## from Lect.VII

- correlations
- (small or large) fluctuations


## Correlations

[a very simple calculation if the array points(n) is stored]

$$
\begin{aligned}
& \text { do } j=0, \quad j \max \\
& \text { si=0._dp ; si2=0._dp ; sij=0._dp } \\
& \text { do } i=1, n-j \\
& s i=s i+p o i n t s(i) \\
& \text { si2 = si2 + points(i)**2 } \\
& \text { sij }=s i j+\text { points(i)*points(i+j) } \\
& \text { end do } \\
& s i=s i /(n-j) \\
& s i 2=s i 2 /(n-j) \\
& s i j=s i j /(n-j)
\end{aligned}
$$

write(2,*), j, (sij-si**2)/(si2-si**2)
end do

## Correlations - Box-Muller algorithm



## Correlations - Metropolis algorithm




## Correlations - Metropolis algorithm



## Correlations - Metropolis algorithm



## Metropolis algorithm

by construction, Metropolis algorithm has short-range correlations. But it is very general!
and particularly useful for applications in the
canonical ensemble :
fix $\mathrm{N}, \mathrm{V}, \mathrm{T}$ (system in equilibrium with a thermal bath)

# Metropolis algorithm in the canonical ensemble 

## Microstates \& macrostates

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## The canonical ensemble

$(\mathbf{N}, \mathbf{V}, \mathbf{T})$ fixed. The probability $\boldsymbol{P}_{s}$ for a system to be in the microstate $s$ depends only on the energy $\boldsymbol{E}_{s}$ of the microstate and the temperature $\boldsymbol{T}$ (common to the system and the thermal bath):

$$
P_{s}=\frac{1}{Z} e^{-\beta E_{s}},(\text { canonical distribution })
$$

where $\beta=1 / k T$, and $Z$ is a normalization constant.
(in the canonical ensemble the energy is a characteristic of the microstate, not of the macrostate) The higher is the energy of the state, the lower is the probability of finding the system in that state.

Because $\sum P_{s}=1$,

$$
Z=\sum_{s=1}^{M} e^{-E_{s} / k T}
$$

(canonical partition function)
(M: all accessible microstates of the system, having the same $\mathrm{N}, \mathrm{V}, \mathrm{T}$ )

## Averages in the canonical ensemble

The ensemble average of a physical quantity $A$ can be calculated using the canonical distribution function:

$$
\langle A\rangle=\sum_{s=1}^{M} A_{s} P_{s}=\frac{1}{Z} \sum_{s=1}^{M} A_{s} e^{-\beta E_{s}}
$$

(M:
total number of possible microstates)

In practice, approximating the sums over a finite number $m$ of the total number $M$ of accessible microstates:

$$
\langle A\rangle \approx A_{m}=\frac{\sum_{s=1}^{m} A_{s} e^{-\beta E_{s}}}{\sum_{s=1}^{m} e^{-\beta E_{s}}}
$$

(Note: m, not M !)

## Averages in the canonical ensemble

Due to the functional form of $e^{-\beta E_{s}}$, this formulation is highly inefficient (many microstates have very small probability):

$$
\langle A\rangle \approx A_{m}=\frac{\sum_{s=1}^{m} A_{s} e^{-\beta E_{s}}}{\sum_{s=1}^{m} e^{-\beta E_{s}}}
$$

=> much better to use the importance sampling, i.e. simple averages of the desired physical quantity over microstates $s$ generated according to the proper distribution:

$$
A_{m}=\frac{1}{m} \sum_{s=1}^{m} A_{s}
$$

## Averages in the canonical ensemble

How to generate microstates properly distributed ?
=> generate a Markov chain with the transition matrix:

$$
T_{s, \text { trial }}=\min \left[1, \frac{P_{\text {trial }}}{P_{s}}\right]=\min \left[1, \frac{e^{-\beta E_{\text {trial }}}}{e^{-\beta E_{s}}}\right]=\min \left[1, e^{-\beta \Delta E}\right]
$$

where $\Delta E=E_{\text {trial }}-E_{s}$
Peculiarities w.r.t. the application of the Metropolis algorithm already discussed:
$e^{-\beta E_{s}}$ is a monotonic decreasing function of $E_{s}$, therefore we accept all the moves decreasing the energy.

