SEISMOLOGY I

Laurea Magistralis in Physics of the Earth and of the Environment

Wave propagation

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Wave propagation is ruled by:

- SUPERPOSITION PRINCIPLE
- REFLECTION
- REFRACTION
- DIFFRACTION
- DOPPLER EFFECT

GEOMETRICAL SPREADING





Before Maxwell's equations were developed Huygens and Fermat could describe the propagation of light using their empirically determined principles.

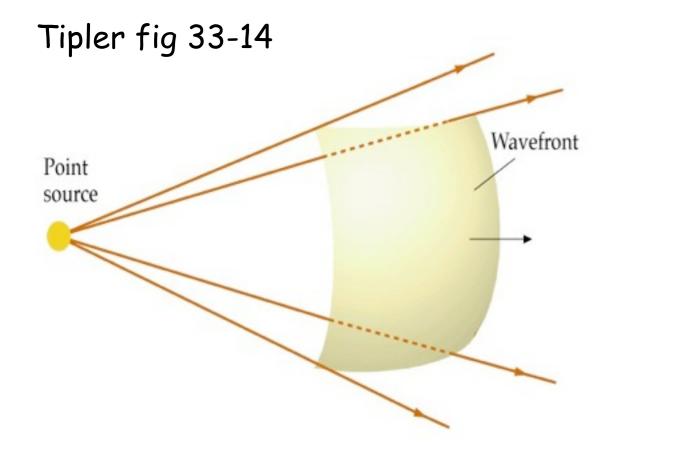
Huygens Principle

All points on a given wavefront are taken as point sources for the production of spherical secondary waves, called wavelets, which propagate outwards with speeds characteristic of waves in that medium.

After some time has elapsed the new position of the wavefront is the surface tangent to the wavelets.







Spherical wavefront from a point source

Huygens construction for spherical plane wave wave (b) (a)

Tipler fig 33-16





This is a general principle for determining the paths of light rays.

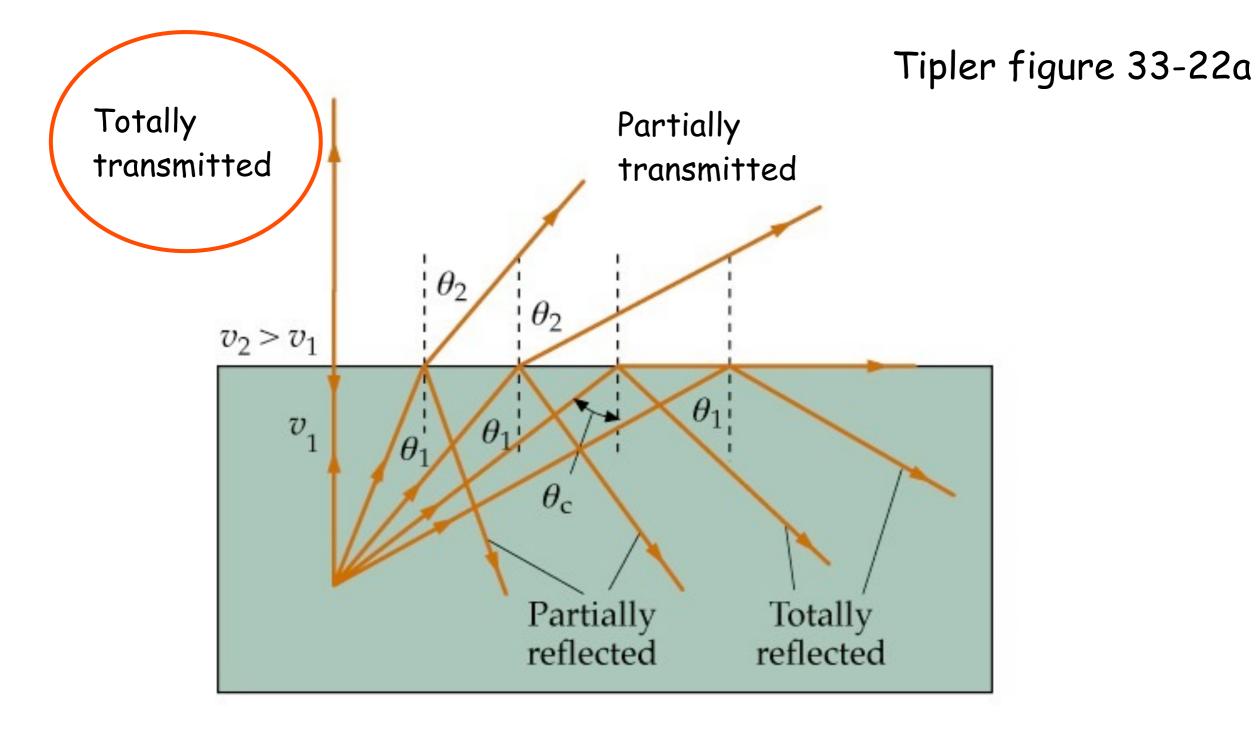
When a light ray travels between any two points P and Q, its actual path will be the one that takes the least time.

Fermat's principle is sometimes referred to as the "principle of least time".

An obvious consequence of Fermat's principle is that when light travels in a single homogeneous medium the paths are straight lines.











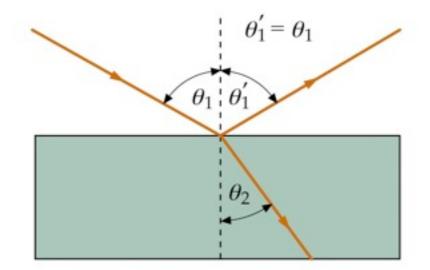
Index of refraction :

A transparent medium is characterised by the **index of refraction** n, where n is defined as the ratio of the speed of light in a vacuum c, to the speed of light in the medium v

$n = \frac{c}{v}$	air	n=1.0003
	water	n=1.33
	diamond	n=2.4

Law of reflection :

$$\theta_1 = \theta_1$$







For normal incidence at a boundary (ie $\theta_1 = \theta_1' = 0$)

the reflected intensity I is given by:

$$\mathbf{I} = \left(\frac{\mathbf{n}_1 - \mathbf{n}_2}{\mathbf{n}_1 + \mathbf{n}_2}\frac{1}{\dot{\mathbf{j}}} \mathbf{I}_o\right)$$

 \mathbf{I}_{o} = incident intensity, n_{1} and n_{2} are the refractive indexes of medium 1 and 2

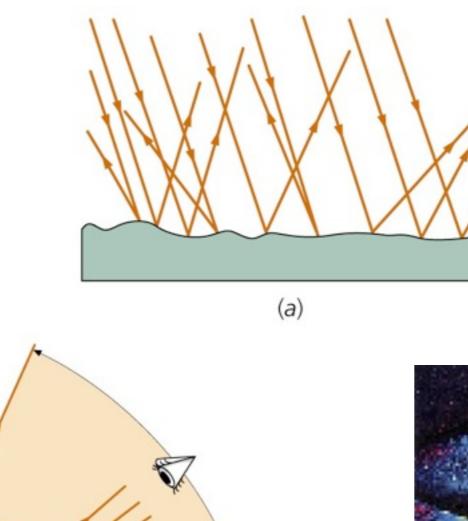
eg: for an air-glass interface, $n_1 = 1$ and $n_2 = 1.5$ giving $I = I_0/25$



Reflection of light







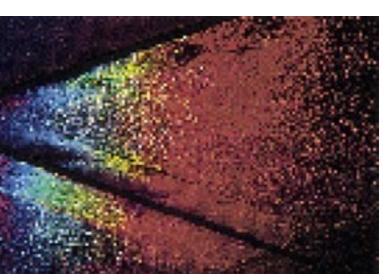
Diffuse reflection (Tipler fig 33-20)



reflection

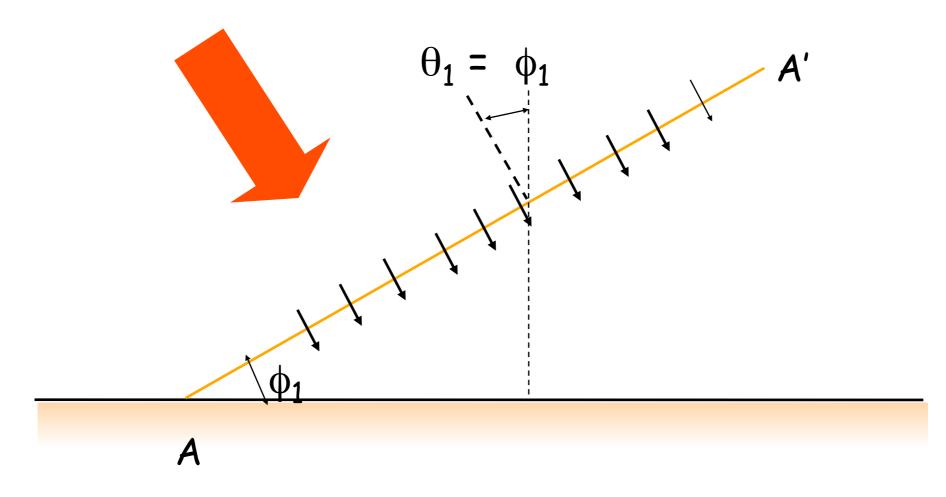
(Tipler fig 33-19)

