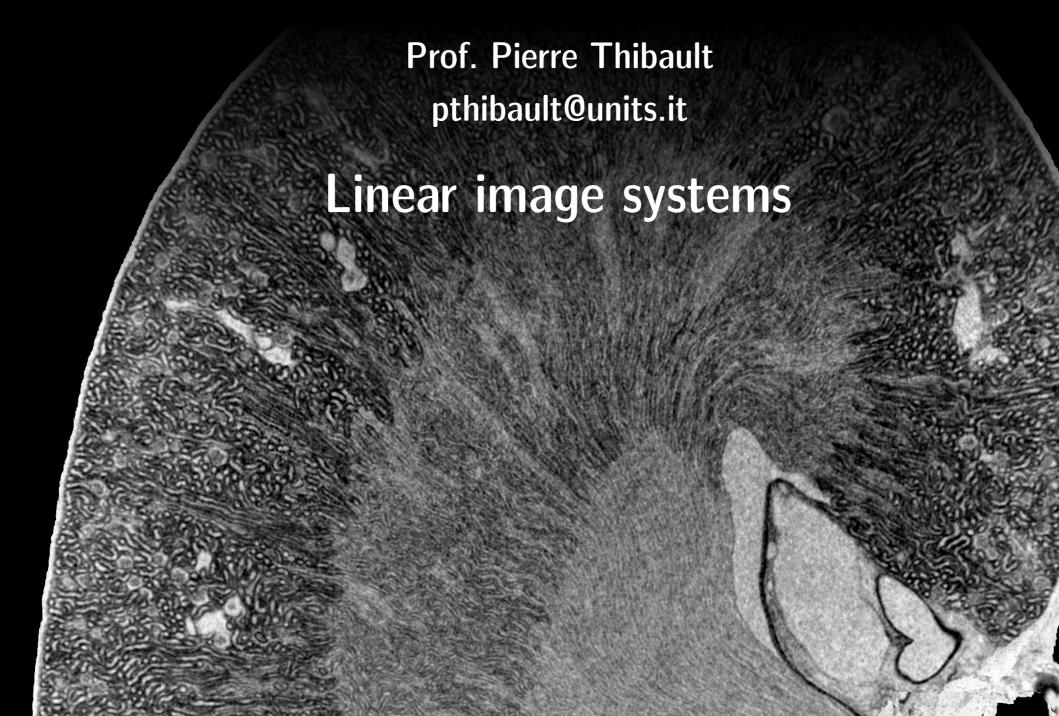
Image Processing for Physicists



Overview

- Definition of resolution
- Imaging systems:
 - Linear transfer model
 - Noise

Resolution

"the smallest detail that can be distinguished"

- No unique definition
 - Numerical aperture
 - Pixel size
 - Other criteria (PSF, MTF)
- What is "detail"?
- What is "distinguish"?

Resolution

1280 x 1280





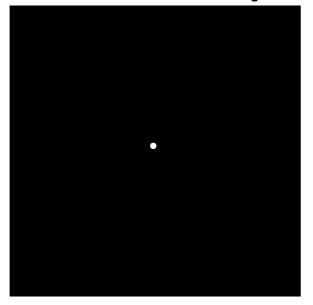
- **not** simply given by pixel size (i.e. sampling rate)
- light quality, optics quality, detector quality, algorithm quality, noise, ...

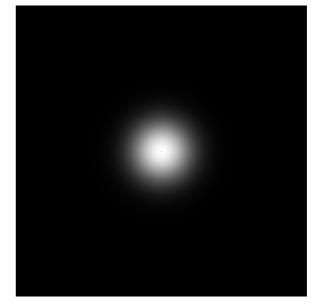
Linear translation-invariant systems

• Point spread function ("impulse response")

• LTI system: convolution with PSF

Point spread function

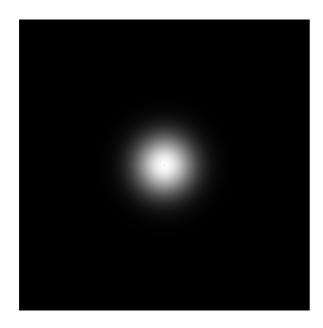








PSF and resolution

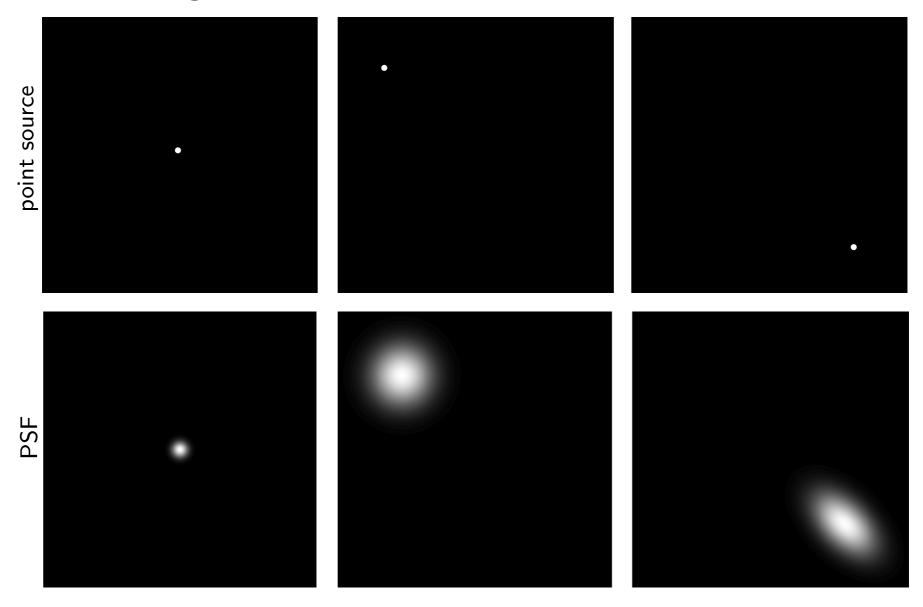


Measurement of the PSF

• Direct measurement from impulse

Line-spread function

PSF and translation invariance



- Not translation invariant o PSF depends on position o not a convolution
- Useful to model system imperfections, lens aberrations, ...

