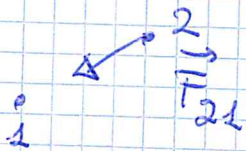


THEORY OF POTENTIAL  
 PLEASE, REFER TO BINNEY AND TREHARNE  
BOOK

①

372  
 NUM

You know that



$$\vec{F}_{21} = - \frac{G m_1 m_2}{r_{12}^2} \hat{r}_{12}$$

$\vec{F}$  is a conservative force  
 all "central" forces

$$W_{\text{work}} = \int_A^B \vec{F} \cdot d\vec{s}$$

$$\dots \hat{r} \cdot d\vec{s}$$

$$\dots \hat{r} d\hat{r} \rightarrow \int_{r_A}^{r_B}$$

$W = \text{potential energy}$

$$\Delta W = - W_{\text{work}} = - \int \vec{F} \cdot d\vec{s}$$

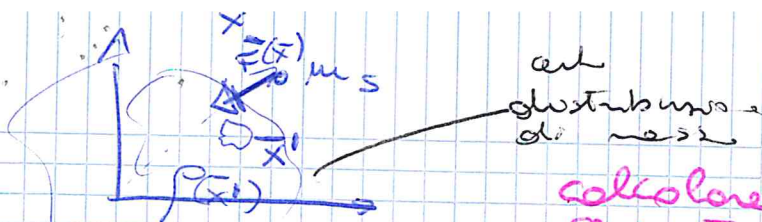
$$W = - \frac{G m_1 m_2}{r_{12}} + \text{const}$$

$\rightarrow 0 \quad r \rightarrow \infty$

1 2 3 ~ particles

binding energy

$$W = - \frac{G m_1 m_2}{r_{12}} - \frac{G m_2 m_3}{r_{23}} - \frac{G m_1 m_3}{r_{13}}$$



calcolare la forza generata da  $\rho(x')$  su  $m_s$

$$\delta \vec{F}(\vec{x}) = G m_s \frac{\vec{x}' - \vec{x}}{|\vec{x}' - \vec{x}|^3} \delta m(\vec{x}') = G m_s \frac{\vec{x}' - \vec{x}}{|\vec{x}' - \vec{x}|^3} \rho(\vec{x}') d^3(x')$$

the overall force  $\leftarrow$  integrating in the whole space

$$\vec{F}(\vec{x}) = \int \delta \vec{F}(\vec{x})$$

$$\vec{F}(\vec{x}) = m_s \vec{g}(\vec{x})$$

gravitational field

$$\vec{g}(\vec{x}) \equiv G \int d^3 \vec{x}' \frac{(\vec{x}' - \vec{x})}{|\vec{x}' - \vec{x}|^3} \rho(\vec{x}')$$

2.2

We define the gravitational potential

$$\phi(\vec{x}) \equiv -G \int d^3 \vec{x}' \frac{\rho(\vec{x}')}{|\vec{x}' - \vec{x}|}$$

2.3

$$\vec{\nabla} \phi(\vec{x}) = -G \vec{\nabla}_x \int d^3 \vec{x}' \frac{\rho(\vec{x}')}{|\vec{x}' - \vec{x}|} =$$

the  $\int$  does not depend on  $\vec{x}$  ...

$$= -G \int d^3 \vec{x}' \rho(\vec{x}') \vec{\nabla}_x \frac{1}{|\vec{x}' - \vec{x}|}$$

but  $\vec{\nabla}_x \left( \frac{1}{|\vec{x}' - \vec{x}|} \right) = \frac{\vec{x}' - \vec{x}}{|\vec{x}' - \vec{x}|^3}$

See other papers \*

su "credit to Romano Memetti"

$$\Rightarrow = -G \int d^3 \vec{x}' \rho(\vec{x}') \frac{\vec{x}' - \vec{x}}{|\vec{x}' - \vec{x}|^3} \Rightarrow$$

$$\vec{g}(\vec{x})$$

$$\vec{g}(\vec{x}) = -\vec{\nabla} \phi$$

2.5

the potential is very useful ...

a scalar field is easier to visualize than a vector field

Now

completion

$$\bar{\nabla}(\bar{g}(\bar{x})) = -\bar{\nabla}\bar{\nabla}\phi = -\bar{\nabla}^2\phi$$

divergence

$$\bar{\nabla}(\bar{g}(\bar{x})) = G \int d^3\bar{x}' \bar{\nabla}_x \left( \frac{\bar{x}' - \bar{x}}{|\bar{x}' - \bar{x}|^3} \right) \rho(\bar{x}') \quad (2.6)$$