Mirror







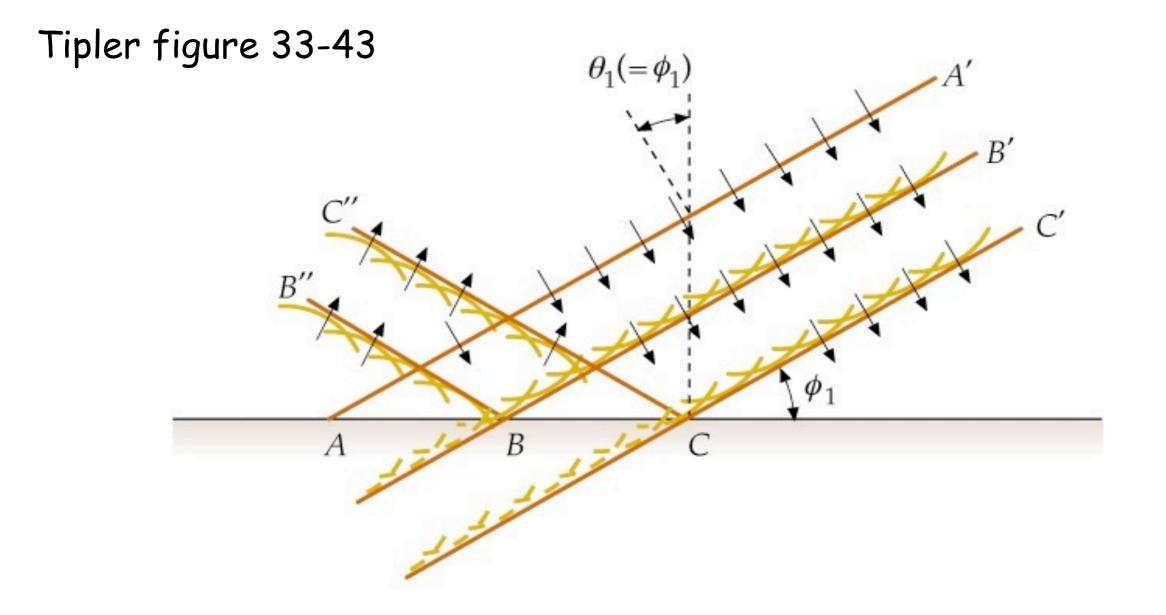


AA' is a wavefront of incident light striking a mirror at A

The angle between the wavefront and the mirror = the angle between the normal to the wavefront and the vertical direction (normal to the mirror)





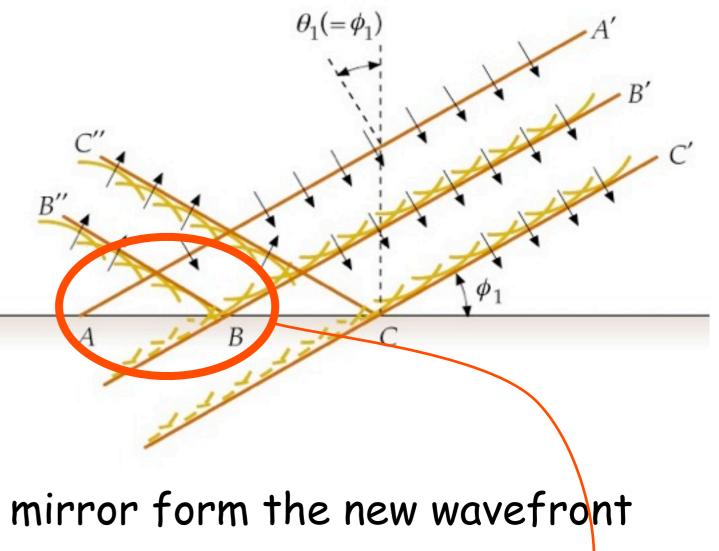


According to Huygens each point on the wavefront can be thought of as a point source of secondary wavelets





The position of the wavefront after a time t can be found by constructing wavelets of radius ct with centres on AA'

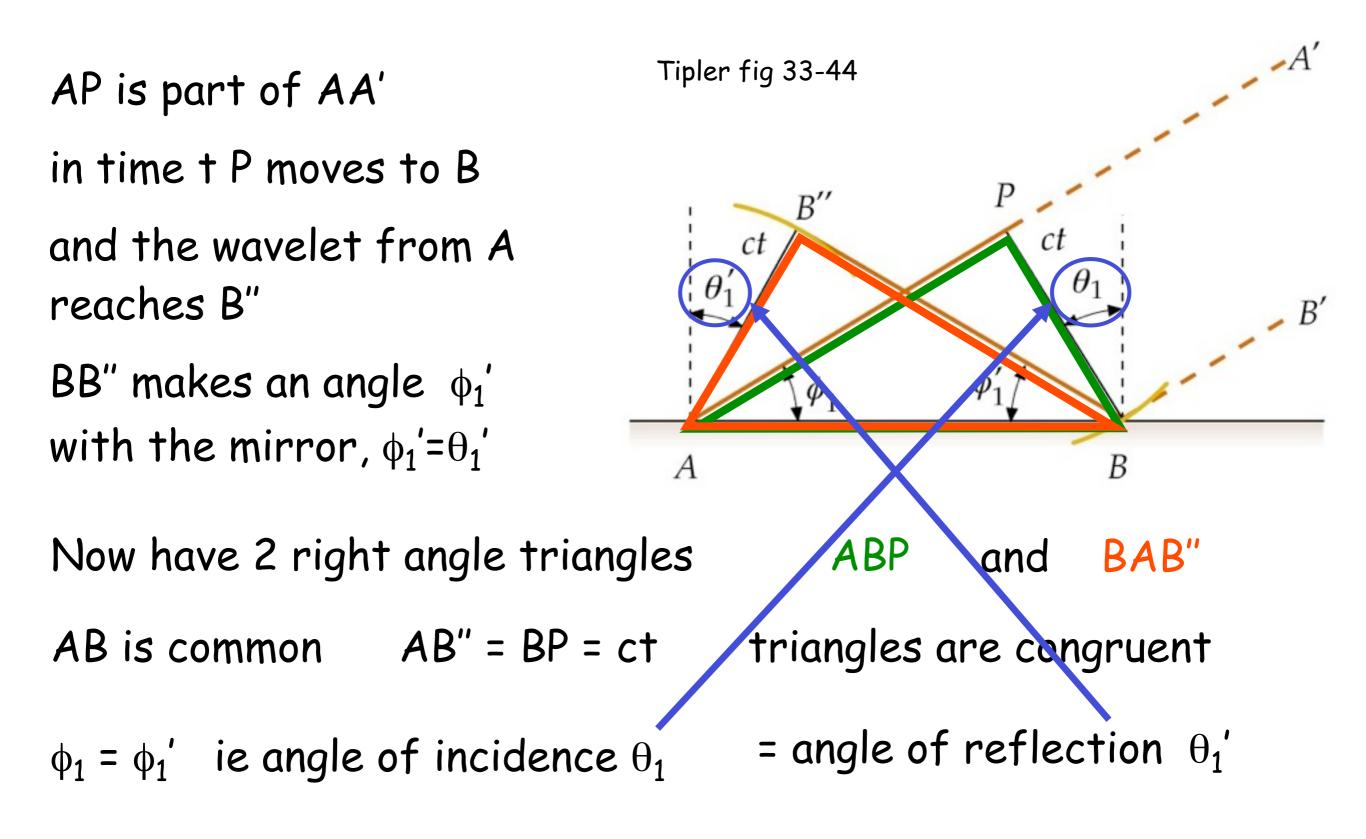


Waves that do not strike the mirror form the new wavefront BB' (and similarly CC')

Consider a small portion of these waves......





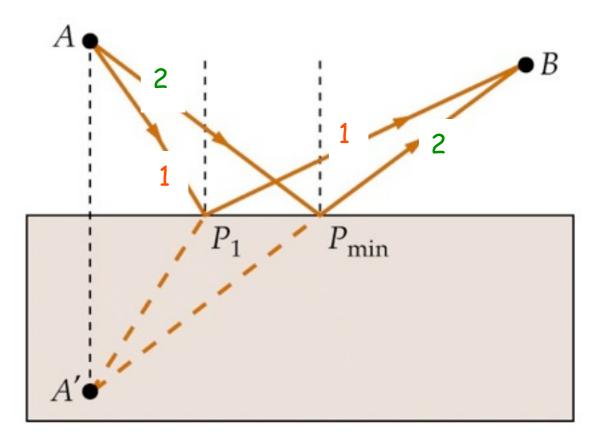






Light can travel from A to B (via mirror) on path 1 or path 2.