The Fourier picture

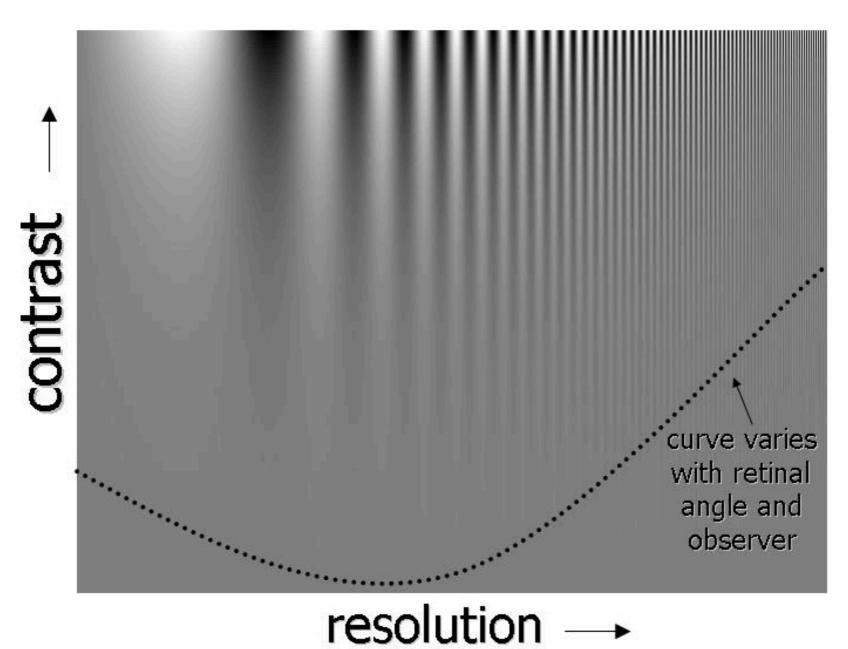
Optical transfer function

Response of a system to an oscillating signal with well-defined frequency

Modulation transfer function

Amplitude change of an oscillating signal for a given frequency

Eye MTF



Campbell-Robson curve

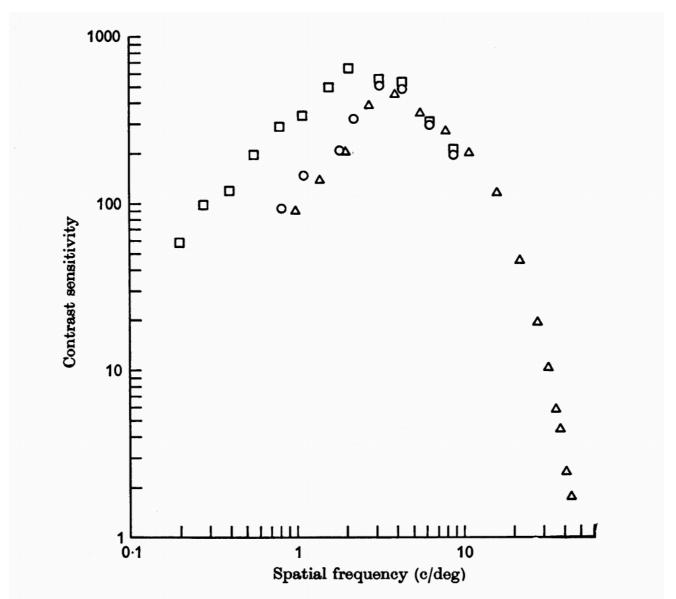
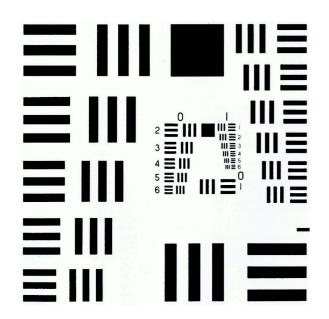
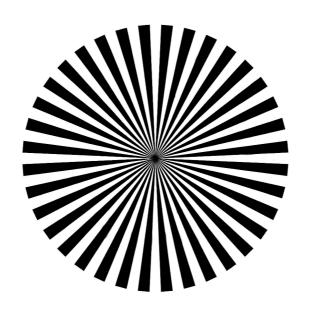
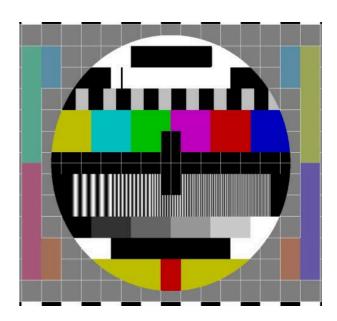


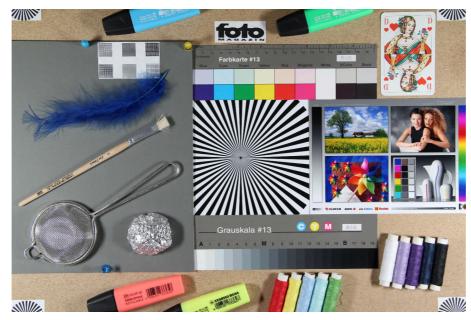
Fig. 2. Contrast sensitivity for sine-wave gratings. Subject F.W.C., luminance 500 cd/m^2 . Viewing distance 285 cm and aperture $2^{\circ} \times 2^{\circ}$, \triangle ; viewing distance 57 cm, aperture $10^{\circ} \times 10^{\circ}$, \square ; viewing distance 57 cm, aperture $2^{\circ} \times 2^{\circ}$, \bigcirc .

