# Integral pulse height distribution (curva di conteggio)



# at determines Resolution?

## gnal Variance >> Baseline Variance



⇒ Electronic (baseline) noise not important

Examples: 
High-gain proportional chambers

Scintillation Counters with High-Gain PMTs

e.g. 1 MeV  $\gamma$ -rays absorbed by NaI(Tl) crystal Number of photoelectrons:  $N_{pe} \approx 8 \cdot 10^4 \, [\text{MeV}^{-1}] \times E_{\gamma} \times QE \approx 2.4 \cdot 10^4$ Variance typically:  $\sigma_{pe} = N_{pe}^{-1/2} \approx 160 \text{ and } \sigma_{pe} / N_{pe} \approx 5 - 8\%$ Signal at PMT anode (assume Gain=  $10^4$ ):  $Q_{sig} = G_{PMT} N_{pe} \approx 2.4 \cdot 10^8 \text{ el anc}$   $\sigma_{sig} = G_{PMT} \sigma_{pe} \approx 1.2 \cdot 10^7 \text{ el}$ whereas electronic noise easily <  $10^4 \text{ el}$ 

### nal Variance << Baseline Variance



⇒ Electronic (baseline) noise critical for resolution

Examples: • Gaseous ionization chambers (no internal gain)

Semiconductor detectors

e.g. in Si : Number of electron-hole pairs  $N_{ep} = \frac{E_{dep}}{3.6 \text{ eV}}$ Variance  $\sigma_{ep} = \sqrt{F \cdot N_{ep}}$  (where F= Fano factor  $\approx 0.1$ ) For 50 keV photons:  $\sigma_{ep} \approx 40 \text{ el} \Rightarrow \sigma_{ep} / N_{ep} = 7.5^{\circ}10^{-4}$ 

Obtainable noise levels are 10 to 1000 el.

Baseline fluctuations can have many origins ...

pickup of external interference

artifacts due to imperfect electronics

... etc.,

but the (practical) fundamental limit is electronic noise.

#### 1.10 Detection limits and resolution

In addition to signal fluctuations originating in the sensor, the minimum detection limit and energy resolution are subject to fluctuations introduced by the electronics. The gain can be controlled very precisely, but electronic noise introduces baseline fluctuations, which are superimposed on the signal and alter the peak amplitude. Figure 1.24 (left) shows a typical noise waveform. Both the amplitude and time distributions are random.

When superimposed on a signal, the noise alters both the amplitude and time dependence. Figure 1.24 (right) shows the noise waveform superimposed on a small signal. As can be seen, the noise level determines the minimum signal whose presence can be discerned.

#### DETECTOR SYSTEMS OVERVIEW



FIG. 1.24. Waveforms of random noise (left) and signal + noise (right), where the peak signal is equal to the rms noise level (S/N = 1). The noiseless signal is shown for comparison.

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# Spectroscopy:

- Experiments to determine the decay products of excited states and their interaction
- -The excitation energies of a system increase as size decrease. To produce these excited states high energy particles are needed.

Absolute energies may vary a lot But resolution usually still a critical issue



solution and Electronic Noise 10<sup>6</sup> NaI(TI) SCINTILLATOR 105 olution: the ability to iguish signal levels 104 hy? COUNTS cognize structure in amplitude spectra Comparison between NaI(TI) 10<sup>3</sup> and Ge detectors Ge DETECTOR 10<sup>2</sup> 10 . Philippot, IEEE Trans. Nucl. Sci. NS-17/3 (1970) 446 500 1500 2 0 1000 DOV /LaVA

b) Improve sensitivity

Signal to background ratio improves with better resolution

(signal counts in fewer bins compete with fewer background counts)



G.A. Armantrout et al., IEEE Trans. Nucl. Sci. NS-19/1 (1972) 107

li di Frascati of INFN, wh text year. The experiment, v ram, has some special f ypernuclear physics experivill operate at a  $e^+e^-$  collide Vhat follows will describe it

#### tructure of $\Lambda$ -hypernuc

A-hypernucleus  ${}^{A}_{\Lambda}Z$  is a be rons and a  $\Lambda$  hyperon. The nade by the (A-1) nucleons a nucleus  ${}^{(A-1)}Z$  and the /he  $\Lambda$  hyperon, carrying the istinguishable baryon an nposed by the Pauli princip im states already filled up  $\vee$ hyperon, embedded in a splore nuclear structure.