If we want to apply Fermat's principle we need to know at which point P the wave must strike the mirror so that APB takes the least time.



Tipler figure 33-46

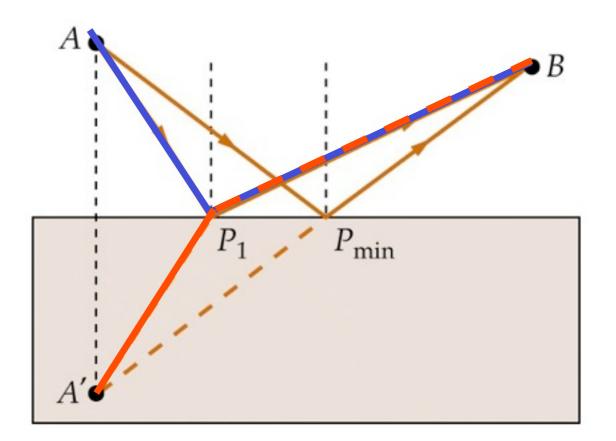
As light is travelling in the same medium at all times the shortest time will also be the shortest distance.





Where ever P is located the distance APB = A'PBwhere A' is the position of the image of the light source.

As the position of P varies the shortest distance will be when $A(P=P_{min})B$ lie in a straight line.



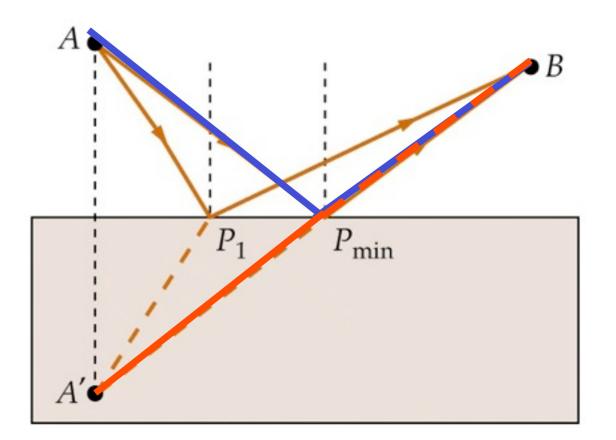
Tipler figure 33-46





Where ever P is located the distance APB = A'PBwhere A' is the position of the image of the light source.

As the position of P varies the shortest distance will be when $A'(P=P_{min})B$ lie in a straight line.

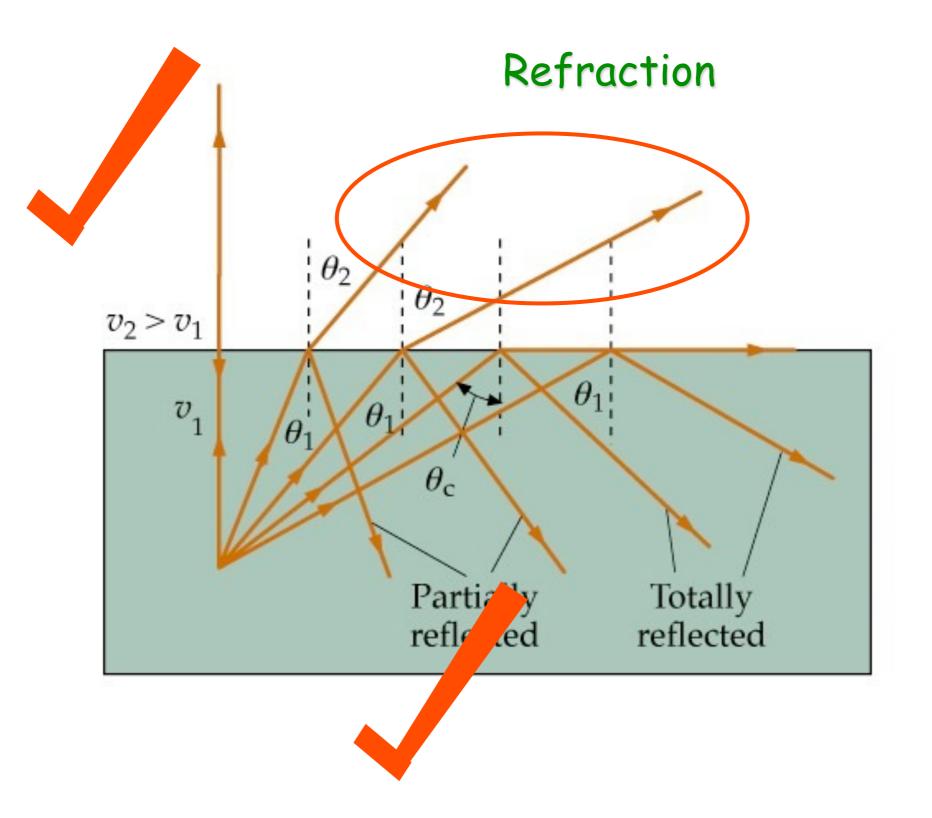


Tipler figure 33-46

This will be when the angle of incidence = angle of reflection









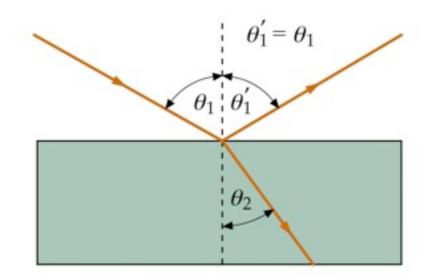


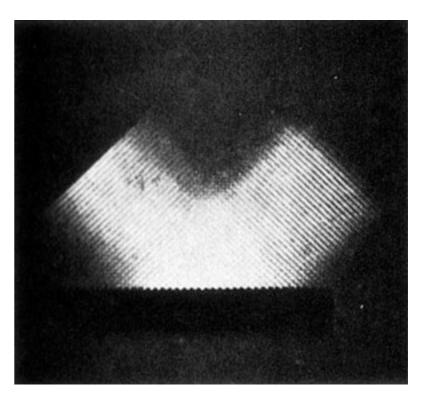
Snell's law of refraction :

if v is the speed of the light in the medium

$$\frac{1}{v_1}\sin\theta_1 = \frac{1}{v_2}\sin\theta_2$$

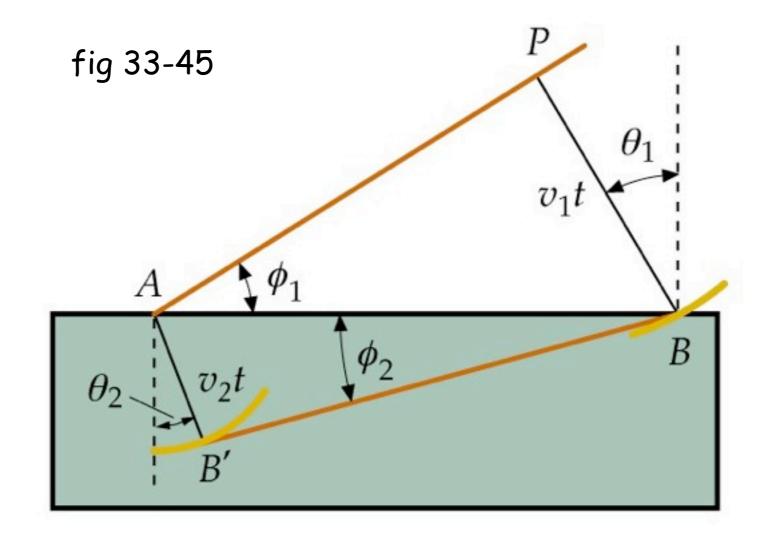
 $n_1 \sin \theta_1 = n_2 \sin \theta_2$









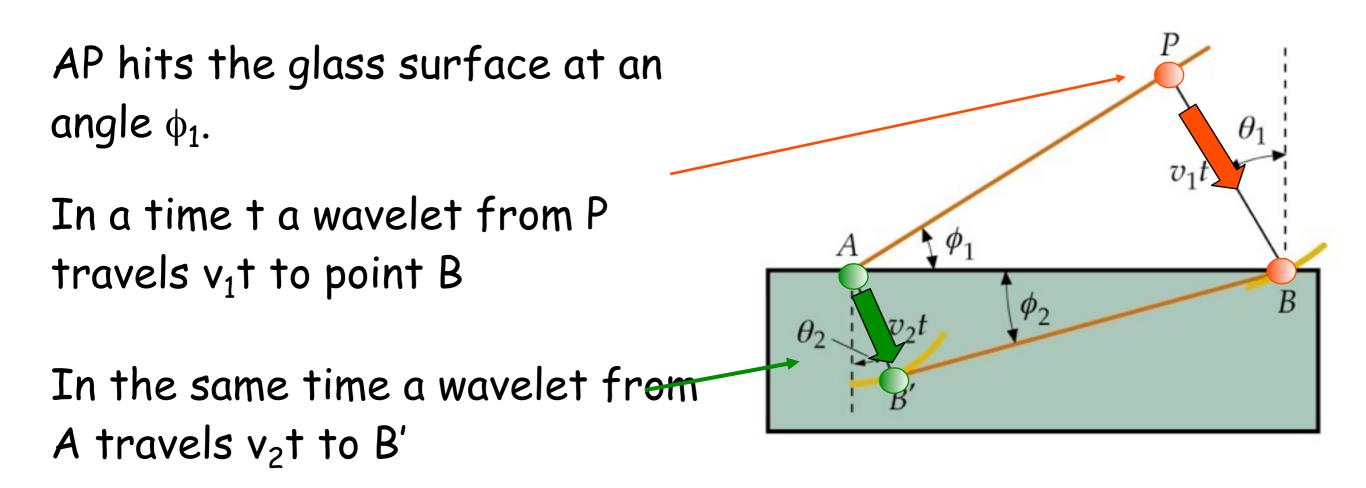


Consider a plane wave incident on a glass interface.