## Metropolis algorithm in the canonical ensemble

1. Establish an initial microstate.
2. Make a random trial change in the microstate. For example, choose a spin at random and flip it. Or choose a particle at random and displace it a random distance.
3. Compute $\Delta E \equiv E_{\text {trial }}-E_{\text {old }}$, the change in the energy of the system due to the trial change.
4. If $\Delta E$ is less than or equal to zero, accept the new microstate and go to step 8 .
5. If $\Delta E$ is positive, compute the quantity $w=e^{-\beta \Delta E}$.
6. Generate a random number $r$ in the unit interval.
7. If $r \leq w$, accept the new microstate; otherwise retain the previous microstate.
8. Determine the value of the desired physical quantities.
9. Repeat steps (2) through (8) to obtain a sufficient number of microstates.
10. Periodically compute averages over microstates.

## Metropolis algorithm in the canonical ensemble

## A few remarks:

I) Because it is necessary to evaluate only the ratio $P_{j} / P_{i}=e^{-\beta \Delta E}$, it is not necessary to normalize the probability.
$\left(P_{j} / P_{i}=\pi_{j} / \pi_{i}\right)$

## 2) ERGODICITY implicitly assumed!

## 3) TEMPERATURE:

If $E_{B}>E_{A}$, accept the new (higher energy) configuration with probability $p=$ $e^{-\left(E_{B}-E_{A}\right) / T}$. This means that when the temperature is high, we don't mind taking steps in the "wrong" direction, but as the temperature is lowered, we are forced to settle into the lowest configuration we can find in our neighborhood.

## Metropolis algorithm in the canonical ensemble

Don't confuse $P_{s}$ with $\mathbf{P}\left(\mathbf{E}_{\mathrm{s}}\right)$, since $\boldsymbol{\Omega}\left(\mathrm{E}_{\mathrm{s}}\right)$ microstates with the same energy $\mathbf{E}_{\mathbf{s}}$ can exist, in general (unless in particular cases where $E_{s}$ uniquely identifies the microstate). Therefore:

$$
P\left(E_{s}\right)=\frac{\Omega\left(E_{s}\right)}{Z} e^{-\beta E_{s}}
$$

where the degeneracy of the energy level is accounted for, whereas:

$$
P_{s}=\frac{1}{Z} e^{-\beta E_{s}}
$$

## Some programs:

on
\$/home/peressi/comp-phys/VIII-canonical/ [do: \$cp /home/peressi/.../VIII-canonical/* .]

## boltzmann_metropolis.f90

and also (see later):
box.f90
entropy.f90
simulated_annealing.f90

## Boltzmann distribution in the canonical ensemble

The Metropolis algorithm really produces microstates with the Boltzmann distribution: application to ideal classical ID gas (ex. n. 4 Lect.VII) 1 free particle: Energy: $E=\frac{1}{2} m v^{2}$ in this case, velocity or energy labels a microstate (the energy with a factor of 2 , due to $+/-$ sign of $v$ ); we generate different microstates by random variations of the velocity and we accept/reject with Metropolis

Important quantities are the probabilities:
$P(v) d v$ that the system has a velocity between $v$ and $v+d v$ or $P(E) d E$ that the system has an energy between $E$ and $E+d E$

## ideal classical ID gas

A particle moving randomly has in each direction a distribution of the component of the velocity:

$$
\begin{align*}
f\left(v_{x}\right) & =\left(\frac{m}{2 \pi k_{B} T}\right)^{1 / 2} e^{-m v_{x}^{2} / 2 k_{B} T}  \tag{1}\\
\left\langle v_{x}^{2}\right\rangle & =\int_{-\infty}^{+\infty} v_{x}^{2} f\left(v_{x}\right) d v_{x}=\frac{k_{B} T}{m} \tag{2}
\end{align*}
$$

In 1D:

$$
f(v) 2 d v=P(E) d E
$$

that gives: $P(E)=\frac{1}{\left(\pi k_{B} T\right)^{1 / 2}} \frac{1}{\sqrt{E}} e^{-E / k_{B} T}$

In 3D:

$$
\begin{equation*}
P(E)=\frac{2}{\sqrt{\pi}} \frac{1}{\left(k_{B} T\right)^{3 / 2}} \sqrt{E} \exp \left(-\frac{E}{k_{B} T}\right) \tag{3D}
\end{equation*}
$$

## Boltzmann distribution

## in the canonical ensemble

```
# T
    # < <E
    # <v0>
    # dvmax : 2.00000
    # deltaE : 5.000000E-02
    # nbin : 79
==> boltzmann.1K <==
    # nMCsteps: 1000
    # <E> : . }50126
    # <v> : 7.456664E-02
    # accept. : . }69200
    # sigma : .713780
\sigma/\sqrt{}{n}=0.022
==> boltzmann.10K <==
    # nMCsteps: 10000
    # <E> : . }50758
    # <v> : 3.366172E-02
    # accept. : . }70770
    # sigma : .726145
\sigma/\sqrt{}{n}=0.007
==> boltzmann.1M <==
    # nMCsteps: 1000000
    # <E> : . }50013
    # <v> : 1.833840E-04
    # accept. : . }69383
    # sigma : .707472
\[
\sigma / \sqrt{n}=0.0007
\]
```