We know that

$$\bar{\nabla}_x \left( \frac{\bar{x}' - \bar{x}}{|\bar{x}' - \bar{x}|^3} \right) = -\frac{3}{|\bar{x}' - \bar{x}|^3} + \frac{3(\bar{x}' - \bar{x})(\bar{x}' - \bar{x})}{|\bar{x}' - \bar{x}|^5} \quad (2.7)$$

see the papers

then  $\bar{x}' - \bar{x} \neq 0$

$$\frac{3(\bar{x}' - \bar{x})(\bar{x}' - \bar{x})}{|\bar{x}' - \bar{x}|^5}$$

i.e.

$$\bar{\nabla}_x \left( \frac{\bar{x}' - \bar{x}}{|\bar{x}' - \bar{x}|^3} \right) = 0 \quad (\bar{x}' \neq \bar{x}) \quad (2.8)$$

So the contribution to the integral (2.6) comes from the point  $\bar{x}' = \bar{x}$  around this point  $\rho(\bar{x}') \sim \text{const.}$

$$(2.6) \rightarrow \bar{\nabla} \bar{g}(\bar{x}) = G \rho(\bar{x}) \int_{|\bar{x}' - \bar{x}| \leq h} d^3\bar{x}' \bar{\nabla}_x \left( \frac{\bar{x}' - \bar{x}}{|\bar{x}' - \bar{x}|^3} \right)$$

~~substituting the variables~~

$$= G \rho(\bar{x}) \int_{|\bar{x}' - \bar{x}| \leq h} d^3\bar{x}' \bar{\nabla}_x \left( \frac{\bar{x}' - \bar{x}}{|\bar{x}' - \bar{x}|^3} \right)$$

↑ small

$$= G \rho(\bar{x}) \int_{|\bar{x}' - \bar{x}| \leq h} d^3\bar{x}' \frac{(\bar{x}' - \bar{x})}{|\bar{x}' - \bar{x}|^3}$$

divergence theorem

$$\int_V \bar{\nabla} \cdot \vec{F} d^3\bar{x} = \int_S \vec{F} \cdot d\vec{s}$$

but over the sphere  $\vec{s} = (\bar{x}' - \bar{x})$   
 $d^2s = (\bar{x}' - \bar{x}) h d^2\Omega$

$$\bar{\nabla} \bar{g}(\bar{x}) = -G \rho(\bar{x}) \int d^2\Omega = -4\pi G \rho(\bar{x})$$

eq. 2.5

$$\bar{\nabla}^2 \phi = -4\pi G \rho$$

Poisson Eq.  
 (2.10) diff equation for isolated system  $\phi \rightarrow 0$  as  $|\bar{x}| \rightarrow \infty$

oriented /

In the special case  
Poisson eq  $\rightarrow$

$\rho=0$

$\nabla^2 \phi = 0$

Laplace eq.

(2.11)

BT 2  
NUR

We integrate eq. (2.10)

$\int \nabla^2 \phi d^3x = \int 4\pi G \rho(x) d^3x$

$M = \int \rho(x) d^3x$   
 $\hookrightarrow$  mass

$\int \nabla(\nabla \phi) d^3x = 4\pi G M$

2  
ole

$\nabla \phi = -\vec{g}$

the integral of  
the normal component  
on  $\bar{S}$   $\rightarrow$   $4\pi G M$   
contained  
in  $\bar{S}$

divergence  
theor.

$\int d^2\bar{S} \nabla \phi = 4\pi G M$

$\int \vec{g} d^2\bar{S} = -4\pi G M$

(remember?  $\int_S \vec{E} d^2\bar{S} = \frac{Q}{\epsilon_0}$  for  $\vec{E}$  field)

$\int d^2\bar{S} \nabla \phi = 4\pi G M$

integrale  
della componente  
normale di  $\nabla \phi$  (2.12)

The Gauss theorem

su una  $\partial V$  superficie  
chiusa equivale a  $4\pi G \rho$   
messa contenuta

$\vec{g} = -\nabla \phi$   
 $\vec{g}$  comes from  $\nabla \phi$  so is conservative! in  
(un gradiente del potenziale) quella  
superficie

lavoro fatto contro le forze gravitazionali  
per assemblare una certa configurazione  
di masse  $\vec{E}$  indipendente dai  
dettagli dell'assemblaggio

