

forming one class are the planetary waves, in which the time evolution is prompted by the weak but important *planetary effect*.

As we may recall from Section 2-6, on a spherical earth (or planet or star, in general), the Coriolis parameter,  $f$ , is proportional to the rotation rate,  $\Omega$ , times the sine of the latitude,  $\varphi$ :

$$f = 2\Omega \sin \varphi.$$

Large wave formations such as alternating cyclones and anticyclones contributing to our daily weather and, to a lesser extent, the Gulf Stream meanders span several degrees of latitude; for them, it is necessary to consider the meridional change in the Coriolis parameter. If the coordinate  $y$  is oriented northward and is measured from a reference latitude  $\varphi_0$  (say, a latitude somewhere in the middle of the wave under consideration), then  $\varphi = \varphi_0 + y/a$ , where  $a$  is the earth's radius (6371 km). Considering  $y/a$  as a small departure, the Coriolis parameter can be expanded in a Taylor series:

$$f = 2\Omega \sin \varphi_0 + 2\Omega \frac{y}{a} \cos \varphi_0 + \dots \quad (6-15)$$

Retaining only the first two terms, we write in traditional notation

$$f = f_0 + \beta_0 y, \quad (6-16)$$

where  $f_0 = 2\Omega \sin \varphi_0$  is the reference Coriolis parameter and  $\beta_0 = 2(\Omega/a) \cos \varphi_0$  is the *beta parameter*. Typical midlatitude values on Earth are  $f_0 = 8 \times 10^{-5} \cdot \text{s}^{-1}$  and  $\beta_0 = 2 \times 10^{-11} \text{ m}^{-1} \cdot \text{s}^{-1}$ . The Cartesian framework where the beta term is not retained is called the *f-plane*, and that where it is retained is called the *beta plane*. The next step in order of accuracy is to retain the full spherical geometry (which we will avoid throughout this book). Rigorous justifications of the beta-plane approximation can be found in Veronis (1963, 1981), Pedlosky (1987), and Verkley (1990).

Note that the beta-plane representation is validated at mid latitudes only if the  $\beta_0 y$  term is small compared to the leading  $f_0$  term. In terms of the motion's meridional length scale  $L$ , this implies

$$\beta = \frac{\beta_0 L}{f_0} \ll 1, \quad (6-17)$$

where the dimensionless ratio can be called the *planetary number*.

The governing equations, having become

$$\frac{\partial u}{\partial t} - (f_0 + \beta_0 y)v = -g \frac{\partial \eta}{\partial x}, \quad (6-18a)$$

$$\frac{\partial v}{\partial t} + (f_0 + \beta_0 y)u = -g \frac{\partial \eta}{\partial y}, \quad (6-18b)$$

$$\frac{\partial \eta}{\partial t} + H \left( \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} \right) = 0, \quad (6-18c)$$