## note:

```
- Accuracy of ~ \(1 \%\) on <E> and \(10 \%\) on <v> : NMCS=1000 is enough
- NOT ENOUGH to well reproduce the BOLTZMANN DISTRIBUTION! (1M needed!)
- ACCEPTANCE RATIO: constant, depends only on dvmax
- SIGMA also

\section*{Boltzmann distribution in the canonical ensemble}
many particles: Energy: \(E=\sum_{i=1}^{N} \frac{1}{2} m_{i} v_{i}^{2}\) in this case, the energy is NOT a label of a microstate (there are several microstates with the same total energy)

Note: the energy histogram is NOT the distribution of microstates!
\(P(E)=\sum_{\substack{\text { states } \\ \text { with } E_{s}=E}} P_{s}\) with \(P_{s}=\frac{1}{Z} e^{-\beta E_{s}}\)
\(P(E) \propto=e^{\frac{-(E-\langle E\rangle)^{2}}{2 \sigma^{2}}}\) with \(\langle E\rangle\) average over all the microstates
What is \(\mathrm{P}(\mathrm{E})\) ? (exercise n. 4 Lec.VII - many particles)
\(\mathrm{T}=1\), \(N=18\), dvnack \(=1\), nHCstepa \(=16^{*} \ldots\)

\(\mathrm{T}=2, \mathrm{~N}=18\), dvnack \(=1\), nifstepa \(=10^{\mathrm{N}} \ldots\)


\title{
Statistical averages and stochastic fluctuations
}

\section*{Averages}
\[
<G>_{T}=\frac{1}{T} \sum_{t=1}^{T} G_{t}
\]

Thermally equilibrated averages:


But in practice \(T\) is finite, and \(<G>_{T}\) oscillates(varying \(T\) ): divide \(T\) into intervals \(A, B, C \ldots\) of length \(L\) and sum(block averages):
\(<G>_{T}=\frac{L}{T} \sum_{I=A, B, C, \ldots}<G>^{(I)}\)
\[
\left(\sigma_{s} / \sqrt{s}\right)
\]
\[
\left(\sigma_{n} / \sqrt{n}\right)
\]
\(\Delta<G>_{T}=\left[\frac{L}{T} \sum_{I}\left(<\left(G^{(I)}\right)^{2}>-\left(<G^{(I)}>^{2}\right)\right]^{1 / 2} \longrightarrow \rightarrow \sim \frac{1}{\sqrt{T}}\right.\)
Note: not always \(\Delta<G>_{T}\) is a good indicator of the actual error! (remind "ergodicity")

\section*{Stochastic fluctuations}

\section*{Fluctuations are always present, due to the nature of the system, also when evolving towards equilibrium.}

A simple example: non-interacting classical particles in a box (gas diffusion)


A box is divided into two parts communicating through a small hole. One particle randomly can pass through the hole per unit time, from the left to the right or viceversa.
\(\mathrm{N}_{\text {left }}(\mathrm{t})\) : number of particles present at time t in the left side Given \(N_{\text {left }}(0)\), what is \(N_{\text {left }}(\mathrm{t})\) ?


\section*{Stochastic fluctuations}


How to reduce fluctuations?
- more particles
- average over many simulation runs

What can we do with fluctuations?

\section*{Intrinsic energy fluctuations in the canonical ensemble - I}
 Consider the thermal capacity:
we have
\[
C_{v}=\frac{\partial\langle E\rangle}{\partial T}=\ldots=-\frac{1}{k T^{2}} \frac{\partial\langle E\rangle}{\partial \beta}
\]
\[
\begin{array}{r}
\frac{\partial\langle E\rangle}{\partial \beta}=-\frac{1}{Z^{2}} \frac{\partial Z}{\partial \beta} \sum_{s} E_{s} e^{-\beta E_{s}}-\frac{1}{Z} \sum_{s} E_{s}^{2} e^{-\beta E_{s}}=\langle E\rangle^{2}-\left\langle E^{2}\right\rangle=-\left\langle(\delta E)^{2}\right\rangle \\
\text { where } \quad \delta E \equiv E-\langle E\rangle
\end{array}
\]

Result:
\[
C_{v}=\frac{\left\langle(\delta E)^{2}\right\rangle}{k_{B} T^{2}}
\]

The thermal capacity (or specific heat if considered for each particle) is related to the intrinsic stochastic energy fluctuations

\section*{Intrinsic energy fluctuations in the canonical ensemble - II}

Since:
\[
C_{v}=\frac{\left\langle(\delta E)^{2}\right\rangle}{k_{B} T^{2}}
\]
if \(N\) is the number of particles, we have:
\[
\frac{\sqrt{\left\langle(\delta E)^{2}\right\rangle}}{\langle E\rangle}=\frac{\sqrt{k_{B} T^{2} C_{v}}}{\langle E\rangle} \propto \frac{\sqrt{N}}{N} \sim \frac{1}{\sqrt{N}}
\]
i.e., the relative energy fluctuations reduce when \(N\) is large
(correct; in the thermodynamic limit: \(E \rightarrow\) const., macro \(\sim\) micro)

Macrostates, microstates and entropy at equilibrium

\title{
Macroscopic systems towards equilibrium
}

Again the simple example of non-interacting classical particles in a box (gas diffusion)


A box is divided into two parts communicating through a small hole. One particle randomly can pass through the hole per unit time, from the left to the right or viceversa.
\(\mathrm{N}_{\text {left }}(\mathrm{t})\) : number of particles present at time t in the left side Given \(N_{\text {left }}(0)\), what is \(N_{\text {left }}(\mathrm{t})\) ?
microcanonical ensemble :
fix E, N,V (isolated closed system)

\section*{Microcanonical ensemble -a fundamental postulate-}

Given an isolated system in equilibrium, it will assume with equal probability each of its accessible microstates
(i.e., a system in equilibrium does not have any preference for any of its available microstates)