# Potential energy

$\rho(\vec{x})$   $\phi(\vec{x})$   $\delta\phi$   
 so there is an amount of potential energy

a mass is already in place like bring an additional small mass  $\delta m$  from  $\infty$  to  $\vec{x}$ , the work done is  $\delta m \phi(\vec{x})$

So a small increment of density  $\delta\rho(\vec{x})$  gives a change in the potential energy (a small change) *in potential*

$$\delta W = \int d^3x \delta\rho(\vec{x}) \phi(\vec{x}) \quad (2.13)$$

Poisson eq.  $\nabla^2(\delta\phi) = 4\pi G(\delta\rho) \rightarrow$

$$\delta W = \int d^3x \phi(\vec{x}) \frac{1}{4\pi G} \nabla^2(\delta\phi) =$$

$$= \frac{1}{4\pi G} \int d^3x \phi(\vec{x}) \nabla^2(\delta\phi) =$$

$$= \frac{1}{4\pi G} \int d^3x \underbrace{\phi(\vec{x})}_P \underbrace{\nabla^2(\delta\phi)}_{\vec{F}}$$

do integrate x part  
 $\int P \vec{F} d^3x = P \vec{F} - \int P \vec{F}' d^3x$   
 $\frac{\partial}{\partial x} P$

<sup>4</sup> extended theorem of divergence sec Appendix of Binney and Tremaine

$$\int_V \nabla \cdot \vec{F} d^3x = \int_S \vec{F} \cdot d\vec{S} - \int_V (\vec{F} \cdot \nabla) \phi d^3x$$

$$\delta W = \frac{1}{4\pi G} \int \phi \nabla^2(\delta\phi) d^3x - \frac{1}{4\pi G} \int d^3x \nabla(\delta\phi) \cdot \nabla \phi(\vec{x}) \quad (2.15)$$

$\Rightarrow$  why  $\phi \propto r^{-1}$   $|\nabla(\delta\phi)| \propto r^{-2}$  as  $r \rightarrow \infty$   $\Rightarrow$  surface where  $r \rightarrow \infty$

$\underbrace{\phi \nabla(\delta\phi)}_{\text{integrand}} \propto r^{-3}$  while the area  $\propto r^2$

Moreover, for the second term

$$\vec{\nabla}\phi \cdot \vec{\nabla}(\delta\phi) = \vec{\nabla}\phi \cdot \delta(\vec{\nabla}\phi) =$$

$$\left( \begin{aligned} \text{In fact } \delta(\vec{\nabla}\phi) &= \vec{\nabla}(\phi + \delta\phi) - \vec{\nabla}\phi = \\ &= \vec{\nabla}\phi + \vec{\nabla}(\delta\phi) - \vec{\nabla}\phi = \vec{\nabla}(\delta\phi) \end{aligned} \right)$$

$$\equiv \frac{1}{2} \delta(\vec{\nabla}\phi \cdot \vec{\nabla}\phi)$$

remember? In fact  
For a vector  $\vec{v}$   $\frac{1}{2} d|\vec{v}|^2 = \frac{1}{2} d(\vec{v} \cdot \vec{v}) = \frac{1}{2} 2 \vec{v} d\vec{v}$

$$\rightarrow = \frac{1}{2} \delta |\vec{\nabla}\phi|^2$$

2.15  $\rightarrow$

$$\delta W = -\frac{1}{4\pi G} \int d^3x \frac{1}{2} \delta |\vec{\nabla}\phi|^2 =$$

(sum of variations  
= variation of the sum)

$$= -\frac{1}{8\pi G} \delta \left( \int d^3x |\vec{\nabla}\phi|^2 \right) \quad \text{2.16}$$