Measurement of MTF





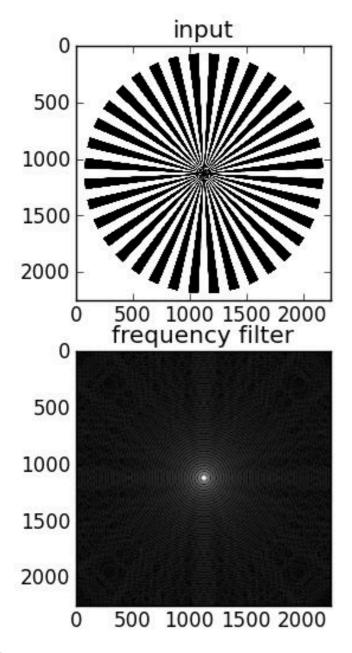


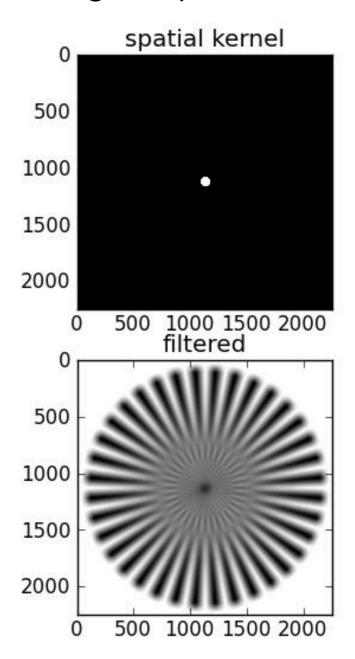


source: http://fotomagazin.de

Phase transfer function

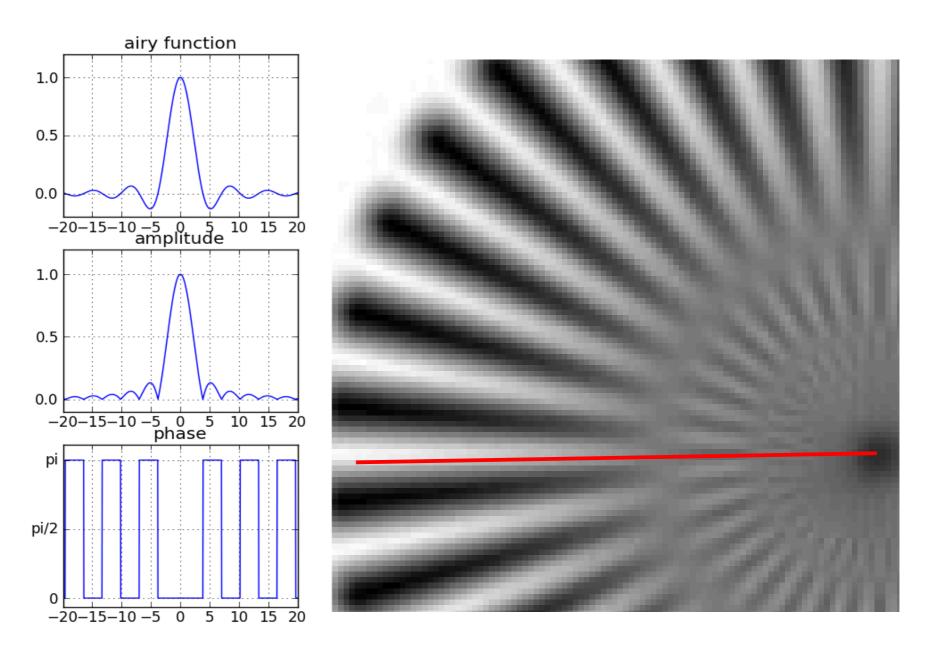
describes how an oscillating signal changes in phase due to system





Phase transfer function

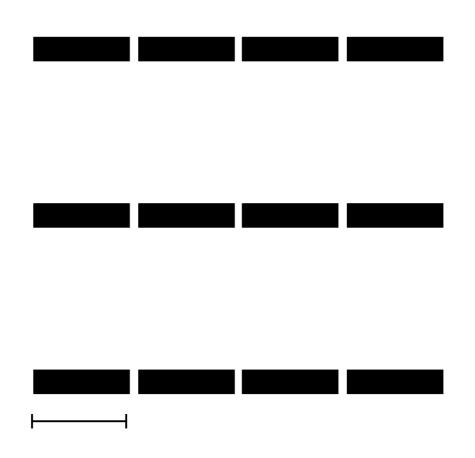
describes how an oscillating signal changes in phase due to system



MTF of an ideal pixel

Pixel MTF

Modulation transfer function of a single detector pixel

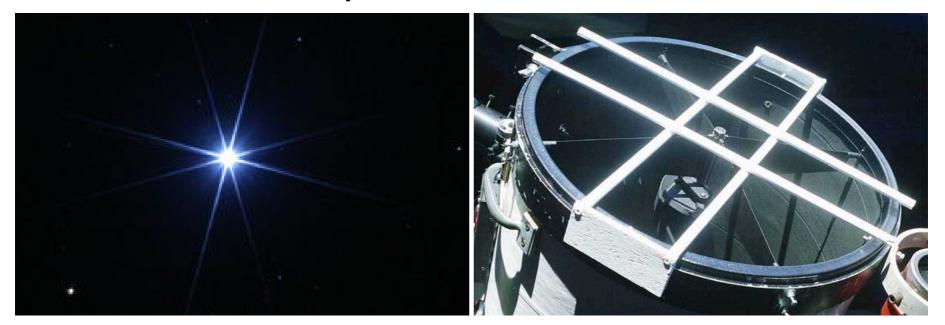


Imaging as a linear filter



PSF examples

isolated stars are essentially PSFs

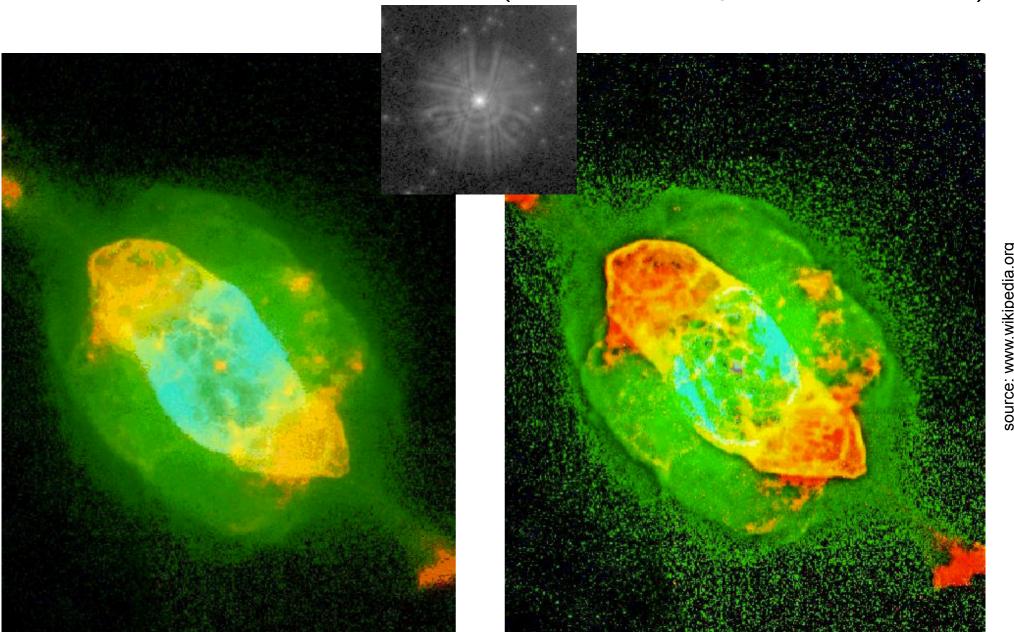




source: www.apod.nasa.gov

PSF examples

Hubble flawed mirror deconvolution (correction for spherical aberration)