Example: suppose a macrostate defined by (N,V,E); if \(\Omega(N, V, E)\) is the \# of microstates with energy \(E \div d E\), for the microstate \(s\) the probability of occurrence is:
\[
P_{s}= \begin{cases}1 / \Omega, & \text { if } s \text { is accessible } \\ 0, & \text { otherwise }\end{cases}
\]

Consequence: for a system at equilibrium, the thermodynamic state (macrostate) which could result from the largest number of microstates is also the most probable macrostate of the system.

\section*{Macroscopic systems towards equilibrium}


Another version: particles blue/red in both sides
(interdiffusion of two gases):
per unit time, one from each side is picked at random and put in the other side: \(\mathrm{N}_{\text {left }}{ }^{\text {blue }}(\mathrm{t})+\mathrm{N}_{\text {left }}{ }^{\text {red }}(\mathrm{t})=\) constant; \(\mathrm{N}_{\text {left }}{ }^{\text {red }}(\mathrm{t})=\) ?

\section*{Approach to equilibrium}


1


3


2


4

A box divided into left/right; total number N of non-interacting particles.
macrostate: specified by the number of particles \(n\) on the left side;
microstate: specified by the specific list of the \(n\) particles on the left side

\section*{Equilibrium and entropy}

The macrostate is specified by the number of particles on the left side, \(n\), and the corresponding microstates are:
\[
\text { number of microstates }=\frac{N!}{n!(N-n)!}=\binom{N}{n}=\binom{N}{N-n}
\]

The total number of microstates for \(\mathrm{N}=10\) is \(2^{10}=1024\)
\begin{tabular}{|c|c|c|}
\hline n & \# of microstates & \(\log (\#\) of micr. \\
\hline 0 & 1 & 0.00 \\
\hline 1 & 10 & 2.30 \\
\hline 2 & 45 & 3.81 \\
\hline 3 & 120 & 4.79 \\
\hline 4 & 210 & 5.35 \\
\hline 5 & 252 & 5.53 \\
\hline 6 & 210 & 5.35 \\
\hline 7 & 120 & 4.79 \\
\hline 8 & 45 & 3.81 \\
\hline 9 & 10 & 2.30 \\
\hline 10 & 1 & 0.00 \\
\hline
\end{tabular}
\(\mathrm{n}=5\) is the macrostate with the largest number of microstates!
( \(\mathrm{n}=5\) is the most probable macrostate)
the most"random"!
Equilibrium =
Maximum number of possible microstates \(=\) Maximum entropy

\section*{Equilibrium and entropy}

We could simply generate microstates and count them...
but also some smart algorithm (more suitable for computer simulations) are possible....

\section*{Entropy: Coincidence method (S.K. Ma, J. Stat. Phys. 26, 22 I (I98I))}

Equilibrium = Maximum entropy \(=\) Maximum number of possible microstates
Too much effort to enumerate all of them!

Alternative procedure (good for computing):
A system evolving in time will duplicate a microstate, before or later... The longer it takes for duplication, the fewer are the microstates in the corresponding macrostate. Hence, the lower is the entropy. Idea: measure the ratio of the number of pairs of duplicated microstates to the total number of possible pairs; entropy is the log of the inverse ratio.
E.g.: suppose as in the previous slide \(\mathrm{N}=10\), and the macrostate \(\mathbf{n = I}\); consider 20 different microstates labelled with the "name" of the particle:


Possible pairs: \(20^{*}(20-I) / 2=190\). Here: 6 pairs for particle "2"; I pair with particle " 10 " etc etc... Sum all of them: get 15 .

Ratio \(=15 / 190\), Entropy: \(S \propto \log (190 / I 5) \sim 2.5\)

Remind the definition of entropy:
\(S=-k_{B} \sum_{s} P_{s} \ln P_{s}\)
in the canonical ensemble
\(S=k_{B} \log \Omega \quad\) in the microcanonical ensemble, where all the microstates corresponding to a macrostate have the same energy
( \(\Omega\) is the number of microstates)

\section*{Metropolis method in the canonical ensemble and the simulated annealing}

\section*{Metropolis and simulated annealing - I}
- Stochastic search for global minimum. Monte Carlo optimization.
- The concept is based on the manner in which liquids freeze or metals recrystallize. Sufficiently high starting temperature and slow cooling are important to avoid freezing out in metastable states.

\title{
Metropolis and simulated annealing - II
}
- Thermodynamic system at temperature T, energy E.
- Perturb configuration (generate a new one).
usual Metropolis procedure in the canonical ensemble
- Compute change in energy dE . If dE is negative the new configuration is accepted. If dE is positive it is accepted with a probability given by the Boltzmann factor : \(\exp (-d E / k T)\).
- The process is repeated many times for good sampling of configuration space.
- then the temperature is slightly lowered and the entire procedure repeated, and so on, until a frozen state is achieved at \(\mathrm{T}=0\).

\section*{Example}

\section*{in simulated_annealing.f90: minimization of}
\[
f(x)=(x+0.2) * x+\cos (14.5 * x-0.3)
\] considered as an energy function and using a fictitious temperature


Function to be minimized: \(\mathbf{f}(\mathbf{x})\); Starting point: \(\mathbf{x}, \mathbf{f x}=\mathbf{f}(\mathbf{x})\)
```