The binding energy  $B_{\Lambda}$  of tis ground state is defined

#### $B_{\Lambda} = M$

here  $M_{core}$  is the mass (in 1 is mass of the  $\Lambda$  particle an Z, experimentally measure ope of about 1 MeV/(unit of the heavy hypernuclei. The which the  $\Lambda$  particle is conjual to the nuclear radius and the 55 MeV typical value of This is consistent with a  $\Lambda$ icleon-nucleon one. Indeect teraction, the zero isospir.

by by external systems of the  $\pi$  or the  $\rho$  with a nucleon and deterines the lack of strong tensor components in the interaction. In relative weakness of the  $\Lambda$ -nucleon interaction entails that the ell structure is not disrupted by the insertion of the  $\Lambda$  in the inclus and the lack of Pauli effects allows all the nuclear single rticle states to be populated by the  $\Lambda$ . In Figure 2, the so called *egrè table*" of the hypernuclei shows the 35 hypernuclei known present.

Experiments of hypernucleus production by "strangeness change" and "associated production" processes can produce pernuclei in which the  $\Lambda$  populates different single particle tes. The latter technique is particularly suitable for populating v lying  $\Lambda$  states, thanks to the high recoil momentum transred to the  $\Lambda$  particle in the reaction.

A beautiful representation of this process is given in Figure 3, iere the excitation spectrum of  $\overset{8}{}\overset{8}{}^{P}Y$ , obtained by the "associated pduction" reaction  $\overset{89}{}^{P}(\pi^{+}, K^{+})\overset{8}{}\overset{8}{}^{P}Y$  at the KEK laboratory in van, is shown. The spectrum demonstrates how, starting from leutron in the  $g_{92}$  state, it is possible to accommodate a  $\Lambda$  particle the hypernuclear states f, d, p and even in the ground state s.

These measurements constitute the spectacular confirmation, a textbook level, of the validity of the independent particle odel or shell model of the nucleus. In non-strange nuclei, the servation of single particle states is only possible for the states the most external nucleon orbits. In fact, due to the Pauli prinle and pairing interactions, deeply bound nucleon single rticle states are so fragmented as to be essentially unobservable. e present experimental data on hypernuclear binding energies d detailed spectroscopic features are limited in quantity and

tion of the YN interaction.

Figure 4 reports a recent measurement of the splitting of the  $5/2^+-3/2^+$  doublet in  ${}^{ABe}_{A}$  by the BNL-AGS E930 experiment [4], measuring  $\gamma$  rays emitted in the nuclear transitions with the new germanium detector array Hyperball. This new technique allowed the energy resolution on low lying hypernuclear levels to be improved from a few MeV to a few keV, even if the count rate resulting is still quite low, ~200  $\gamma$ 's per month of data taking. The spacing of the two levels was measured to be  $31 \pm 2$  keV, incompatible with the prediction of the meson exchange models.

state nucleon in all but the l mesonic decay modes of tl through the weak interac  $(\Lambda + n \rightarrow n + n + 176 \text{ MeV})$ process is possible only in hyr of  $\Lambda^{3}$ , stable against the mes which is available inside a hyr

The study of the non-mes importance, since it provide four fermion, strangeness cha





precise data on the free YN interaction, which are very difficult to



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# **Response function**



There are actually 2 monoenergetic radiation Sources...





Acceptance effect on proton momentum distribution



Original distribution

**Reconstructed distribution** 



# Ap acceptance



# Dead time measurements



# TOF difference needed to mass discriminate particles



# How a shower looks like



Electron shower in lead. 7500 gauss in cloud chamber. CALTECH



Electron shower in lead. Cloud chamber. W.B. Fretter, UCLA

#### Electron showers

Hadron showers







# EM showers: transverse profile

# Transverse shower profile

- Multiple scattering make electrons move away from shower axis
- Photons with energies in the region of minimal absorption can travel far away from shower axis

Molière radius sets transverse shower size, it gives the average lateral deflection of critical energy electrons after traversing  $1{\rm X}_{\rm 0}$ 

$$R_{M} = \frac{21 MeV}{E_{C}} X_{0} \qquad R_{M} \propto \frac{X_{0}}{E_{C}} \propto \frac{A}{Z} (Z >> 1)$$