AP represents a portion of the incident wave - we can use Huygens' construction to calculate the transmitted wave.







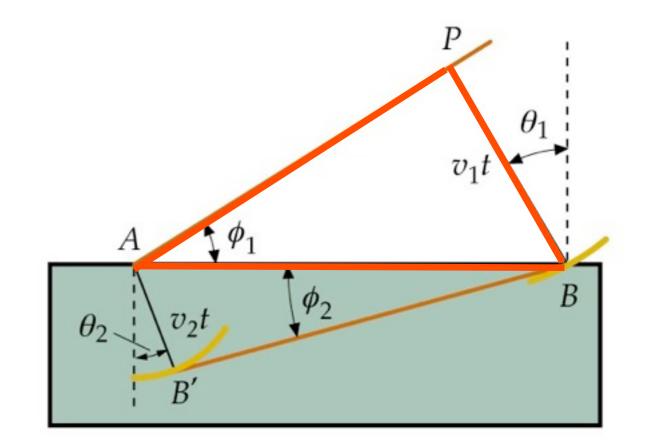




AP hits the glass surface at an angle ϕ_1 .

In a time t a wavelet from P travels v_1 t to point B

In the same time a wavelet from A travels v_2 t to B'



BB' is not parallel to AP because $v_1 \neq v_2$

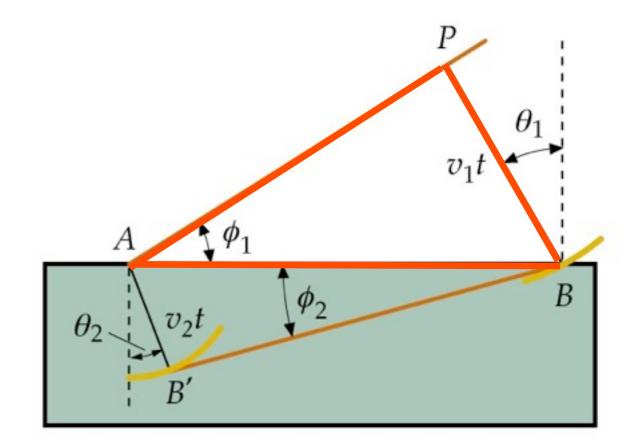
In
$$\triangle APB$$
 sin $\phi_1 = \frac{v_1 t}{AB}$ or $AB = \frac{v_1 t}{\sin \phi_1}$





$$AB = \frac{v_1 t}{\sin \phi_1}$$

but $\phi_1 = \theta_1$
$$\therefore AB = \frac{v_1 t}{\sin \theta_1}$$







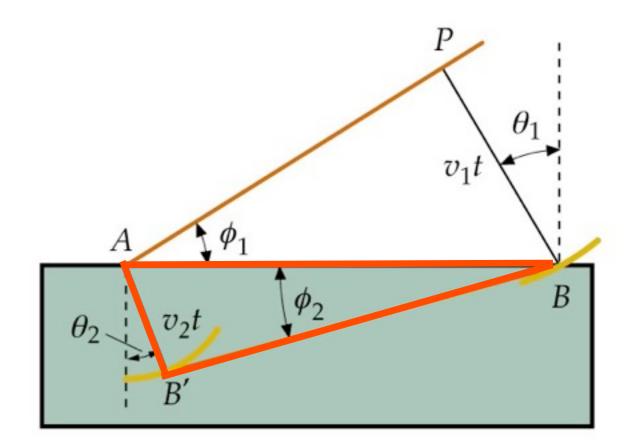
$$AB = \frac{v_1 t}{\sin \phi_1}$$

but $\phi_1 = \theta_1$
$$\therefore AB = \frac{v_1 t}{\sin \theta_1}$$

Similarly in $\Delta ABB'$

$$\sin \phi_2 = \frac{v_2 t}{AB}$$

or
$$AB = \frac{v_2 t}{\sin \theta_2}$$







$$\frac{v_1 t}{\sin \theta_1} = \frac{v_2 t}{\sin \theta_2}$$

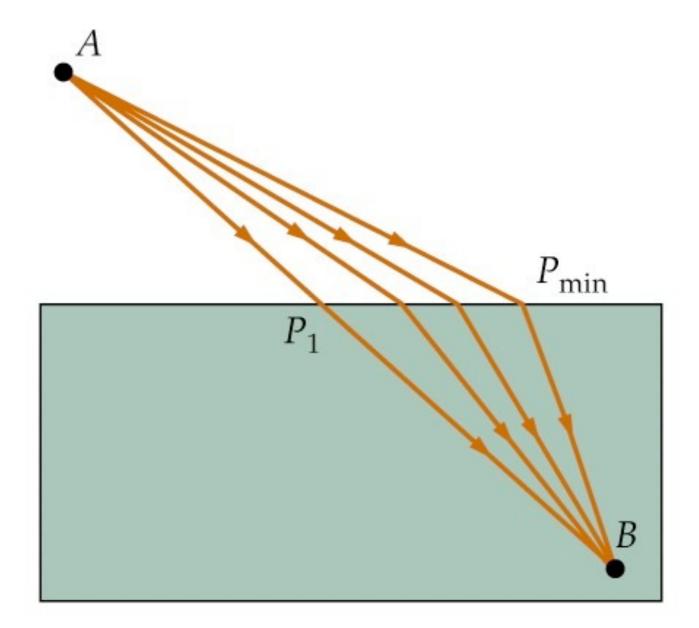
$$\frac{\sin\theta_1}{v_1} = \frac{\sin\theta_2}{v_2}$$

But $v_1 = c / n_1$ and $v_2 = c / n_2$

$$n_1 \sin \theta_1 = n_2 \sin \theta_2$$







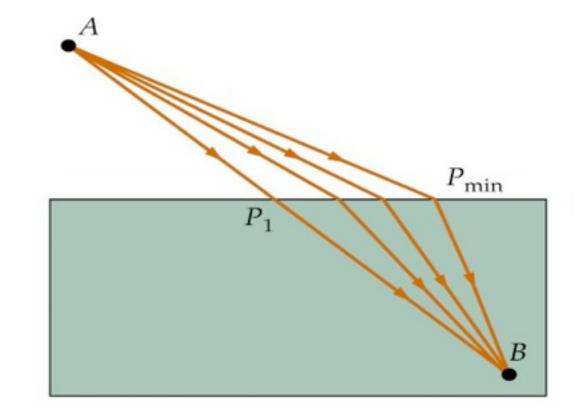
Can also use Fermat's principle to derive Snell's Law (more complicated but important)





Several possible paths from A (in air) to B (in glass)

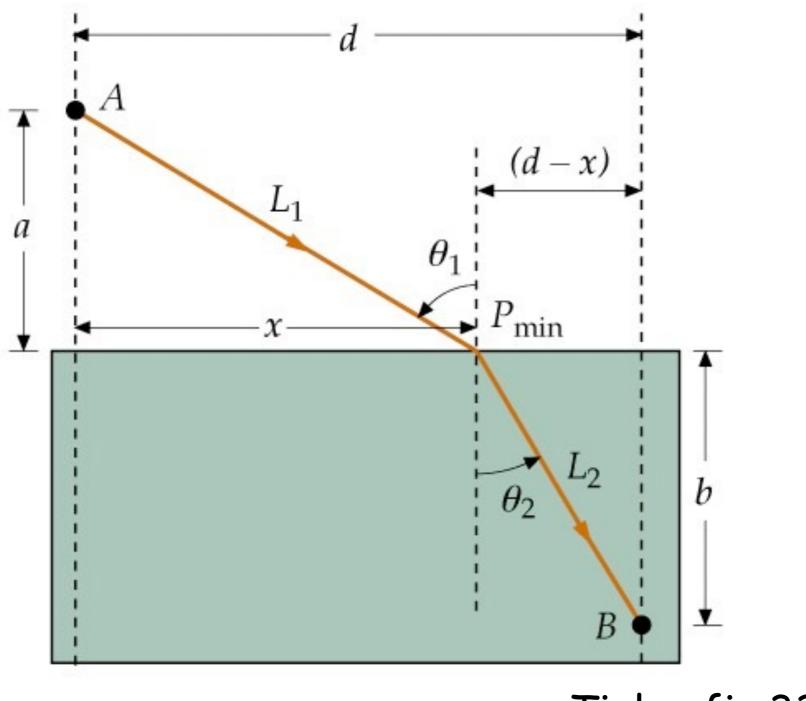
Remember that light travels more slowly in glass than in air so $A-P_1$ -B (straight line) will not have the shortest travel time.