                    initial (high) temperature: temp
    Annealing schedule:
annealing temperature reduction factor:
number of steps per block:
'ad hoc' parameter for trial move:

```

\section*{temp}
tfactor (<1) nsteps
scale
```

DO WHILE (temp > 1E-5) ! anneal cycle

```
```

DO istep = 1, nsteps

```
DO istep = 1, nsteps
    CALL RANDOM_NUMBER(rand) ! generate 2 random numbers; dimension(2) :: rand
    CALL RANDOM_NUMBER(rand) ! generate 2 random numbers; dimension(2) :: rand
    x_new = x + scale*SQRT(temp)*(rand(1) - 0.5) ! stochastic move
    x_new = x + scale*SQRT(temp)*(rand(1) - 0.5) ! stochastic move
    fx_new = func(x_new) ! new object function value
    fx_new = func(x_new) ! new object function value
    IF (EXP(-(fx_new - fx)/temp) > rand(2)) THEN ! success, save
    IF (EXP(-(fx_new - fx)/temp) > rand(2)) THEN ! success, save
        fx = fx_new
        fx = fx_new
        x = x_new
        x = x_new
    END IF
    END IF
    IF (fx < fx_min) THEN
    IF (fx < fx_min) THEN
        fx_min = fx
        fx_min = fx
        x_min = x
        x_min = x
        PRINT '(3ES13.5)', temp, x_min, fx_min
        PRINT '(3ES13.5)', temp, x_min, fx_min
    END IF
    END IF
END DO
END DO
    temp = temp * tfactor ! decrease temperature
    temp = temp * tfactor ! decrease temperature

```

initial T: 10 ( }\mp@subsup{\textrm{K}}{\textrm{B}}{\prime}\mathrm{ units)
initial x: 1.000000
initial f(x): 1.137208

```
final T: 2.50315E-01
final \(x: \quad-1.95067 \mathrm{E}-01\)
final f(x):-1.00088E+00
```

