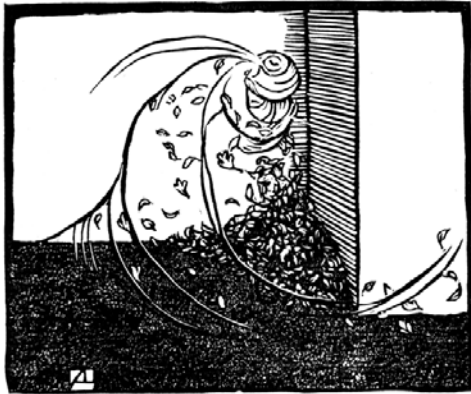


# Vortices occur on small scales

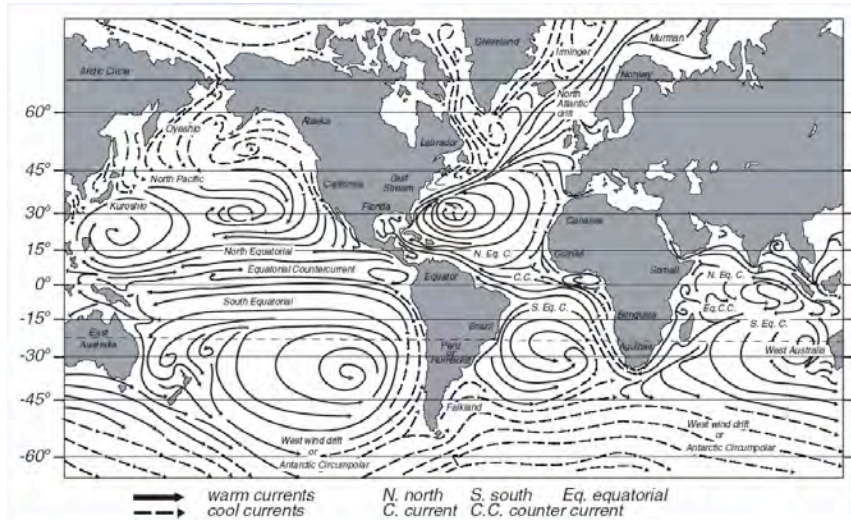
Leaves whirling in the wind



Von Karman street behind bridge pillars

# Vortices occur in the oceans

Large scale wind-driven circulation in the upper ocean



## Vortices occur in the atmosphere

Satellite image of tropical cyclone Irma (2017)



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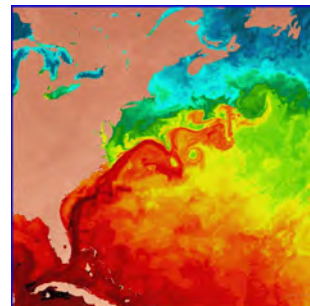
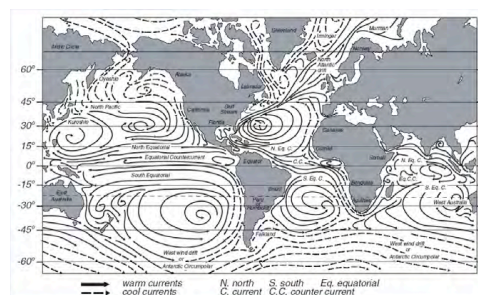
Utrecht University



## Vortices are omni-present in geophysical flows

Because of:

- (two-dimensional) turbulence
- Rotation of the Earth
  - In particular its latitudinal variation



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## Definition of vorticity

3D:

$$\vec{\zeta} \equiv \nabla \times \vec{u}$$

2D:

$$\zeta = \vec{e}_3 \cdot \nabla \times \vec{u} = \frac{\partial v}{\partial x} - \frac{\partial u}{\partial y}$$

- In 3D: curl of velocity field.
- In 2D: vertical component only.