If we move to the right the path in the glass is shorter, but the overall path is longer - how do we choose the shortest route ?



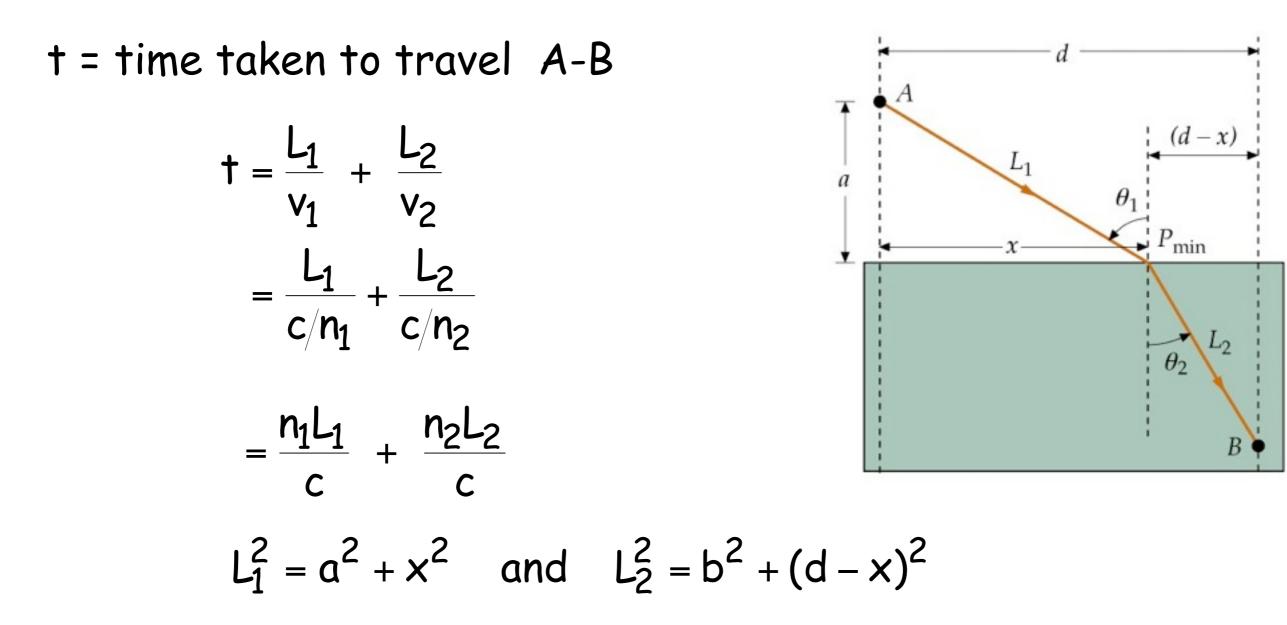




Tipler fig.33-48



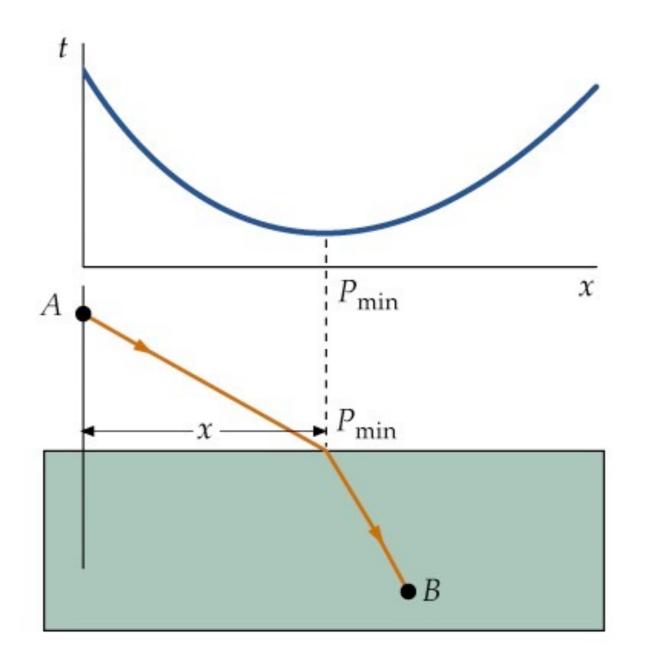




We can combine these three equations and plot the time as a function of x







Tipler fig.33-49

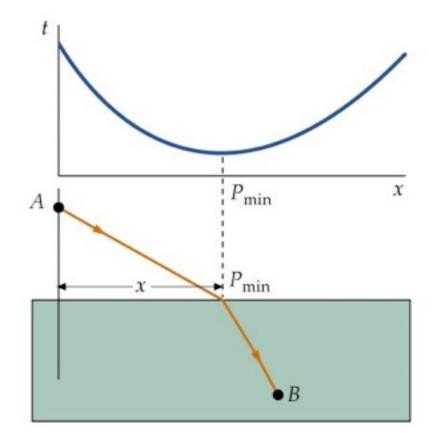




Find the minimum time taken by finding when

$$\frac{dt}{dx} = 0$$

$$\frac{dt}{dx} = \frac{1}{c} \left(\frac{n_1 dL_1}{dx} + \frac{n_2 dL_2}{dx} \right) = 0$$



$$\frac{dL_1}{dx} = \frac{1}{2}(a^2 + x^2)^{-1/2} \times 2x$$
$$= \frac{x}{L_1}$$

but
$$\frac{x}{L_1} = \sin \theta_1$$



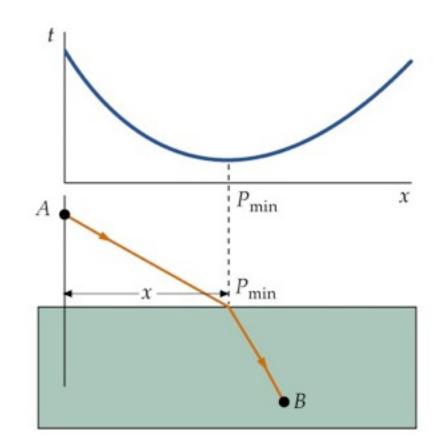


Similarly for L₂ $\frac{dL_2}{dx} = \frac{1}{2}(b^2 + (d - x)^2)^{-1/2} \times -2(d - x)$ $= -\frac{(d - x)}{L_2}$

but
$$-\frac{d-x}{L_2} = -\sin\theta_2$$

we want
$$\frac{dt}{dx} = 0$$

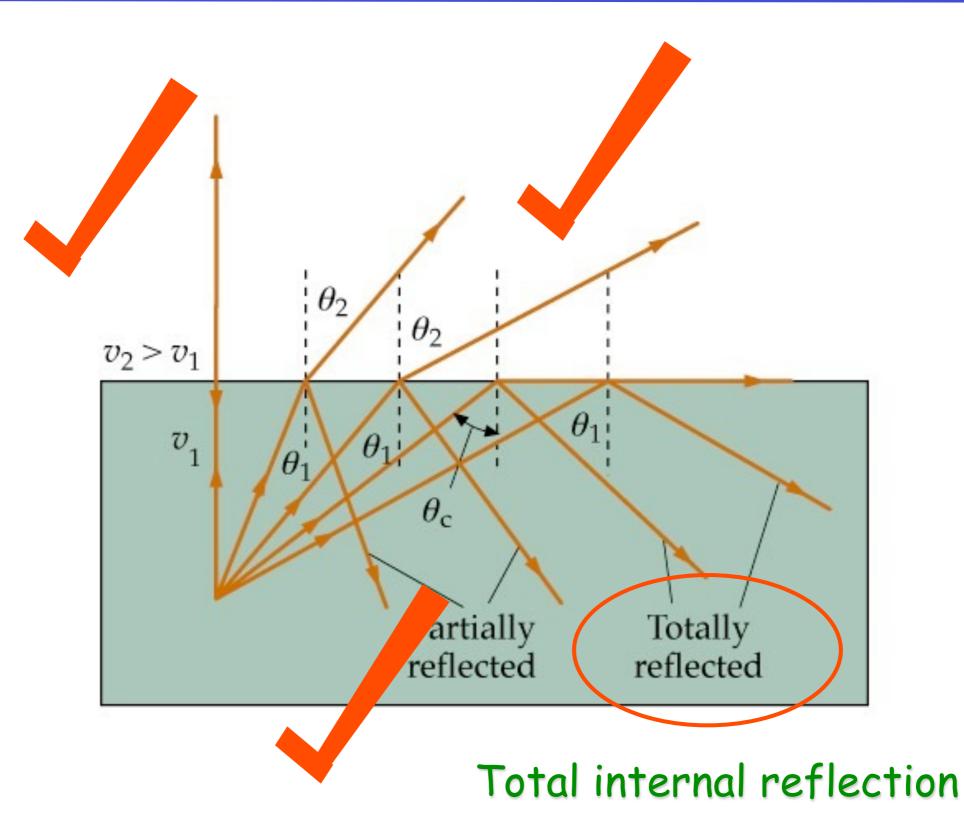
$$n_1 \sin \theta_1 + n_2(-\sin \theta_2) = 0$$