75%  $E_0$  within  $1R_M$ , 95% within  $2R_M$ , 99% within  $3.5R_M$ 

# charged particles only



# all particles



## Electromagnetic calorimeters.



Figure 27.18: An EGS4 simulation of a 30 GeV electron-induced cascade in iron. The histogram shows fractional energy deposition per radiation length, and the curve is a gamma-function fit to the distribution. Circles indicate the number of electrons with total energy greater than 1.5 MeV crossing planes at  $X_0/2$  intervals (scale on right) and the squares the number of photons with  $E \ge 1.5$  MeV crossing the planes (scaled down to have same area as the electron distribution).



Fig. 7.20. Schematic representation of an electromagnetic cascade. The wavy lines are photons and the solid lines electrons or positrons.





**Figure 11.2** Shower profiles in lead. The number of electrons should be multiplied by a normalization factor of 0.79. (D. Müller, Phys. Rev. D 5: 2677, 1972.)

**Figure 11.3** Longitudinal development of electromagnetic showers different materials. Right scale shows radii for 90% shower containmer (C. Fabjan and T. Ludlam, adapted with permission from the Annu Review of Nuclear and Particle Science, Vol. 32, © 1982 by Annu Reviews, Inc.)











# Electromagnetic showers

# Energy resolution: Limitations







Shower simulation



Figura 6.16: Distribuzione dei punti rappresentativi sullo scatterplot (CER vs momento) delle particelle cariche rivelate dallo spettrometro CHAOS, ottenuta dopo la procedura di identificazione. Nel grafico sono mostrate le particelle con tutti i gradi di identificazione (PID)

# Hadronic showers

#### Hadronic shower development

- General comment: Complexity of of hadronic and nuclear processes produce multitude of effects that determine the functioning and performance of hadron calorimeters
  - Many channels compete in the development of hadronic showers
  - Larger variations in the deposited and visible energy
  - More complicated to optimize
- Sizeable electromagnetic (e) besides hadronic (h) shower contribution mainly from π<sup>0</sup> decay (1/3 of pions)
- Invisible energy due to delayed emitted photons in nuclear reactions, soft neutrons and binding energy
- Visible energy smaller for hadronic (h) than for electromagnetic (e) showers: Ratio of response e/h > 1
- Larger intrinsic fluctuations for hadronic than electromagnetic showers
- Improvements: Increase visible energy to get e/h=1: Compensation (Compensation for the loss of invisible energy)!
- Discussed instr. effects for e showers also hold for h showers

Step 1: Production of energetic hadrons with a mean free path given by the nuclear interaction length:



dE/dx nuclear cascade dE/dx nuclear cascade

Step 2: Hadronic collisions with material nuclei (significant part of the primary part of primary energy is consumed in nuclear processes):

Evaporation

Evaporation followed by evaporation



# Processes contributing to the energy deposition in hadronic cal.

Process	Percent of total
Secondary proton ionization	31.6
Electromagnetic cascade $(\pi^0)$	21.0
Nuclear binding energy plus neutrino energy	20.6
Secondary $\pi^{\pm}$ ionization	8.2
Neutrons with $E > 10 \text{ MeV}$	4.9
Neutrons with $E < 10 \text{ MeV}$	3.9
Residual nuclear excitation energy	3.7
Z > 1 ionization	2.4
Primary proton ionization	2.3
Other	1.4

Table 11.2. Average fractional energy deposition for a 10-GeV proton in an iron/liquid argon calorimeter

Source: T. Gabriel and W. Schmidt, Oak Ridge National Laboratory report, ORNL/TM-5105, 1975.

Hadronic Calorimeters (are [very] difficult to model)

20 GeV  $\pi$  in copper (simulation)



J.P. Wellisch



Hadronic Showers (m, n, p, ...) inelastic hadron interactions Propagation : very LARGE  $\rightarrow$  multi particle production Nuclear disintegration

Neutrino

Weak interactions secondaries : mostly hadrons

# Nuclear Interaction Length $\lambda_i$

is the average distance a high-energy hadron has to travel inside a medium before a nuclear interaction occurs.

Probability not to have interacted after a path  $z P = e^{-z/\lambda_i}$ 



# Longitudinal distribution of energy in hadron calorimeter.

