \text{ and } \mu
$$
\n
\n- \n
$$
N \text{ (or } n = \frac{N}{V} \text{) as a function of } T \text{ and } \mu
$$
\n
\n

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[The Sommerfeld theory of conduction in metals](#page-41-0)

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<span id="page-27-0"></span>Applications of the Fermi-Dirac distribution

#### Limiting form of  $f_i^N$

• In the ground-state:

$$
\begin{cases} f_{\mathbf{k}\sigma} = 1 & \varepsilon(\mathbf{k}) < \varepsilon_F \\ f_{\mathbf{k}\sigma} = 0 & \varepsilon(\mathbf{k}) > \varepsilon_F \end{cases}
$$

For the  $f_i^N$  distribution we have

$$
\begin{cases} f_{\mathbf{k}\sigma} = 1 & \varepsilon(\mathbf{k}) < \mu \\ f_{\mathbf{k}\sigma} = 0 & \varepsilon(\mathbf{k}) > \mu \end{cases}
$$

- Therefore  $\lim_{T\to 0} \mu = \varepsilon_F$
- $\mu \sim \varepsilon_F$  also at room temperature

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<span id="page-28-0"></span>Applications of the Fermi-Dirac distribution

Total energy of the electron gas

At any T,  $U = 2 \sum_{\mathbf{k}} \varepsilon(\mathbf{k}) f(\varepsilon(\mathbf{k}))$ :

$$
f(\varepsilon(\bm{k}))=\frac{1}{e^{\frac{(\varepsilon-\mu)}{k_{B}T}}+1}
$$

Defining  $u = \frac{U}{V}$ V

$$
u=\int \frac{d\mathbf{k}}{4\pi^3}\varepsilon(\mathbf{k})f(\varepsilon(\mathbf{k}))
$$

From  $N = 2 \sum_{\mathbf{k}} f(\varepsilon(\mathbf{k}))$ :

$$
n=\int \frac{d\mathbf{k}}{4\pi^3}f(\varepsilon(\mathbf{k}))
$$

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<span id="page-29-0"></span>Applications of the Fermi-Dirac distribution

#### Density of levels

• Working in spherical coordinates:

$$
\int \frac{d\mathbf{k}}{4\pi^3} F(\varepsilon(\mathbf{k})) = \int_0^\infty \frac{k^2 dk}{\pi^2} F(\varepsilon(\mathbf{k})) = \int_{-\infty}^\infty d\varepsilon g(\varepsilon) F(\varepsilon)
$$

 $g(\epsilon)$ : density of levels (per unit volume)

$$
g(\varepsilon) = \begin{cases} \frac{m}{\hbar^2 \pi^2} \sqrt{\frac{2m\varepsilon}{\hbar^2}} & \varepsilon > 0\\ 0 & \varepsilon < 0 \end{cases}
$$

At the Fermi level:  $g(\varepsilon_F) = \frac{mk_F}{\hbar^2\pi^2}$ 

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#### <span id="page-30-0"></span>Thermal properties of the free-electron gas

Applications of the Fermi-Dirac distribution

#### Density of levels

**•** Alternatively:  $g(\varepsilon) = \begin{cases} \frac{3}{2} & \text{if } \varepsilon \leq \frac{3}{2} \\ 0 & \text{if } \varepsilon \leq \frac{3}{2} \end{cases}$ 2 n  $\frac{n}{\varepsilon_F}$  $\left(\frac{\varepsilon}{\varepsilon_H}\right)$  $(\frac{\varepsilon}{\varepsilon_{\mathsf{F}}})^{1/2}$   $\varepsilon > 0$ 0  $\varepsilon < 0$ At the Fermi level:  $g(\varepsilon_F) = \frac{3}{2}$ n εF

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<span id="page-31-0"></span>Applications of the Fermi-Dirac distribution

Total energy and density of the electron gas

$$
\bullet \ \ u = \int_{-\infty}^{\infty} d\varepsilon g(\varepsilon) \varepsilon f(\varepsilon)
$$

$$
\bullet \; n = \int_{-\infty}^{\infty} d\varepsilon g(\varepsilon) f(\varepsilon)
$$