Now we sum up all of the contributions  $\delta W$

$$W = \int \delta W$$

$$W = -\frac{1}{8\pi G} \int d^3x |\vec{\nabla}\phi|^2 \quad \text{2.17}$$

Energy  
potentielle  
per un corps  
assemble

To obtain an alternative expression for  $W$

$$2.17 \rightarrow W = -\frac{1}{8\pi G} \int d^3x \nabla\phi \cdot \nabla\phi$$

but we know that  $\nabla(\lambda\vec{u}) = \lambda\nabla\vec{u} + \vec{u}\nabla\lambda$

$$\nabla(\phi\nabla\phi) = \phi\nabla(\nabla\phi) + \nabla\phi \cdot \nabla\phi$$

$$\rightarrow \nabla\phi \cdot \nabla\phi = \nabla(\phi\nabla\phi) - \phi\nabla^2\phi$$

$$\rightarrow = -\frac{1}{8\pi G} \left[ \int d^3x \nabla(\phi\nabla\phi) - \int d^3x \phi\nabla^2\phi \right] =$$

div. theorem.

$$\int d^3x \phi \nabla^2\phi$$

$\int d^3x \phi \nabla^2\phi = \int d^3x \phi \nabla^2\phi$

$\int d^3x \phi \nabla^2\phi = \int d^3x \phi \nabla^2\phi$

$\int d^3x \phi \nabla^2\phi = \int d^3x \phi \nabla^2\phi$

use eq. of Poisson

$$\nabla^2\phi = 4\pi G\rho$$

$$= \frac{1}{8\pi G} \int d^3x 4\pi G\rho\phi$$

$$W = \frac{1}{2} \int d^3x \rho(\vec{x})\phi(\vec{x})$$

Potential energy

for a body

2.18

Chandrasekhar

The potential - energy tensor

tensor  $\underline{\underline{W}}$   $\begin{pmatrix} \dots & \dots & \dots \\ \dots & \dots & \dots \\ \dots & \dots & \dots \end{pmatrix}$   $3 \times 3$

general component  $W_{JK} \equiv - \int d^3 \bar{x} \rho(\bar{x}) x_J \frac{\partial \phi}{\partial x_K}$   
 over all space

2.19  
 SS eq. 2.19

2.3  
 $\phi(\bar{x}) = -G \int d^3 \bar{x}' \frac{\rho(\bar{x}')}{|\bar{x}' - \bar{x}|}$   
 definite minus

the trace of  $\underline{\underline{W}}$  sense  $\underline{\underline{W}}$   
 $\text{trace}(\underline{\underline{W}}) = \underline{\underline{W}}$

~~$W_{JK} = G \int d^3 \bar{x} \int d^3 \bar{x}' (\rho(\bar{x}) x_J \frac{\partial}{\partial x_K})$~~

$W_{JK} = G \int d^3 \bar{x} \rho(\bar{x}) x_J \frac{\partial}{\partial x_K} \int d^3 \bar{x}' \frac{\rho(\bar{x}')}{|\bar{x}' - \bar{x}|}$  = 2.20

\* solo componente  $K$  does not depend on  $\bar{x}$

$W_{JK} = G \int d^3 \bar{x} \int d^3 \bar{x}' \rho(\bar{x}) \rho(\bar{x}') \frac{x_J (x'_K - x_K)}{|\bar{x}' - \bar{x}|^3}$  = 2.21 e

$x$  and  $x'$  are dummy variables of integrals.

$W_{JK} = G \int d^3 \bar{x}' \int d^3 \bar{x} \rho(\bar{x}') \rho(\bar{x}) \frac{x'_J (x_K - x'_K)}{|\bar{x} - \bar{x}'|^3}$  = 2.21 b  
 change of the order of integration  
 $= G \int d^3 \bar{x} \int d^3 \bar{x}' \rho(\bar{x}) \rho(\bar{x}') \frac{(-x'_J)(x'_K - x_K)}{|\bar{x}' - \bar{x}|^3}$

2.21 e + b

$2W_{JK} = G \int d^3 \bar{x} \int d^3 \bar{x}' \rho(\bar{x}) \rho(\bar{x}') \cdot (-1) \frac{(x'_J - x_J)(x'_K - x_K)}{|\bar{x}' - \bar{x}|^3}$

$W_{JK} = -\frac{1}{2} G \int d^3 \bar{x} \int d^3 \bar{x}' \rho(\bar{x}) \rho(\bar{x}') \frac{(x'_J - x_J)(x'_K - x_K)}{|\bar{x}' - \bar{x}|^3}$  = 2.22



