The Aurora



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HISTORY

- China ~ "Flying Dragons" (2000 aC)
- Bible

(Antico Testamento, Ezechiele "Ed ecco un vento tempestoso avanzarsi dal Settentrione, una grande nube che splendeva tutt'intorno, un fuoco in cui guizzavano bagliori, e nel centro come lo splendore dell'eletto in mezzo al fuoco")

- Ancient Greece (Senofane "Cumuli di nubi ardenti")
- Halley ~ "Orientation of Auroral Curtains aligned with Earth's Magnetic Field" (1716, Marzo)
- Loomise ~ "Auroral Ovals" (1859)

HISTORY

- Becqerel ~ "Solar Particles" (1878)
- Birkeland ~ "Field Alligned Currents" (1902-3)
- Störmer ~ "Motion of Charged Particles in Earth's Magnetic Field, Height of the Aurora" (1907)
- Vegard ~ "Proton Aurora" (1939)
- Anger ~ "First Global Space-Based Auroral Image (ISIS-2)" (inizio anni '70)

The launch of more advanced spacecraft (Dynamics Explorer, Viking, Polar e Image) made it possible to obtain the global auroral distribution

SOLAR WIND

Flow of charged particles - *plasma* - streaming from the Sun's *corona* in *all directions*



Parameter	Min	Av	Max
Flux (cm ⁻² s ⁻¹)	1	3	100
Vel (km/s)	200	400	900
Density (cm ⁻³)	0.4	6.5	100
Helium %	0	5	25
B (nT)	0.2	6	80

Comparison of the Solar Corona at Solar Maximum and Minimum (White Light Eclipse Images from the High Altitude Observatory)



February 16,1980 Solar Eclipse (Near Solar Maximum)



November 3, 1994 Solar Eclipse (Near Solar Minimum)

MAGNETOSPHERE

Region surrounding a planet where the planet's magnetic field dominates



WHAT CAUSES THE AURORA?

• The aurora is a visible manifestation of the *solar-terrestrial connection*



Solar wind particles, following the geomagnetic field lines, precipitate into the upper atmosphere where interact with atmospheric atoms

WHAT CAUSES THE AURORA?





 $h_{\rm V} = \Delta E$



AURORA SPECTRUM

The aurora is a source of X-rays,UV, IR, radio and optical radiation \rightarrow electron interacts with neutral atoms or molecules wich emits radiation due to disexitation



Green colour \longrightarrow atomic oxygen line at 557.7 nm, 100-200 Km altitude Red colour \longrightarrow atomic oxygen spectral triplet at 630.0, 636.4, 639.1 nm, >200 Km altitude, protons

Violet or Blue colour \longrightarrow molecular nitrogenline at 391.4, 427.0, 470.0 nm, <100 km altitude

AURORAL OVAL

Seen from space, northen and southern lights appear as oval shaped circle with the magnetic pole in the centre

Stretched out untisunward ~67° geomagnetic latitude

Compressed sunward ~78° geomagnetic latitude

THETA-AURORA



AURORA OCCURRENCE

Within the auroral zone, the aurora can be seen every clear winter night

- 27-day intervals
- more frequent in late autumn and early spring
- Northern lights activity corresponds closely to sunspot activity, which follows an 11-year cycle $\longrightarrow BUT$ 1 year delay
- 20-30% less during solar minimum than at solar maximum

THE AURORAL SUBSTORM



T = 0 min \rightarrow Quiet phase: weak arcs in the midnight regions at high geomagnetic latitude (~75°)

 $T = 0 \sim 5 \text{ min} \rightarrow \text{Oneset phase:}$ southernmost arc brightens and moves southward

T = 5~10 min → Expansion phase: the bright arc forms a bulge wich expands northward and moves westward at high speed

Akasofu, 1968

THE AURORAL SUBSTORM



T = 10-30 min → Expansion phase maximum: the bulge expandes westward and northward, irregular pulsating aurora on the morning side

 $T = 30 \sim 60 \text{ min} \rightarrow \text{Recovery}$ phase: weak arcs start to contract toward the north. Pulsating aurora at lower latitudes in the morning side

 $T = 1 \sim 2 \text{ hrs} \rightarrow \text{Recovery}$ phase: the situation retreats to the pre-onset condition

Akasofu, 1968

AURORAL STRUCTURE



- *Arcs*: green, sometimes with red above and purple below, quiet phase
- *Bands or Arcs with Structure*: green, sometimes with red above and purple below, expansion phase
- *Corona*: Geometric perspective effect, expansion phase
- *Diffuse Glows*: greenish at high latitudes, red at lower latitudes, recovery phase, faint

PROTON AURORA

Electron and proton auroras are different and develop differently over time



- equally bright but *less structured*
- Protons quickly become *neutralized* as they combine with electrons
- very important at the *start* of the substorm

EXPERIMENTS - IMAGE

Imager for Magnetopause-to-Aurora Global Exploration

Launch: 25 March 2000 Mission end: Dec 2005

Elliptical polar orbit: Apogee = $7.2 R_T$ Perigee = 1000 Km I = 90°

IMAGE was the first satellite mission dedicated to imaging the Earth's magnetosphere



EXPERIMENTS - THEMIS

Time History of Events and Macroscale Interactions during Substorms

- Launch Date: Feb. 17,2007
- 2-year mission
- 5 identical probes



Determine what physical process in near-Earth space initiates the violent eruptions of the aurora that occur during substorms



AURORA ON OTHER PLANETS

Aurora can occure on every planet or moon with a *megnetic field* and an *athmosphere*

The process of generating auroras is the same throughout the whole solar system even if the configuration, the colours and particularly the rapidly-varying displays, are different from what we see on earth





Saturn Aurora

REFERENCES

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Physics of the Upper Polar Atmosphere, Asgeir Brekke, WILEY

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