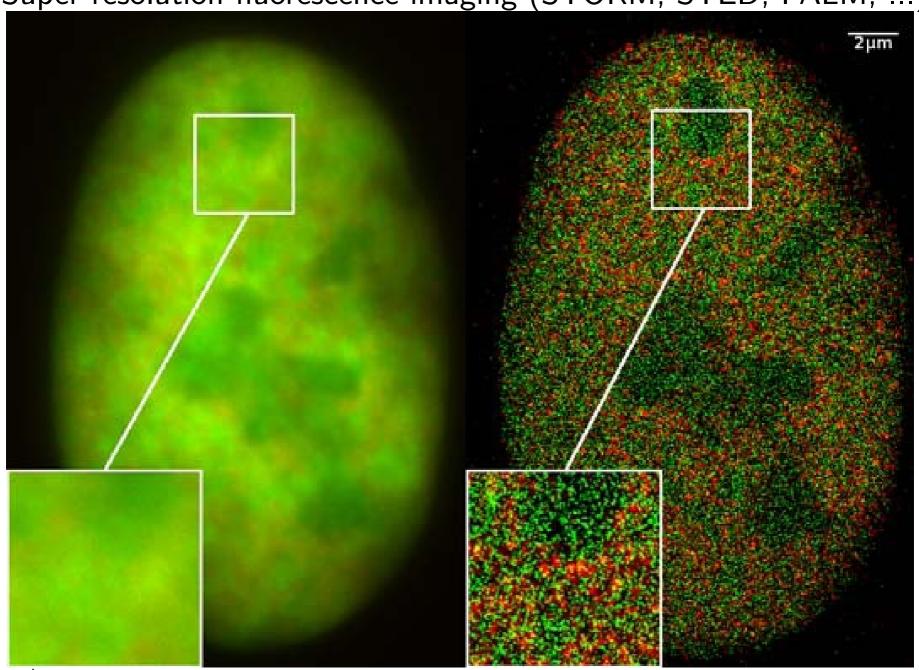


Imaging systems

source: www.wikipedia.org

PSF examples

Super-resolution fluorescence imaging (STORM, STED, PALM, ...)



Contrast and noise

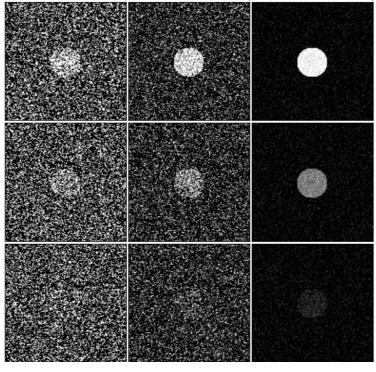
 Intensity operation: higher contrast, higher noise