$$n_1 \sin \theta_1 = n_2 \sin \theta_2$$



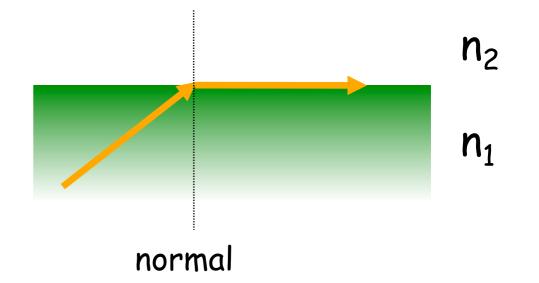








Total internal reflection occurs when light attempts to move from a medium of refractive index n_1 to a medium of refractive index n_2 and $n_1 > n_2$



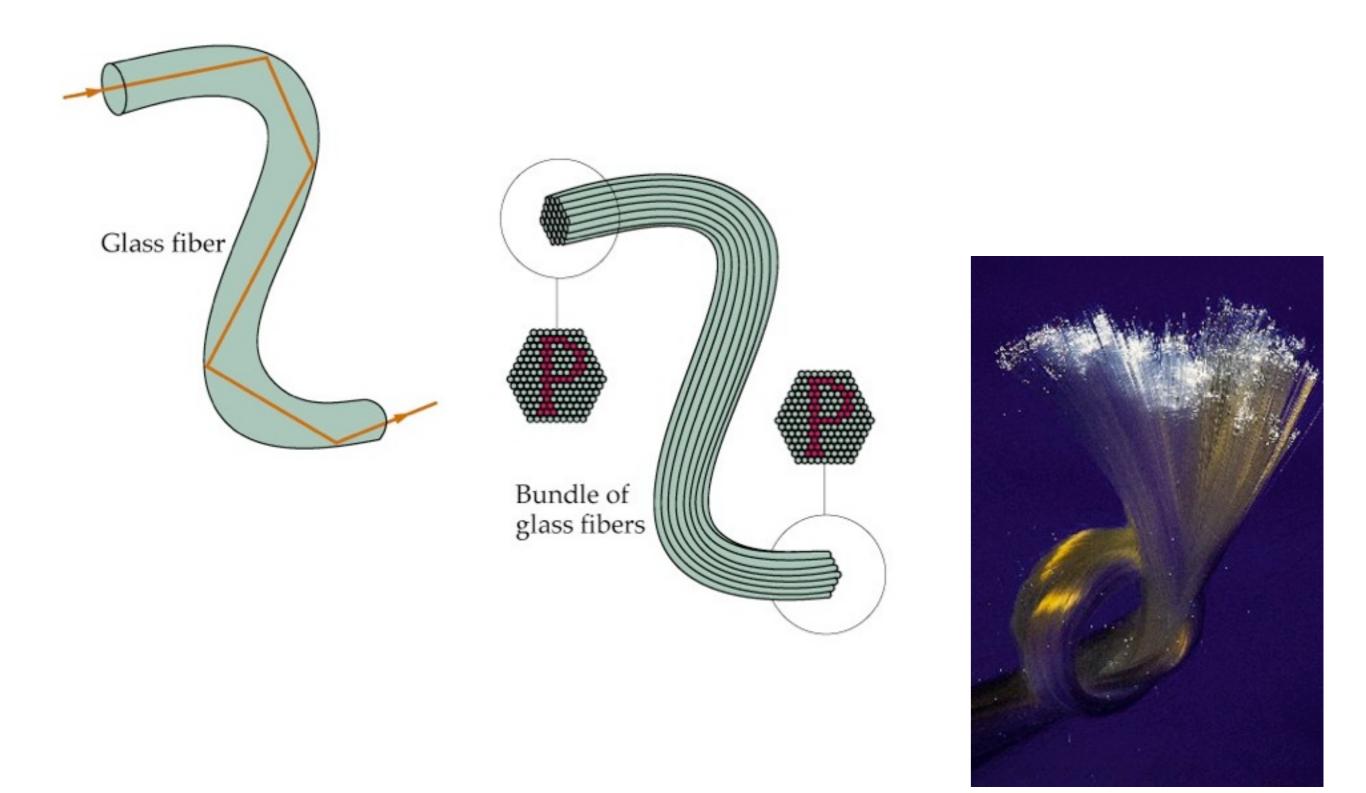
At a particular angle of incidence (the critical angle θ_c) the refracted ray will travel along the boundary ie $\theta_2 = 90^{\circ}$