### Interpretation:

- Local measure of rotation: spin of infinitesimal fluid parcel around its own axis.
- positive: counterclockwise, negative: clockwise, cf. cork screw rule
- Local density of circulation per unit infinitesimal area:  
Stokes theorem (Green in 2D):

$$\oint_{\partial A} \vec{u} \cdot d\vec{l} = \iint_A \zeta \, dx \, dy$$

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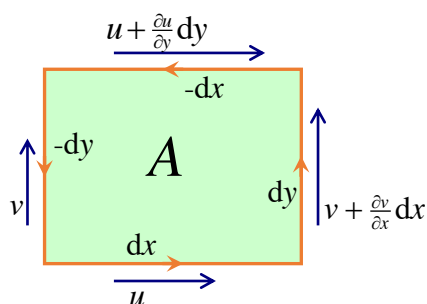
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## Local density of circulation per unit infinitesimal area

Circulation along the red path around infinitesimal area  $A$ :



$$\begin{aligned} \oint_{\partial A} \begin{pmatrix} u \\ v \end{pmatrix} \cdot d\vec{l} &= u \, dx + \\ &+ (v + \frac{\partial v}{\partial x} dx) \, dy \\ &- (u + \frac{\partial u}{\partial y} dy) \, dx - v \, dy = \\ &= (\frac{\partial v}{\partial x} - \frac{\partial u}{\partial y}) \, dx \, dy = \iint_A \zeta \, dx \, dy \end{aligned}$$

Dividing by the area  $A = dx \, dy$ , one finds:  
vorticity is the local circulation per unit area.

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## Vorticity in the geophysical context

On the Earth, even fluids at rest have vorticity:

**Planetary vorticity** due to rigid body rotation of the Earth:

$$\zeta_p = 2\Omega \sin(\theta) = f$$

Fluid motion is measured relative to the Earth's rotation, so

**Relative vorticity** is calculated from velocities relative to the Earth:

$$\zeta_r = \zeta = \frac{\partial v}{\partial x} - \frac{\partial u}{\partial y}$$

In an inertial frame, fluid motion is both with and relative to the Earth,

**Absolute vorticity** is the sum of planetary and relative vorticity:

$$\zeta_a = \zeta + f$$

**Potential vorticity** is a conserved quantity following fluid parcels:

$$\Pi = \frac{\zeta + f}{h}$$

It will be derived later.

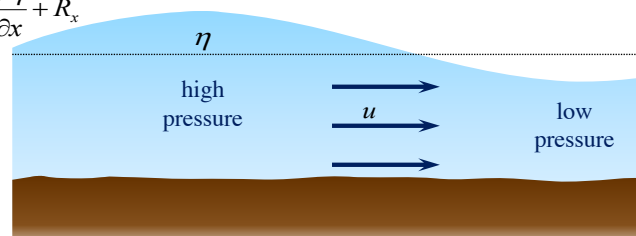
## Depth-averaged 2D shallow water equations

$$\text{Momentum equation: } \rho \left( \frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + v \frac{\partial u}{\partial y} + w \frac{\partial u}{\partial z} - f v \right) = -\frac{\partial p}{\partial x} + r_x$$

Main assumption: vertical scale  $\ll$  horizontal scales of motion

- Velocities (almost) independent of  $z$
- Hydrostatic balance:  $p = \rho g (\eta - z)$

$$\text{Depth-averaged: } \frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + v \frac{\partial u}{\partial y} - f v = -g \frac{\partial \eta}{\partial x} + R_x$$



# Derivation of the vorticity equation

(starting from the depth-averaged 2D shallow water equations)

$$\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + v \frac{\partial u}{\partial y} - f v = -g \frac{\partial \eta}{\partial x} + R_x$$

$$\frac{\partial v}{\partial t} + u \frac{\partial v}{\partial x} + v \frac{\partial v}{\partial y} + f u = -g \frac{\partial \eta}{\partial y} + R_y$$

Now differentiate v-momentum equation with respect to x, u-momentum equation with respect to y and subtract:

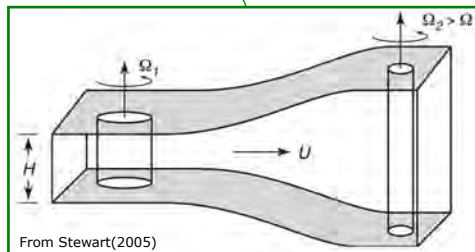
Remember:  $\zeta = \frac{\partial v}{\partial x} - \frac{\partial u}{\partial y}$

# Vorticity equation

$$\frac{D\xi}{Dt} + (\xi + f) \left( \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} \right) + \beta v = \frac{\partial R_y}{\partial x} - \frac{\partial R_x}{\partial y}$$

Local derivative + advection

Stretching or ballerina effect



# Vorticity equation

$$\frac{D\xi}{Dt} + (\xi + f) \left( \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} \right) + \beta v = \frac{\partial R_y}{\partial x} - \frac{\partial R_x}{\partial y}$$