 Contrast-to-noise remains constant









Decreasing noise

Random variables

random variable, sample space

probability density function

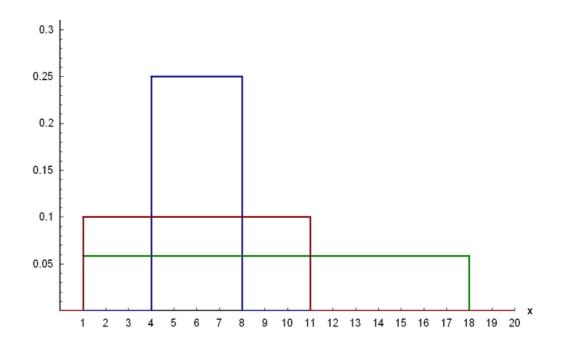
expectation value

variance

Uniform distribution

probability density function

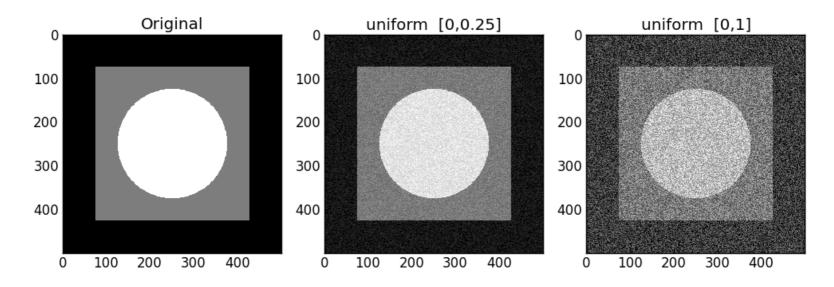
expectation value

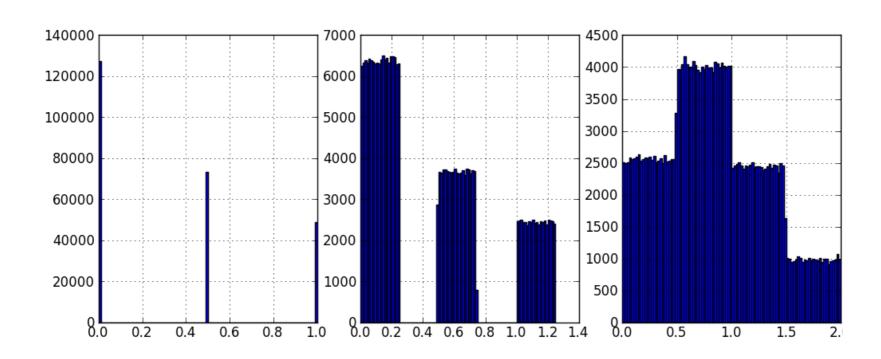


variance

occurrence

Uniform distribution

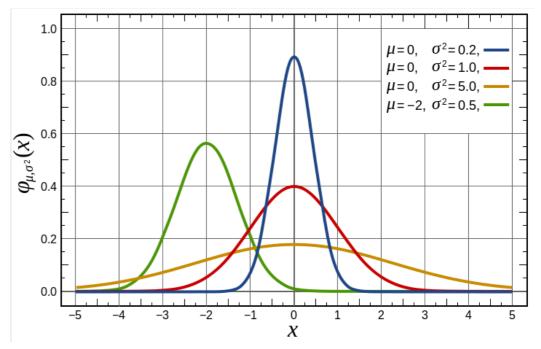




Gaussian distribution

probability density function

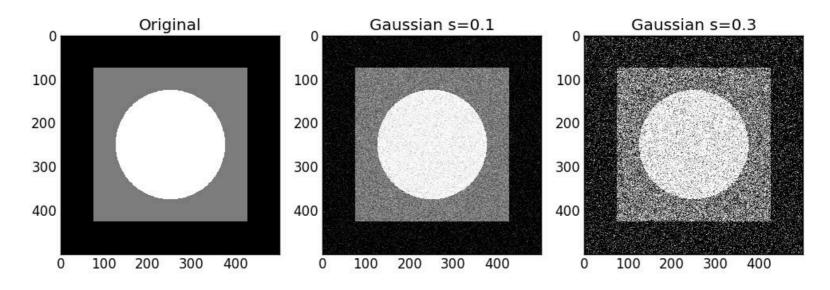
expectation value

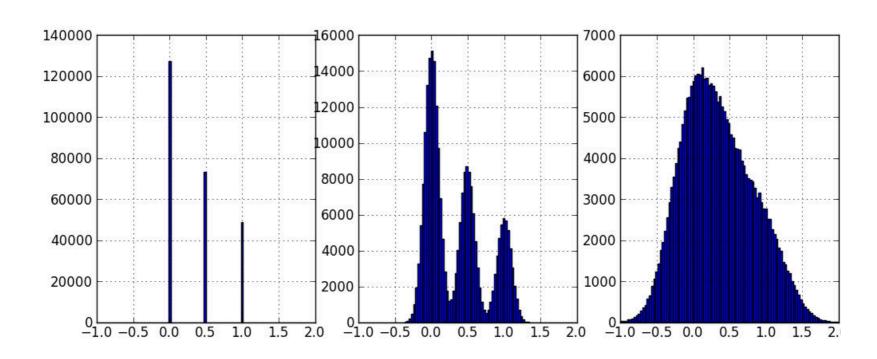


variance

occurrence

Gaussian distribution

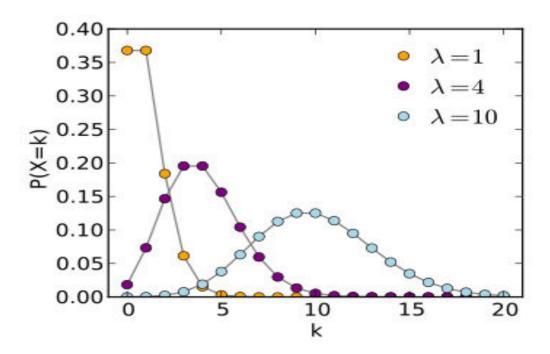




Poisson distribution

probability mass function

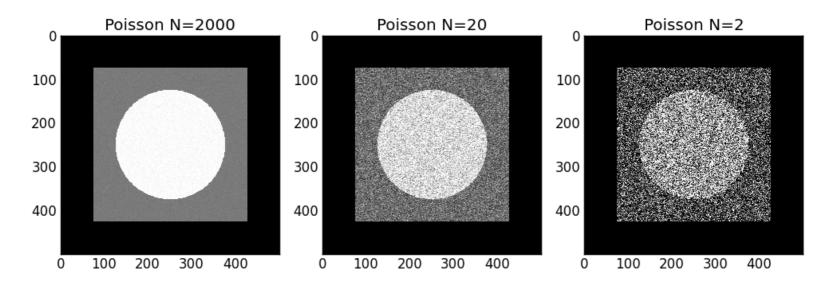
expectation value

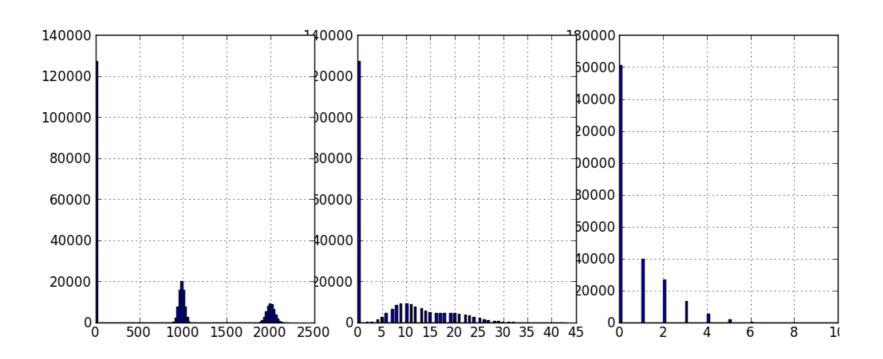


variance

occurrence

Poisson distribution



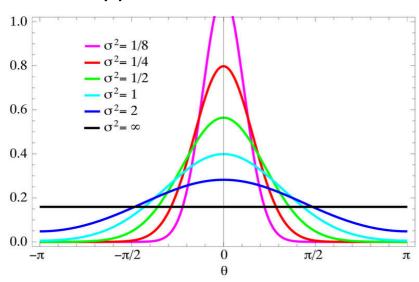


Poisson distribution

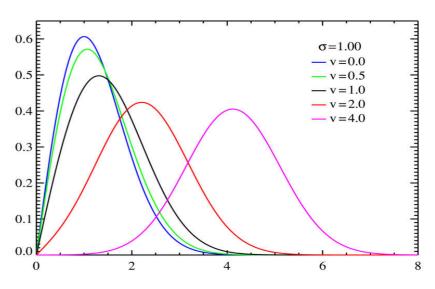


Many other distributions

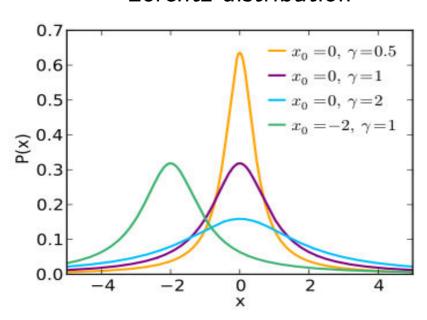
Wrapped normal distribution



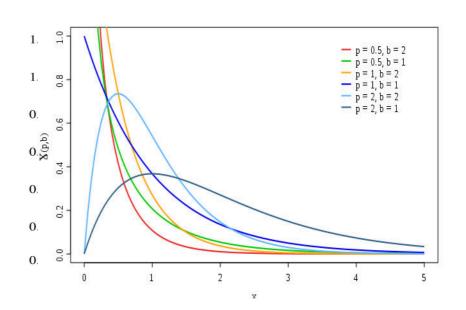
Rice distribution



Lorentz distribution

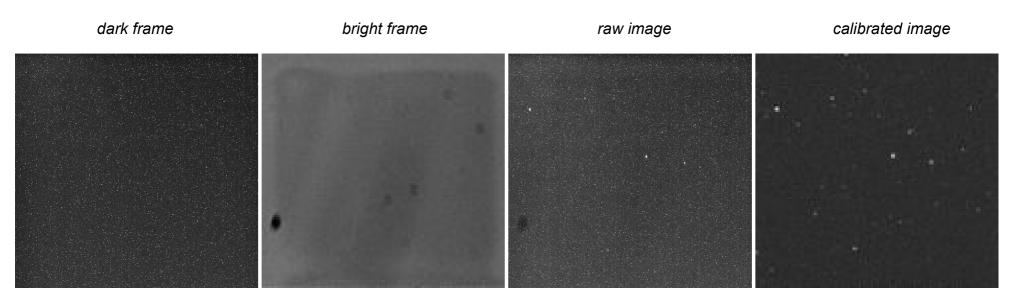


Gamma distribution



Detector noise (CCD)

