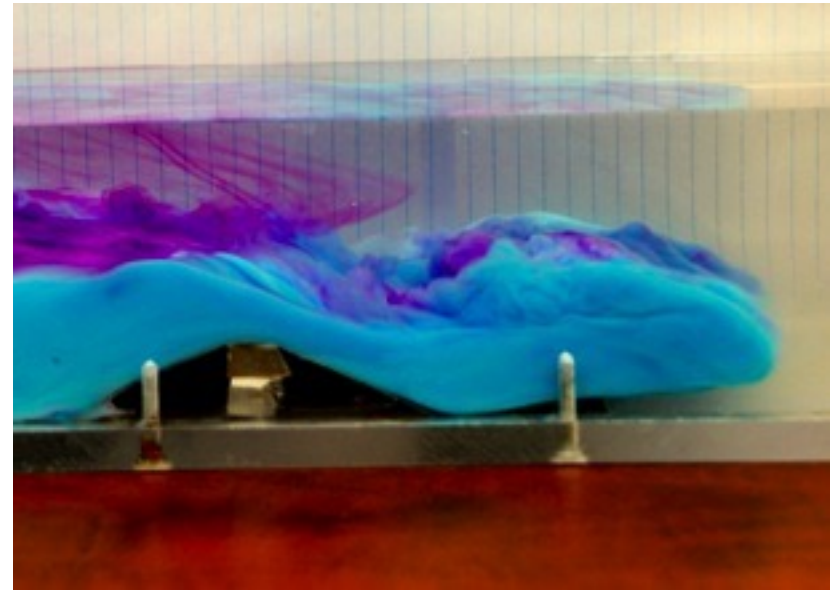


STRATIFICATION



STRATIFICATION induces a certain degree of decoupling between the various fluid masses which have different densities (vertical layering)
=> stratified systems contain more degrees of freedom than homogeneous systems
=> stratified systems exhibit additional types of motion

11.2 Static stability

Let us first consider a fluid in static equilibrium. Lack of motion can occur only in the absence of horizontal forces and thus in the presence of horizontal homogeneity. Stratification is then purely vertical (Figure 11-1).

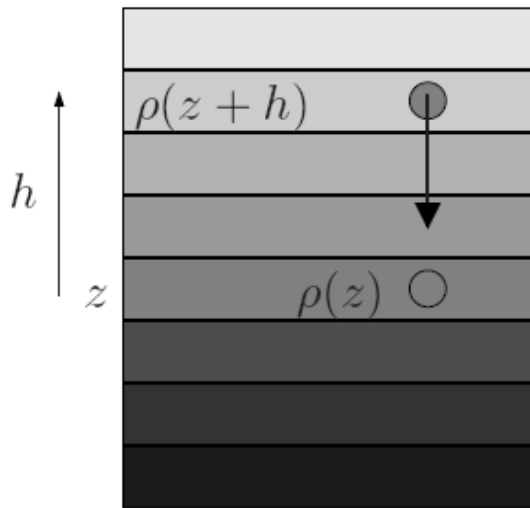


Figure 11-1 When an incompressible fluid parcel of density $\rho(z)$ is vertically displaced from level z to level $z+h$ in a stratified environment, a buoyancy force appears because of the density difference $\rho(z) - \rho(z+h)$ between the particle and the ambient fluid.

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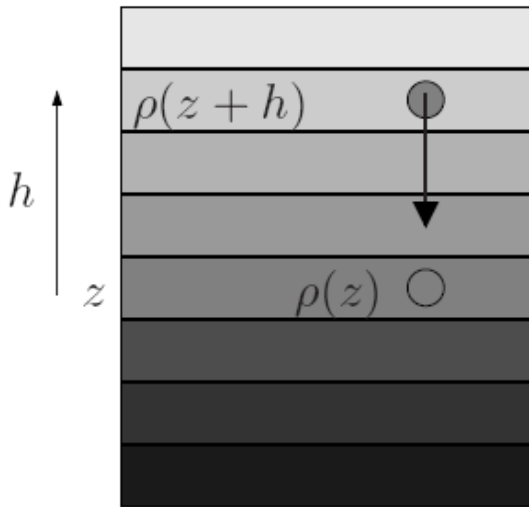


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With **incompressible fluids**, the displaced parcel retains its former density and at the new level is subject to a *net downward force equal to its own weight minus the weight of the displaced fluid* (Archimede's principle):

$$\rho(z) V \frac{d^2 h}{dt^2} = g [\rho(z + h) - \rho(z)] V$$

[...] using the **Boussinesq approx***

$$\rho(z + h) - \rho(z) \simeq \frac{d\rho}{dz} h$$

$$\frac{d^2 h}{dt^2} - \frac{g}{\rho_0} \frac{d\rho}{dz} h = 0$$

=> 2 cases: **STABLE** and **UNSTABLE**

=> $\frac{d\rho}{dz} < 0$ and $\frac{d\rho}{dz} > 0$ [...]