Local derivative + advection

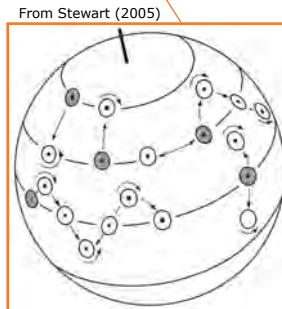
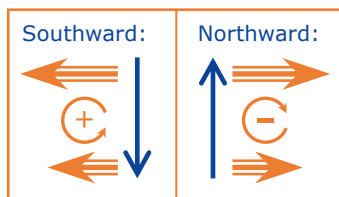
Stretching or ballerina effect

Beta-effect

Frictional torque

Important contributions in depth-integrated equations due to wind stress and bottom drag:

$$\frac{\partial}{\partial x} \left( \frac{\tau_y^w - \tau_y^b}{\rho} \right) - \frac{\partial}{\partial y} \left( \frac{\tau_x^w - \tau_x^b}{\rho} \right)$$



# Intermezzo: Depth-integrated continuity equation

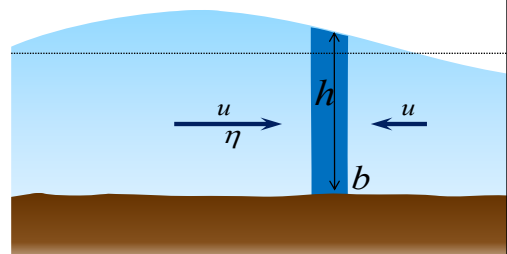
Continuity equation (incompressible flow):  $\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} + \frac{\partial w}{\partial z} = 0$

Integrate over depth from bottom ( $z = b$ ) to surface ( $z = \eta$ ):

$$0 = h \left( \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} \right) + w|_{z=\eta} - w|_{z=b} = h \left( \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} \right) + \frac{D\eta}{Dt} - \frac{Db}{Dt}$$

(remember:  $u, v$  do not depend on  $z$  in shallow water)

Conclusion:  $\frac{Dh}{Dt} + h \left( \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} \right) = 0$



## Derivation of potential vorticity equation

$$\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} = -\frac{1}{h} \frac{Dh}{Dt}$$

to eliminate the convergence term in the vorticity equation:

$$\frac{D\xi}{Dt} - \frac{\xi + f}{h} \frac{Dh}{Dt} + \frac{Df}{Dt} = \frac{\partial R_y}{\partial x} - \frac{\partial R_x}{\partial y}$$

$$\frac{D}{Dt} \left( \frac{\xi + f}{h} \right) = \frac{1}{h} \left( \frac{\partial R_y}{\partial x} - \frac{\partial R_x}{\partial y} \right)$$

So in the absence of friction, potential vorticity is conserved along trajectories!

## Vorticity-driven flows

Introductory:

- a) Zonal flow
- b) Topographic steering
- c) Interaction between vortices

More advanced:

- d) Moving vortices due to  $\beta$ -effect
- e) Large-scale wind driven circulation

## a) Zonal flow in the atmosphere

Remind: Potential vorticity =  $\Pi = \frac{\zeta + f}{h}$

Conserving potential vorticity when:

- velocities are relatively small ( $\zeta \ll f$ ),
- depth  $h$  does not change much,

implies that  $f$  does not change much, i.e. flow is zonal...