- Various sources:
 - shot noise (photon statistics, Poisson)
 - dark current (thermal electronic fluctuations in semiconductor, Poisson)
 - readout noise (fluctuations during amplification and digitization, Gauss)
 - many other imperfections ...
- dark frame measures detector noise, hot pixels, dead pixels
- bright frame measures gain differences and imperfections (dust, etc)



source: H. Raab, Johannes-Kepler-Observatory, Linz

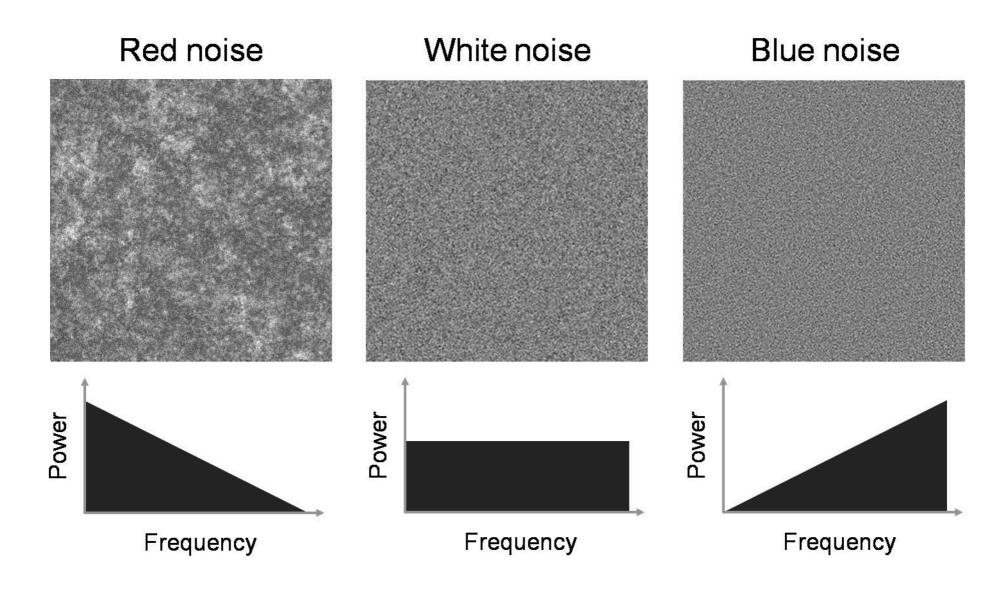
Correlation & Convolution

Noise power spectrum

power spectrum of pure noise image