*not like an air balloon in seawater!

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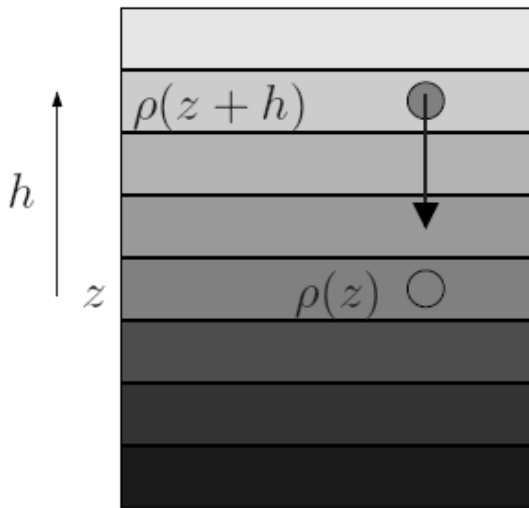


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$$\rho(z + h) - \rho(z) \simeq \frac{d\rho}{dz} h$$

$$N^2 = - \frac{g}{\rho_0} \frac{d\rho}{dz} \quad \leftarrow \quad \frac{d^2 h}{dt^2} - \frac{g}{\rho_0} \frac{d\rho}{dz} h = 0$$

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=> $\frac{d\rho}{dz} < 0$ and $\frac{d\rho}{dz} > 0$ [...]

BUOYANCY FORCE

We will restrict to **STABLY STRATIFIED FLUIDS => N^2**

Can we find an adimensional number with similar role as Ro ?