• valid for any non-interacting electron systems

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<span id="page-32-0"></span>Applications of the Fermi-Dirac distribution

#### The Sommerfeld expansion

- For metals,  $T \ll T_F$  even at room temperature
- For  $T \neq 0$   $f(\varepsilon)$  differs little from its T=0 form
	- region  $\Delta \varepsilon \sim k_B T$  around  $\mu$



Fer[m](#page-26-0)i function f for given  $\mu$  at  $T = 0K$  (top) and room t[emp](#page-31-0)e[rat](#page-33-0)[ur](#page-31-0)[e \(b](#page-32-0)[ot](#page-33-0)[to](#page-25-0)m[\)](#page-40-0)

<span id="page-33-0"></span>Applications of the Fermi-Dirac distribution

#### The Sommerfeld expansion

- Applied to integrals of the type  $\int_{-\infty}^{\infty} H(\varepsilon)f(\varepsilon)d\varepsilon.$
- If  $H(\varepsilon)$  does not vary much for  $\Delta \varepsilon \sim k_BT$  around  $\mu$ 
	- Taylor expansion of  $H(\varepsilon)$  around  $\mu$
	- assumed to converge rapidly for well-behaved  $H(\varepsilon)$

$$
\int_{-\infty}^{\infty} H(\varepsilon) f(\varepsilon) d\varepsilon = \int_{-\infty}^{\mu} H(\varepsilon) d\varepsilon \n+ \frac{\pi^2}{6} (k_B T)^2 H'(\mu) + \frac{7\pi^4}{360} (k_B T)^4 H'''(\mu) + O(\frac{k_B T}{\mu})^6
$$

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<span id="page-34-0"></span>Applications of the Fermi-Dirac distribution

#### Specific heat of the electron gas

• Apply the Sommerfeld expansion to both  $u$  and  $n$ .

$$
u = \int_0^{\mu} \varepsilon g(\varepsilon) d\varepsilon + \frac{\pi^2}{6} (k_B T)^2 [\mu g'(\mu) + g(\mu)] + O(T^4)
$$
  

$$
n = \int_0^{\mu} g(\varepsilon) d\varepsilon + \frac{\pi^2}{6} (k_B T)^2 g'(\mu) + O(T^4)
$$

Also, to order  $\mathcal{T}^2$ :

$$
\int_0^\mu H(\varepsilon)d\varepsilon=\int_0^{\varepsilon_F}H(\varepsilon)d\varepsilon+(\mu-\varepsilon_F)H(\varepsilon_F)
$$

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<span id="page-35-0"></span>Applications of the Fermi-Dirac distribution

#### Specific heat of the electron gas

**•** Therefore:

$$
u = \int_0^{\varepsilon_F} \varepsilon g(\varepsilon) d\varepsilon + \varepsilon_F \left\{ (\mu - \varepsilon_F) g(\varepsilon_F) + \frac{\pi^2}{6} (k_B T)^2 g'(\varepsilon_F) \right\} + \frac{\pi^2}{6} (k_B T)^2 g(\varepsilon_F) n = \int_0^{\varepsilon_F} g(\varepsilon) d\varepsilon + \left\{ (\mu - \varepsilon_F) g(\varepsilon_F) + \frac{\pi^2}{6} (k_B T)^2 g'(\varepsilon_F) \right\}
$$

Note:

$$
\begin{array}{ll}\n\bullet & n = \int_0^{\varepsilon_F} g(\varepsilon) d\varepsilon \\
\bullet & u_0 = \int_0^{\varepsilon_F} \varepsilon g(\varepsilon) d\varepsilon\n\end{array}
$$

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 $\left\{ \begin{array}{ccc} 1 & 0 & 0 \\ 0 & 1 & 0 \end{array} \right.$ 

<span id="page-36-0"></span>Applications of the Fermi-Dirac distribution

#### Deviation of  $\mu$  from  $\varepsilon_F$

From:

$$
0=(\mu-\varepsilon_F)g(\varepsilon_F)+\frac{\pi^2}{6}(k_BT)^2g'(\varepsilon_F)
$$