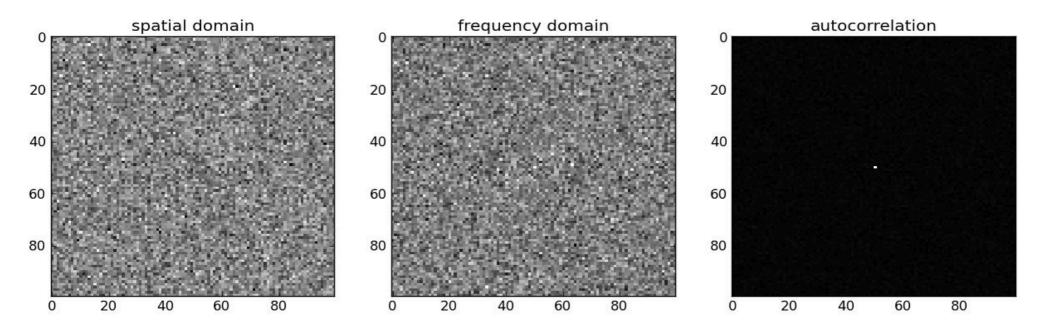
connection to auto-correlation

Noise power spectrum

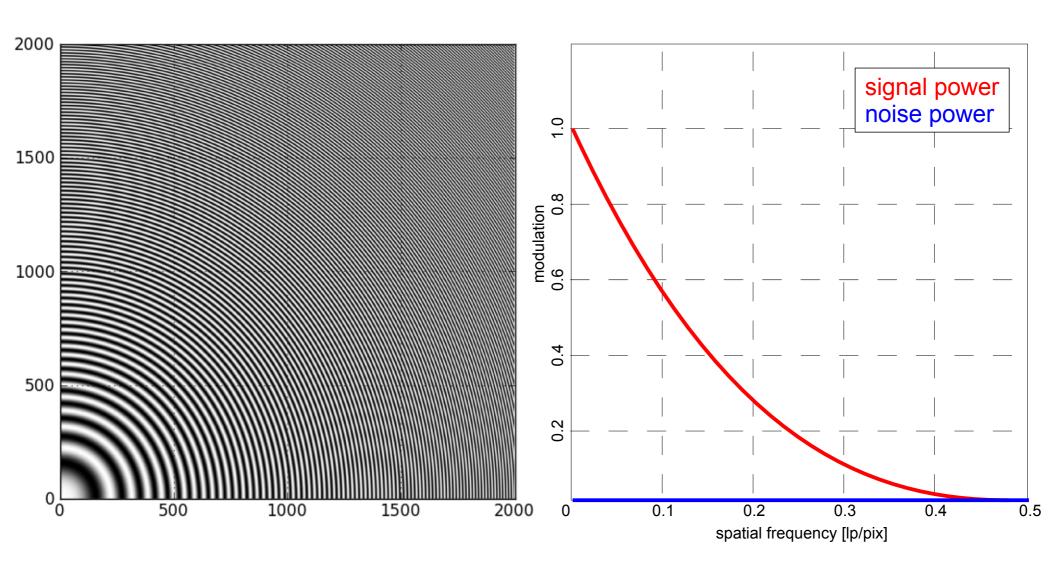


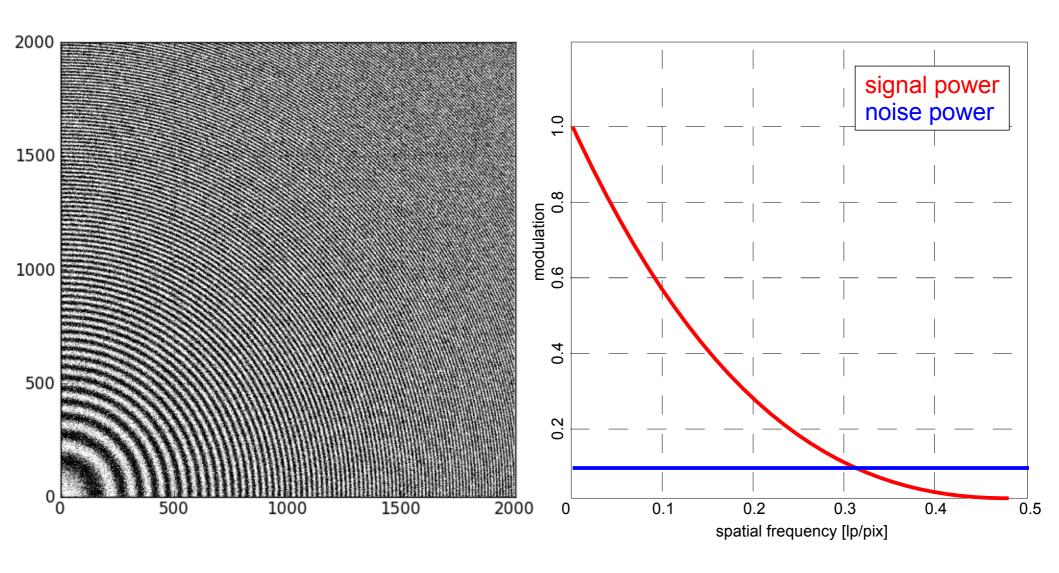
source: http://scien.stanford.edu/pages/labsite/2008/psych221/projects/08/AdamWang/project_report.htm

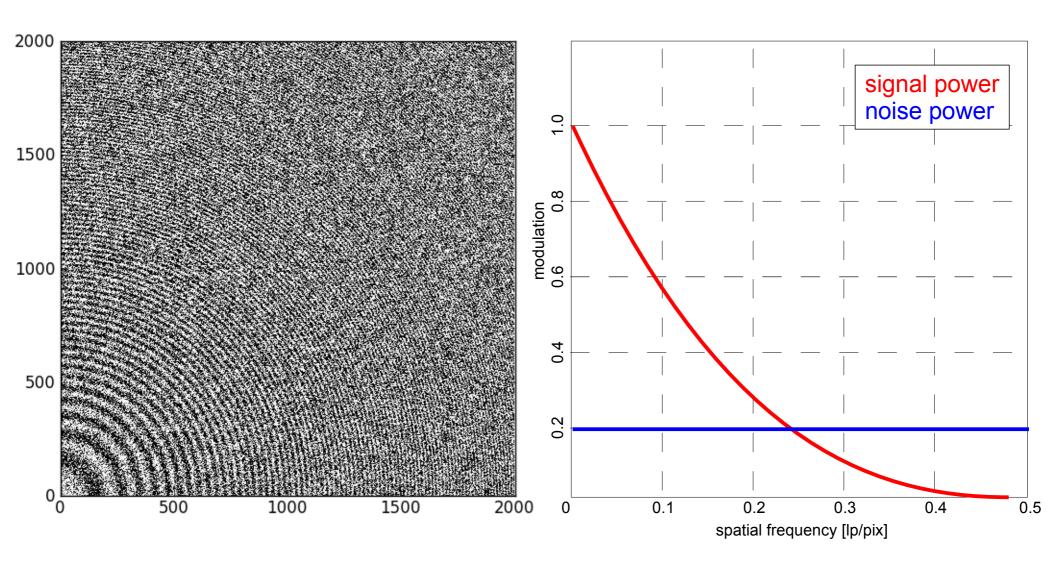
White noise

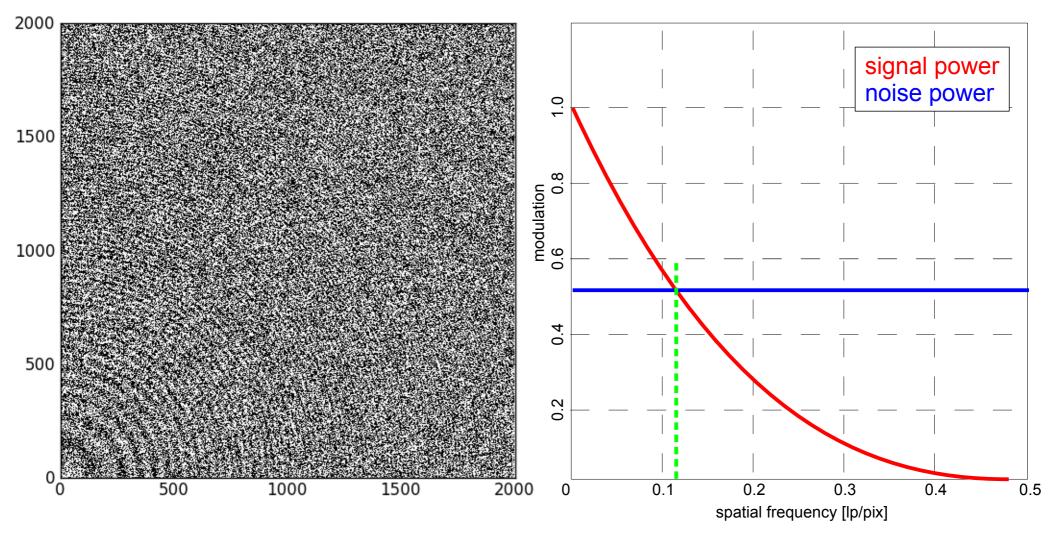


- white noise in spatial domain equals white noise in frequency domain
- white noise is perfectly uncorrelated
- all other types of noise are correlated to some degree
- white noise is an idealization