If the angle of incidence > θ_c the light is entirely reflected



Optical Fibres

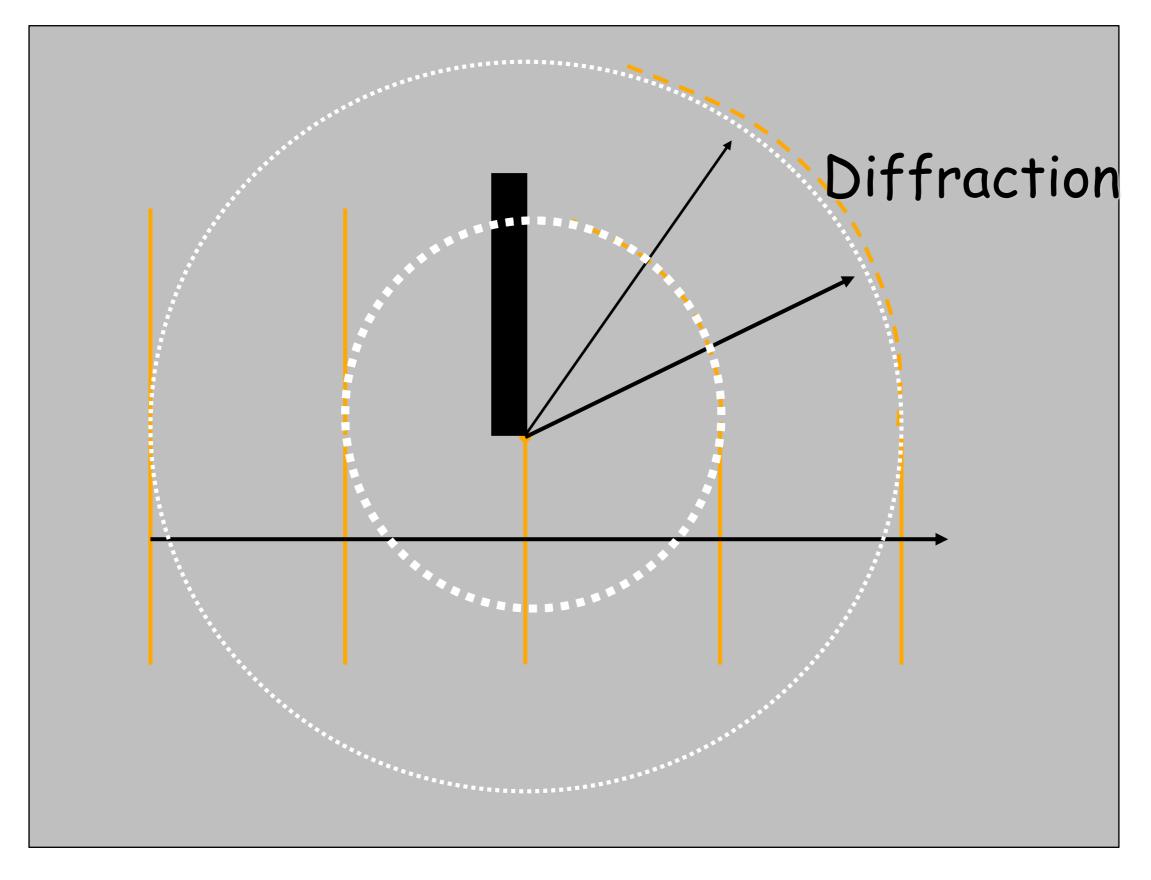






Visualization







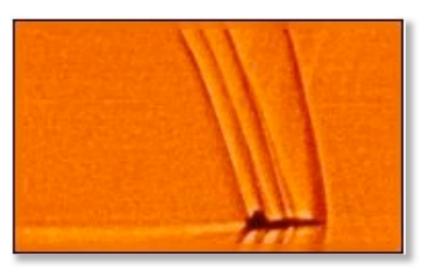
Sound & Doppler effect

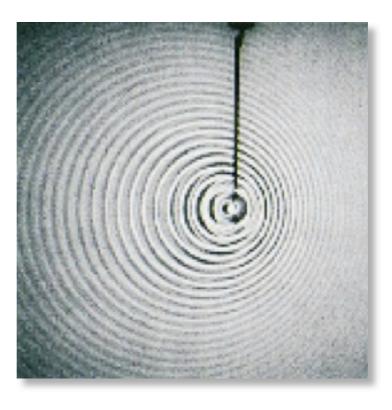


Doppler effect

Shock waves and

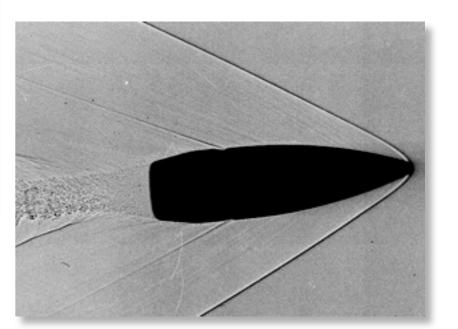
Sonic booms















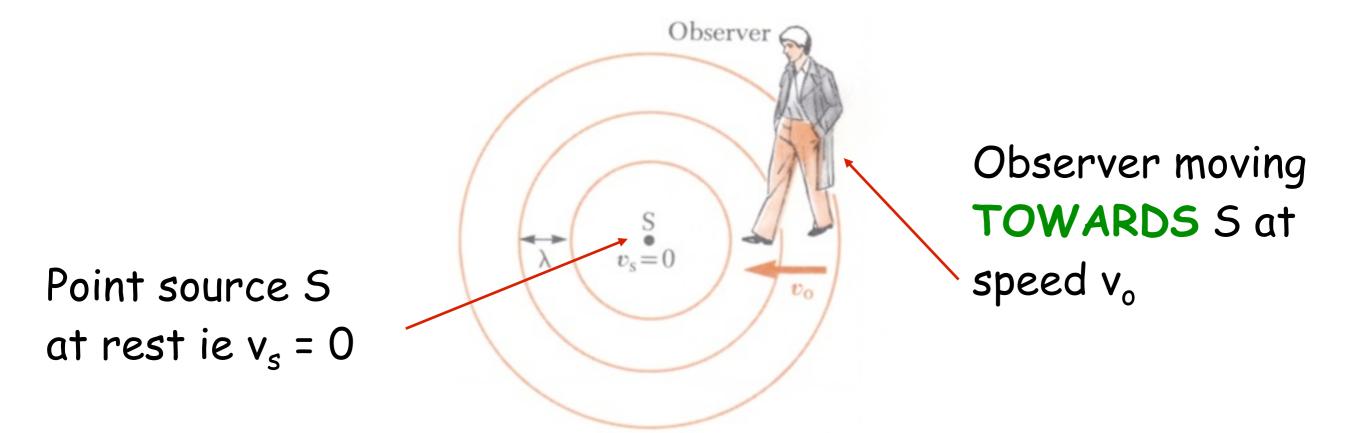
The Doppler effect is experienced whenever there is a relative motion between source and observer.

When the source and observer are moving towards each other the frequency heard by the observer is higher than the frequency of the source.

When the source and observer move away from each other the observer hears a frequency which is lower than the source frequency.

Although the Doppler effect is most commonly experienced with sound waves it is a phenomenon common to all harmonic waves.



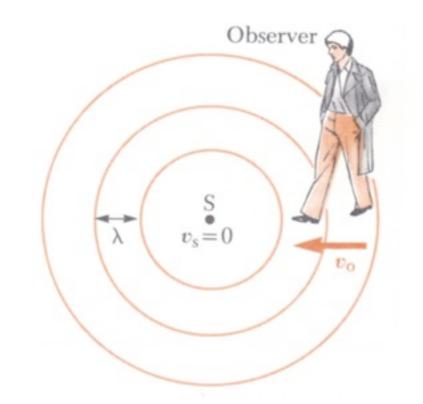


We are dealing with relative speeds ... "at rest" = at rest with respect to the air. We assume the air is stationary.

frequency of source = f velocity of sound = v frequency observed = f' wavelength of sound = λ







If observer O was stationary he would detect f wavefronts per second

When $O \longrightarrow S$ O moves a distance $v_o t$ in t seconds During this time O detects an additional $\frac{v_o t}{\lambda}$ wavefronts ie: an additional $\frac{v_o}{\lambda}$ wavefronts / second

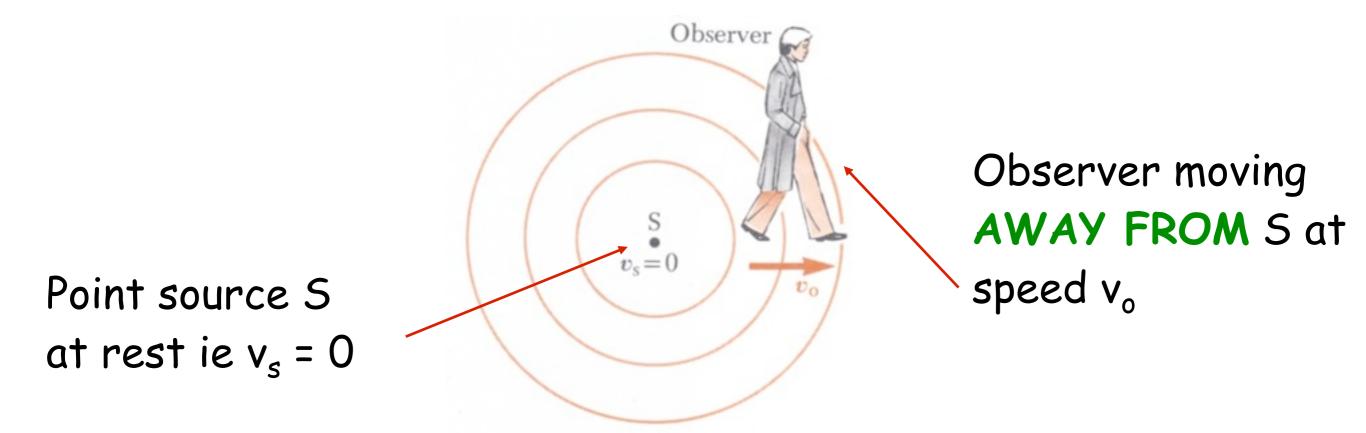




As more wavefronts are heard per second the frequency heard by the observer is increased.

$$f' = f + \Delta f = f + \frac{v_o}{\lambda}$$

but $v = f\lambda$ or $\frac{1}{\lambda} = \frac{f}{v}$
 $\therefore \frac{v_o}{\lambda} = \frac{v_o}{v} f$
 $\therefore f' = f + \frac{v_o}{v} f$ (v + v_o) = speed of waves
 $f' = f \left(\frac{v + v_o}{v} \right)$ relative to O



O now detects fewer wavefronts /second and therefore the frequency is lowered.

The speed of the wave relative to O is $(v - v_o)$

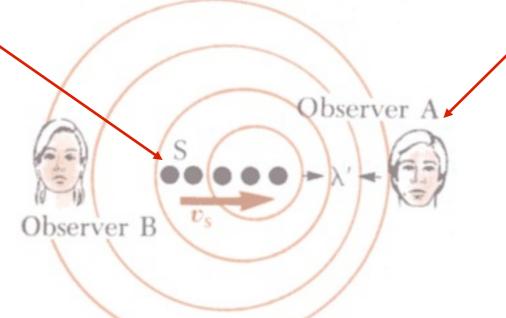
$$\therefore f' = f\left(\frac{v - v_o}{v}\right)$$



 V_{s}

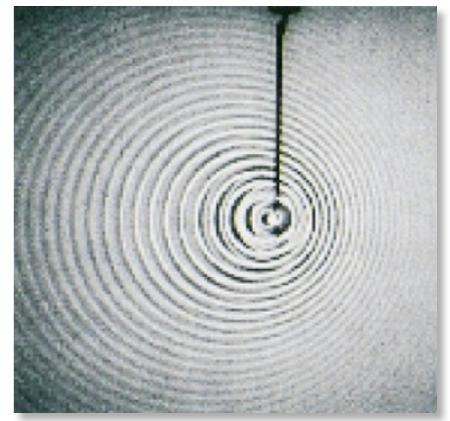


Point source S moving with speed



Source is moving towards O_A at speed v_s .