We get:

$$
\mu = \varepsilon_F - \frac{\pi^2}{6} (k_B T)^2 \frac{g'(\varepsilon_F)}{g(\varepsilon_F)}
$$

$$
= \varepsilon_F \left[ 1 - \frac{1}{3} (\frac{\pi k_B T}{2\varepsilon_F})^2 \right]
$$

• The difference is  $\sim$  0.01% even at room T

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<span id="page-37-0"></span>Applications of the Fermi-Dirac distribution

#### Specific heat capacity,  $c_v$

From:

$$
u=u_0+\frac{\pi^2}{6}(k_BT)^2g(\varepsilon_F)
$$

• We get:

$$
c_{v} = \left(\frac{\partial u}{\partial T}\right)_{v} = \frac{\pi^{2}}{3}k_{B}^{2}Tg(\varepsilon_{F})
$$

$$
= \frac{\pi^{2}}{2}\left(\frac{k_{B}T}{\varepsilon_{F}}\right)nk_{B}
$$

- varies linearly with T
- compare with the classical result  $c_v = \frac{3}{2}nk_B \left(\frac{k_B T}{\varepsilon_F}\right) \sim 10^{-2}$ )
- $\bullet$  electronic contribution is negligible even at room  $\mathsf T$

<span id="page-38-0"></span>Applications of the Fermi-Dirac distribution

Specific heat capacity,  $c_v$ : qualitative considerations

- **•** From the T-dependence of the Fermi function  $f(\varepsilon)$ :
	- nr. of electrons excited (per unit volume):  $\sim g(\varepsilon_F) \times k_B T$
	- excitation energy:  $\sim k_B T$
	- energy density:  $\sim (k_B T)^2 g(\varepsilon_F) \rightarrow c_{\rm v} \sim k_B^2 \, T g(\varepsilon_F)$



Fermi function at  $T \neq 0$ 

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<span id="page-39-0"></span>Applications of the Fermi-Dirac distribution

Experimental verification of  $c_v = \frac{\pi^2}{2}$  $rac{\tau^2}{2} \left( \frac{k_B T}{\varepsilon_F} \right)$  nk<sub>B</sub>

At room T  $c_{\rm v}$  is determined by the ionic contribution  $(\propto T^3$  for  $T \rightarrow 0$ 

 $c_v = \gamma T + AT^3$ 

Experimental data (of c<sub>p</sub>) are fitted to the equation:  $\frac{c_V}{7} = \gamma + AT^2$ 

- $\bullet$  electronic contribution is comparable to the ionic at T of few K
- extrapolate at  $T \rightarrow 0$
- Experimentally  $[C] = [\frac{cal}{Kmol}]$ . Multiply by  $\frac{ZN_A}{n}$ :

$$
C = \frac{\pi^2}{3} ZR \frac{k_B T g(\varepsilon_F)}{n}
$$
  
\n
$$
\implies \gamma = \frac{1}{2} \pi^2 R \frac{Z}{T_F} = 0.169 Z (\frac{r_s}{a_0})^2 \times 10^{-4} \text{cal mol}^{-1} K^{-2}
$$

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<span id="page-40-0"></span>Applications of the Fermi-Dirac distribution

#### Experimental verification of  $c_v = \frac{\pi^2}{2}$  $rac{\tau^2}{2} \left( \frac{k_B T}{\varepsilon_F} \right)$  nk<sub>B</sub>

SOME ROUGH EXPERIMENTAL VALUES FOR THE COEFFICIENT OF THE LINEAR TERM IN T OF THE MOLAR SPECIFIC HEATS OF METALS, AND THE VALUES GIVEN BY SIMPLE FREE ELECTRON THEORY



<span id="page-41-0"></span>

[The Fermi-Dirac distribution](#page-20-0)

[Thermal properties of the free electron gas](#page-26-0)



[The Sommerfeld theory of conduction in metals](#page-41-0)

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<span id="page-42-0"></span>Fermi-Dirac velocity distribution