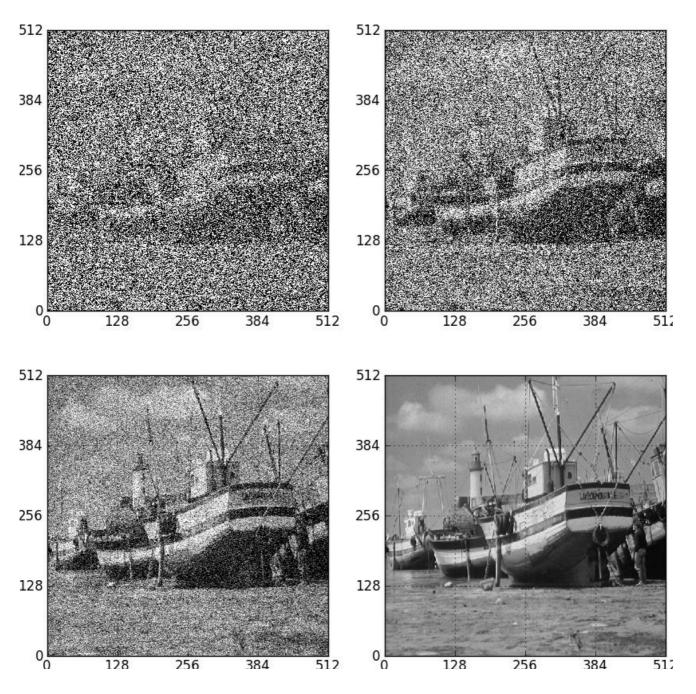






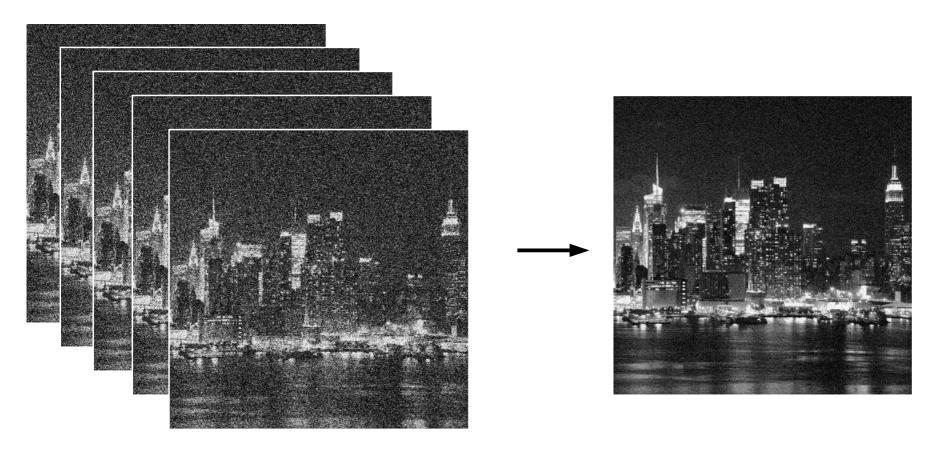


- Noise power exceeds signal power for high frequencies
- Small scale image details are lost in noise first



Noise reduction by averaging

Average multiple images



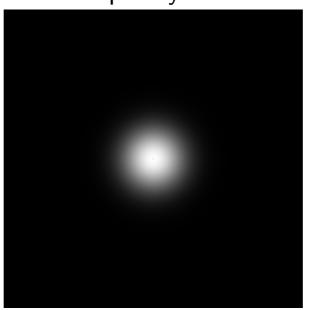
requirement: additive noise, zero mean

Denoising by linear filtering

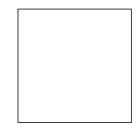
- use spatial convolution or frequency filtering to reduce noise
- noise reduction
 possible, but at cost
 of sharpness
- trade-off between noise reduction and resolution
- need fancier methods

original

frequency filter



convolution kernel



Resulting image



Median filtering

Use median as estimator for fat tail distributions

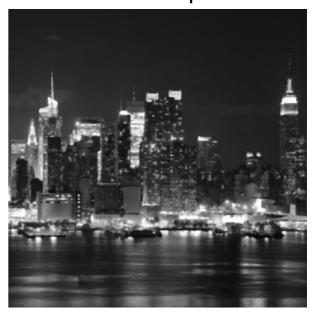
• less sensitive to outliers in pixel ensemble, better edge preservation

Salt and pepper noise

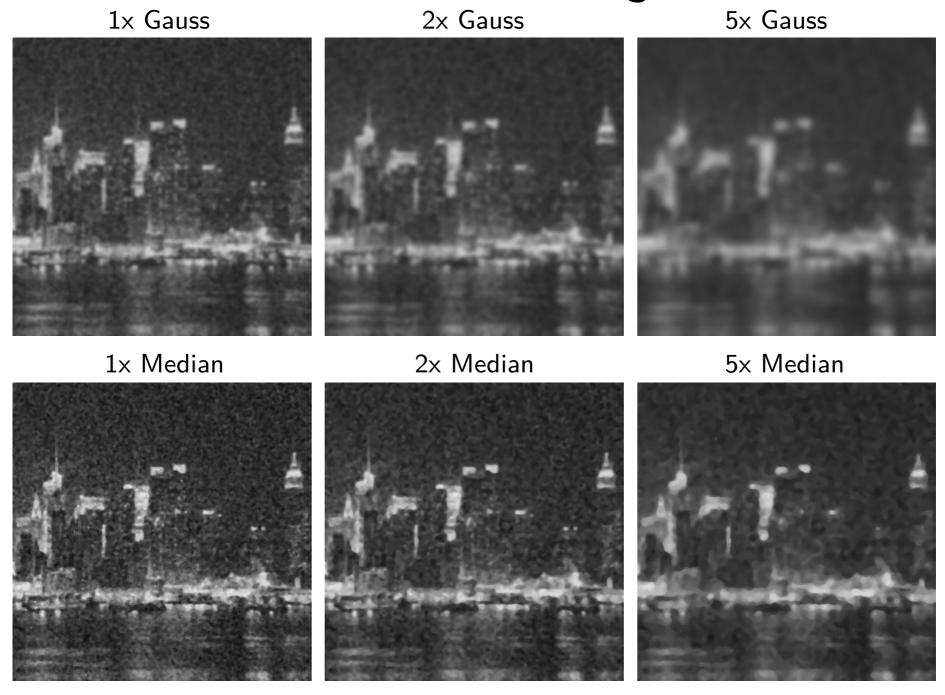
Gauss sigma=1 pixel



Median 1 pixel



Median filtering



Common abbreviations

Abbreviation	Name	Definition
IRF	Impulse response function	Linear operator map of delta function
PSF	Point spread function	Image of point object (optical IRF)
OTF	Optical transfer function	Fourier transform of PSF
PTF	Phase transfer function	Phase part of OTF
MTF	Modulation transfer function	Amplitude of OTF
CTF	Contrast transfer function	MTF for non-sinusoidal objects
PDF	Probability density function	Probability distribution for a given random variable
SPS	Signal power spectrum	Amplitude squared of signal F.T.
NPS	Noise power spectrum	Amplitude squared of noise F.T.
SNR	Signal to noise ratio	Mean signal / mean noise
CNR	Contrast to noise ratio	Mean contrast / mean noise