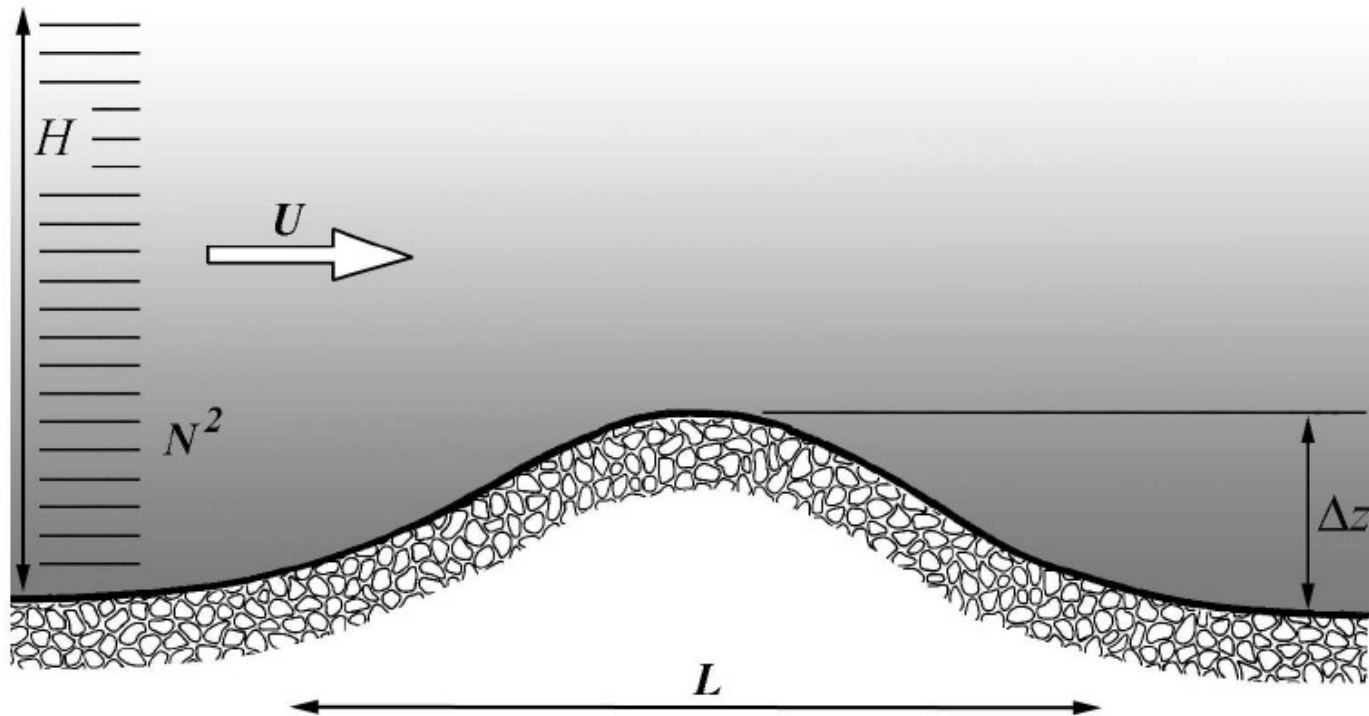


Figure 11-5 Situation in which a stratified flow encounters an obstacle, forcing some fluid parcels to move vertically against a buoyancy force.

Stratification will act to restrict or minimize the vertical displacement, forcing the flow to pass around rather than over the obstacle: the greater the restriction, the greater the importance of stratification

[...]