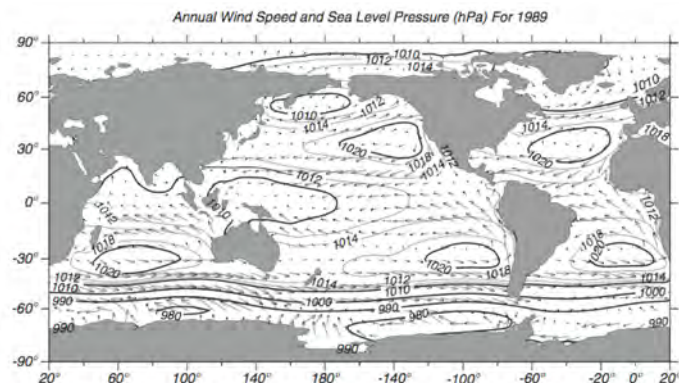
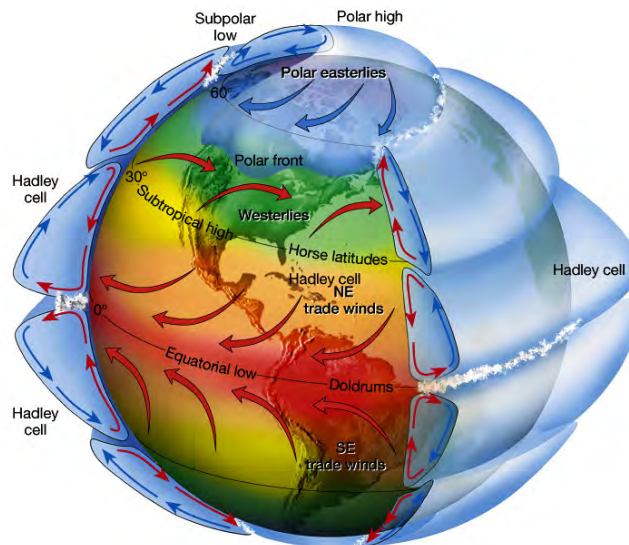


Figure 4.2 Map of mean annual wind velocity calculated from Trenberth et al. (1990) and sea-level pressure for 1989 from the NASA Goddard Space Flight Center's Data Assimilation Office (Schubert et al. 1993). The winds near 140°W in the equatorial Pacific are about 8 m/s.

## a) Zonal flow in the atmosphere



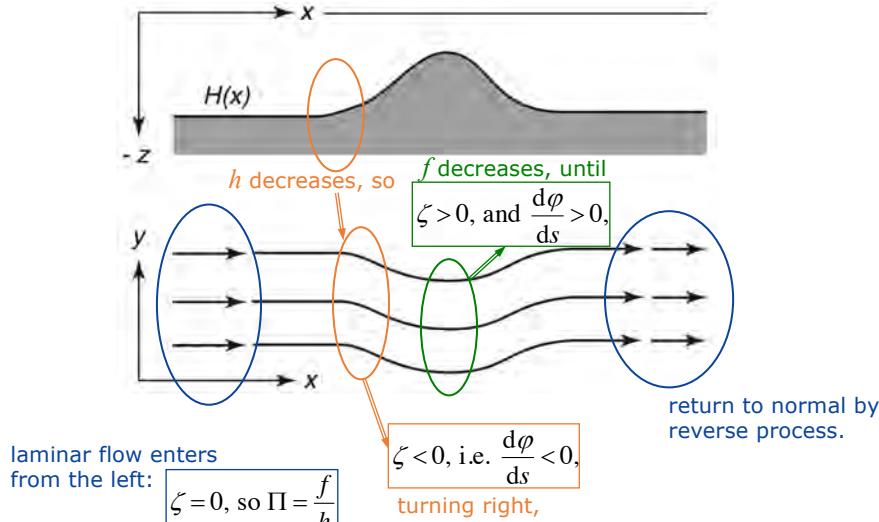
Warm air rises along equator, cold air sinks air at poles, inducing meridional flow; Angular momentum conservation requires huge westerly winds  
 => instabilities develop and flow will be zonal.



## b) Topographic steering

A northern hemisphere example.

Potential vorticity:  $\Pi = \frac{\zeta + f}{h}$



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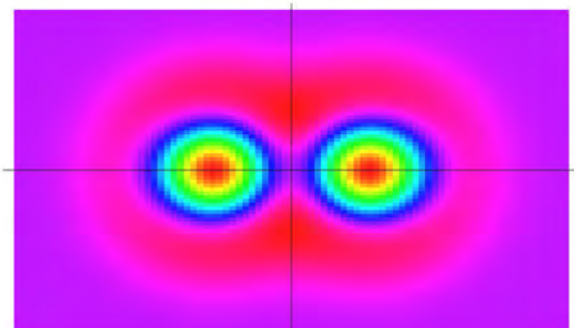


## c) Interacting vortices

Suppose:

- No friction,
- constant depth (i.e. no stretching),
- Length scale small compared to Earth' radius (i.e. no beta).

And we have a vorticity field like this (two positive vortices):



First question: what is the corresponding velocity field?

Vorticity dynamics:  $0 = \frac{D\zeta}{Dt} = \frac{\partial \zeta}{\partial t} + u \frac{\partial \zeta}{\partial x} + v \frac{\partial \zeta}{\partial y}$

Vorticity is just advected with the flow!

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