#### Velocity distribution for electrons in a metal

- **Consider an element of volume dk around k** 
	- number of one-electron levels: 2  $\times$   $(\frac{V}{(2\pi)^3})$   $=$   $\frac{V}{(4\pi^3)}$   $d$   $\bm{k}$
	- probability of occupation:  $f(\varepsilon(\mathbf{k}))$
	- total number of electrons:  $f(\varepsilon(\bm k)) \frac{V}{(4\pi^3)}d\bm k$
	- with velocity  $\bm{v} = \frac{\hbar \bm{k}}{m} \rightarrow d \bm{k} = (\frac{m}{\hbar})^3 d \bm{v}$
- therefore the number of electrons with velocity  $\in (v, v + dv)$  is:

$$
f(\mathbf{v})d\mathbf{v} = \frac{\left(\frac{m}{\hbar}\right)^3}{4\pi^3} \frac{1}{e^{\frac{1/2mv^2-\mu}{k_BT}} + 1} d\mathbf{v}
$$

• probability density (per unit volume)

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<span id="page-43-0"></span>Validity of the classical description

#### $\Delta x \Delta p \sim \hbar$

- Sommerfeld used the Fermi-Dirac velocity distribution in an otherwise classical theory
- Classical description of electron dynamics is valid if:
	- $r$  and  $p$  can be specified as accurately as necessary
	- without violating the uncertainty principle  $(\Delta \times \Delta p \sim \hbar)$
- The classical description is valid if:
	- $\bullet$   $\Delta p << \hbar k_F \sim p$
	- $\Delta x \sim {\hbar\over \Delta p} >> {1\over k_F} \sim r_s \sim 2$ –6Å

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<span id="page-44-0"></span>Validity of the classical description

#### $\Delta x \Delta p \sim \hbar$

- Electronic position must be specified in some instances:
	- for applied electromagnetic fields ( $\Delta x << \lambda$ )
	- $\bullet$  for applied T gradients
- Conclusions of the models were valid if  $\bm{E}$  or  $\bm{\tau}$  vary negligibly in the scale of ∆x
	- valid for UV-vis radiation, not X-rays (QM must be used)
	- usually valid for normal  $\nabla T$ 's
- We assumed  $\Delta x \ll l$ , the mean free path
	- $\bullet$  /  $\sim$  100Å at room T

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<span id="page-45-0"></span>Improvements over Drude's theory

#### The use of Fermi-Dirac velocity distribution

- Affected properties:
	- mean free path
	- thermal conductivity
	- **•** thermopower
- Properties not affected:
	- **•** magnetoresistance
	- Hall coefficient
	- **DC** and AC conductivities

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<span id="page-46-0"></span>Improvements over Drude's theory

#### Mean free path

\n- From 
$$
l = v_F \tau \rightarrow l = \frac{(r_s/a_0)^2}{\rho_\mu} \times 92 \text{ Å}
$$
\n- $l \sim 100 \text{ Å}$  are possible at room T
\n

#### Thermal conductivity

 $\bullet$ 

From 
$$
\kappa = \frac{1}{3}v^2 \tau c_v
$$
  
\n
$$
\frac{\kappa}{\sigma T} = \frac{\pi^2}{3} (\frac{k_B}{e})^2 = 2.44 \times 10^{-8} \text{watt} \cdot \text{ohm}/K^2
$$

#### excellent agreement with exp.

目

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 $\left\{ \begin{array}{ccc} 1 & 0 & 0 \\ 0 & 1 & 0 \end{array} \right.$ 

<span id="page-47-0"></span>Improvements over Drude's theory

**Thermopower** 

With Sommerfeld estimate of the specific heat:

$$
Q = -\frac{\pi^2}{6} \frac{k_B}{e} \left( \frac{k_B T}{\varepsilon_F} \right) = -1.42 \left( \frac{k_B T}{\varepsilon_F} \right) \times 10^{-4} \text{volt/K}
$$

smaller by  $O(\frac{k_BT}{\varepsilon_F})\sim 0.01$  at room  $\sf T$ 

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