8

2.22

→  $\bar{W}$  is symmetric

$$W_{jk} = W_{kj}$$

trace

$$\begin{pmatrix} 0 & & \\ & 0 & \\ & & 0 \end{pmatrix}$$

Sol<sup>3</sup> terms  
 $\bar{S} = K$

$$\text{trace}(\bar{W}) = \sum_{j=1}^3 W_{jj} =$$

$$= \left(-\frac{1}{2}G\right) \int d^3\bar{x} \rho(\bar{x}) \int d^3\bar{x}' \frac{\rho(\bar{x}')}{|\bar{x}' - \bar{x}|^3} |\bar{x}' - \bar{x}|^2$$

def. of  $\phi$

$$= \frac{1}{2} \int d^3\bar{x} \rho(\bar{x}) \phi(\bar{x})$$

is eq. 2.18 for  $W$

2.24

$\text{trace}(\bar{W}) = W$

If we make directly the trace using 2.19

$$\begin{aligned} \sum_{j=1}^3 W_{jj} &= - \sum_{j=1}^3 \int d^3\bar{x} \rho(\bar{x}) x_j \frac{\partial \phi}{\partial x_j} = \\ &= - \int d^3\bar{x} \rho(\bar{x}) \bar{x} \cdot \bar{\nabla} \phi \end{aligned}$$

So

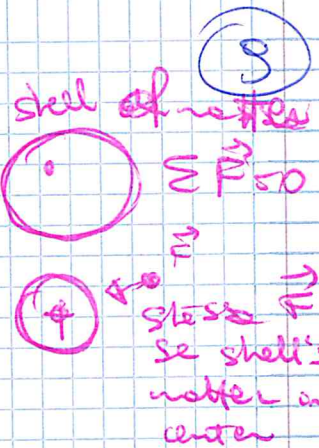
$$W = - \int d^3\bar{x} \rho(\bar{x}) \bar{x} \cdot \bar{\nabla} \phi$$

2.24

Alternative expression for the potential energy of a body

# Spherical systems

Newton's Theorems  $\rightarrow$   
(page 60)



same COROLLARY del  $\vec{\nabla} \phi = 0$

gravitational potential inside an empty shell is constant since

$$\vec{\nabla} \phi = -\vec{g} = 0 \Rightarrow \phi = \text{const.}$$

$$\phi(r) = ?$$

$\leftarrow$  from eq. 2.3 def. of gravit. potential.

$$\phi(\vec{x}) = -G \int d^3x' \frac{\rho(x')}{|\vec{x}' - \vec{x}|}$$

$$\hookrightarrow \phi = -\frac{GM}{R} \quad (2.25)$$

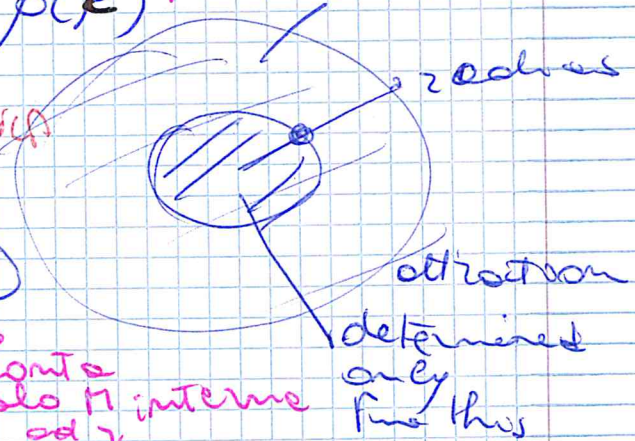
of central & quadrupole moment  
 $\vec{x} = 0 \quad x' = R$

FOR 1 shell!

shell's own center  $\times$  potential density is  $\rho(r)$

$$\hookrightarrow \text{also } \phi = -\frac{G}{R} \int dV \rho(\vec{x}')$$

ATTENZIONE GRAVITAZIONALE DI UNA DISTRIBUZIONE DI DENSITA'  $\rho$  FROM I and II Theorems DIFFERENZA



$$\vec{F}(\vec{x}) = -\frac{GM(r)}{r^2} \hat{e}_r \quad (2.27)$$

where

$$M(r) = 4\pi \int_0^r dr' r'^2 \rho(r')$$

conta solo  $M$  interne ad  $r$

Total gravitational potential is given by the sum of the potential of the shells of mass

$$dM(r) = 4\pi \rho(r) r^2 dr$$

$$\phi(r) = -\frac{G}{r} \int_0^r dM(r') - G \int_r^\infty \frac{dM(r')}{r'}$$

internal shells external

we potential essenza ad un punto  $\rightarrow$