Wavefronts are closer together as a result of the motion of the source.







The observed wavelength λ' is shorter than the original wavelength $\lambda.$

During one cycle (which lasts for period T) the source moves a distance v_sT (= v_s/f)

In one cycle the wavelength is shortened by v_s/f

$$\lambda' = \lambda - \Delta \lambda = \lambda - \frac{v_{s}}{f}$$

but $\lambda = \frac{v}{f}$ and $\lambda' = \frac{v}{f'}$
 $\therefore f' = \left(\frac{v}{\lambda - v_{s}/f}\right)^{\frac{1}{f}}$





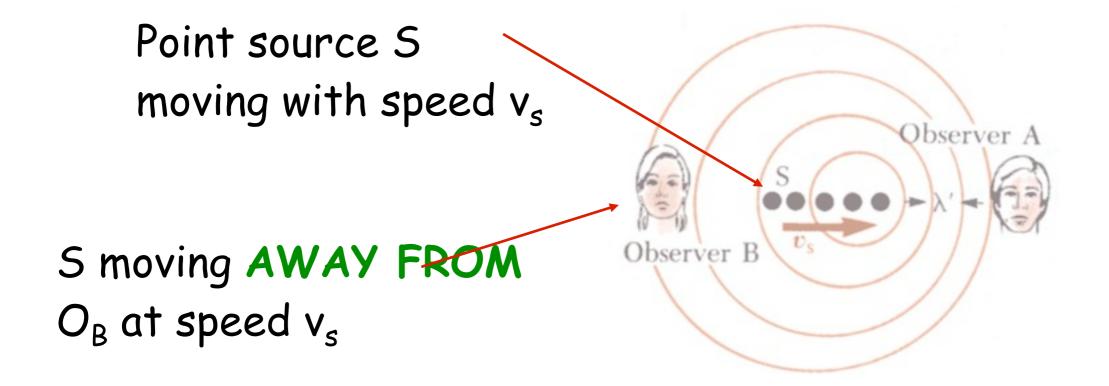
$$f' = \left(\frac{v}{\lambda - v_s/f}\right)^{\frac{1}{j}}$$
$$= \left(\frac{v}{v/f - v_s/f}\right)^{\frac{1}{j}}$$
$$f' = f\left(\frac{v}{v - v_s}\right)^{\frac{1}{j}}$$

ie: the observed frequency is increased when the source moves towards the observer.

Note - the equation breaks down when $v_s \sim v$. We will discuss this situation later.







Source is moving away from O_B at speed v_s .

Wavefronts are further apart, λ is greater and O_{B} hears a decreased frequency given by

$$f' = f\left(\frac{v}{v + v_s}\right) \frac{1}{\dot{f}}$$





Frequency heard when observer is in motion

$$f' = f\left(\frac{v \pm v_o}{v}\right) + O \text{ towards S} \\ - O \text{ away from S}$$

Frequency heard when source is in motion

$$f' = f\left(\frac{v}{v \mp v_s}\right) + S \text{ towards } O$$

+ S away from O

Frequency heard when observer is in motion

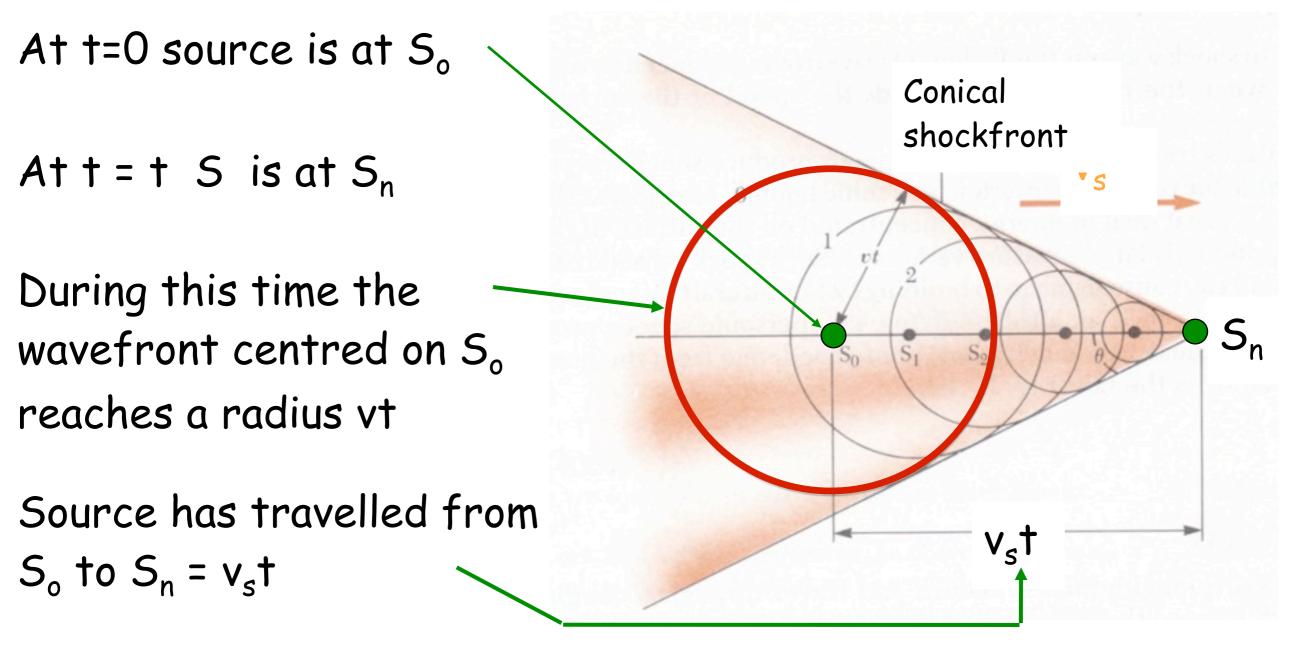
$$f' = f\left(\frac{v \pm v_o}{v \mp v_s}\right) = \frac{1}{2}$$

$$lower signs = away from$$





Go back to situation where source is moving with velocity $v_{\rm s}$ which exceeds wave velocity





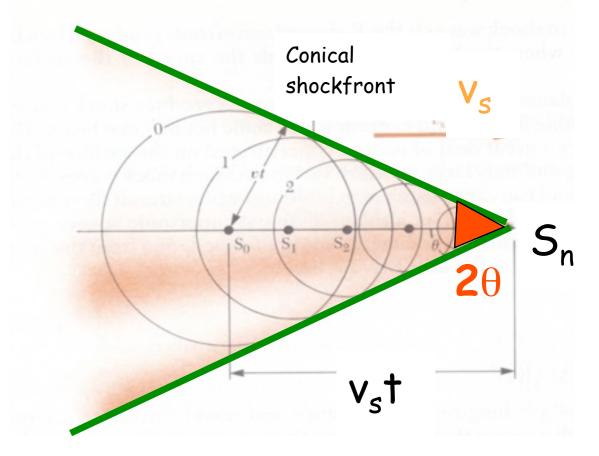


At t=t S is at S_n and waves are just about to be produced here

The line drawn from S_n to the wavefront centred on S_o is tangential to all wavefronts generated at intermediate times

The envelope of these waves is a cone whose apex half angle θ is given by

$$\sin\theta = \frac{vt}{v_s t} = \frac{v}{v_s}$$







The ratio $\frac{v}{v_s}$ is known as the MACH number.

The conical wavefront produced when $v_s > v$ (supersonic speeds) is known as a shock wave.

An aeroplane travelling at supersonic speeds will produce shockwaves.

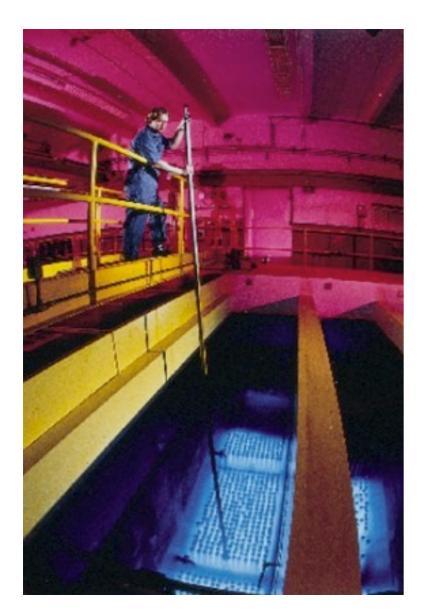
In this photo the cloud is formed by the adiabatic cooling of the shock wave to the dew point.







$$\sin\theta = \frac{vt}{v_s t} = \frac{v}{v_s}$$



This equation can also be applied to a form of electromagnetic radiation called Cerenkov radiation.

A charged particle moves in a medium with speed v that is greater than the speed of light in that medium

The blue glow surrounding the fuel elements in a nuclear reactor is an example of Cerenkov radiation