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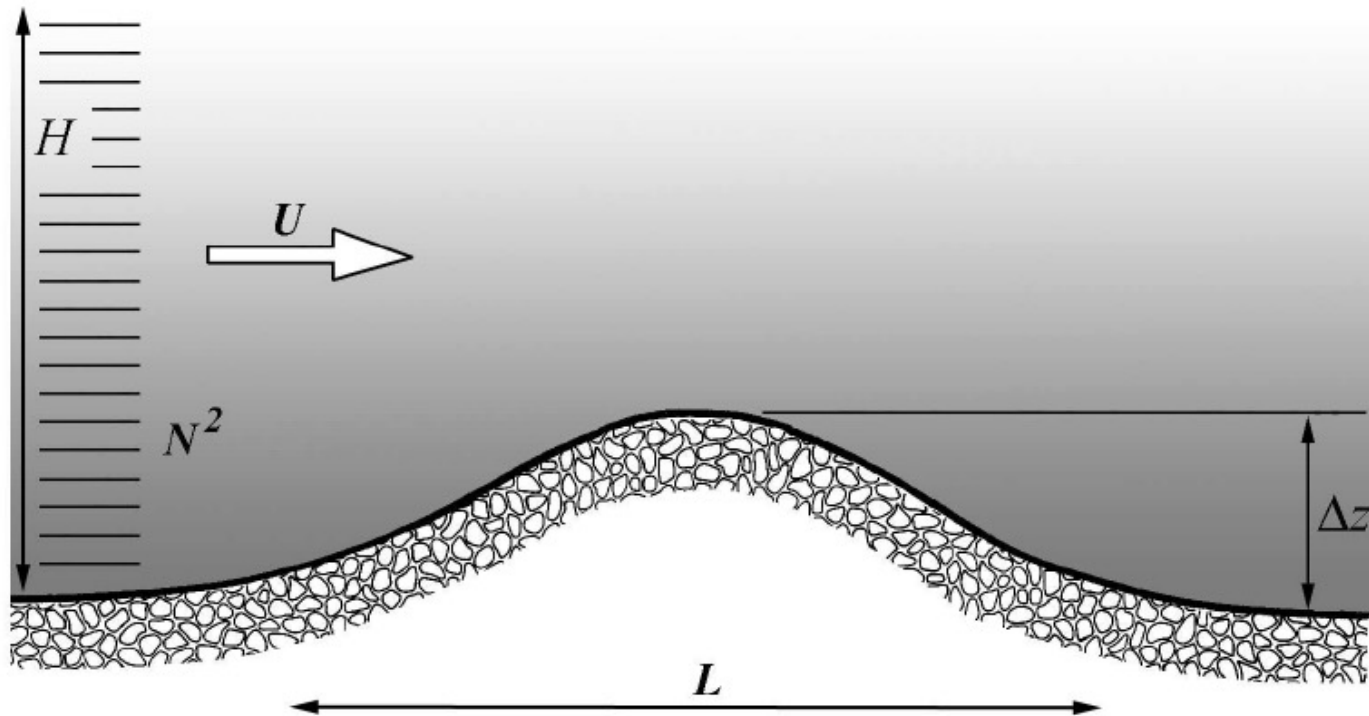
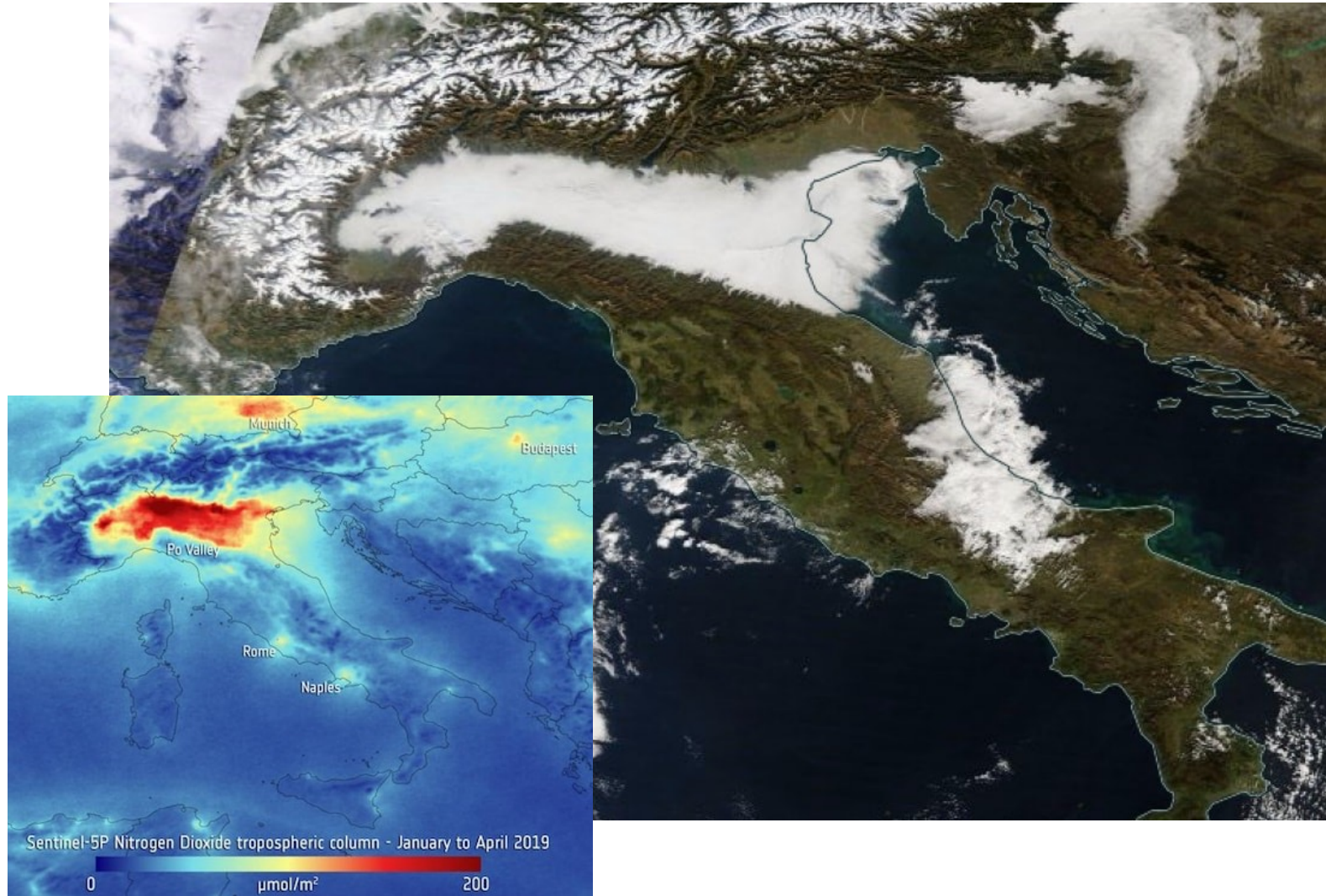


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$$[\dots] \frac{\text{vert.conv.}}{\text{horiz.div.}} = \frac{\partial_z w}{\partial_x u + \partial_y v} = \frac{\Delta z}{H} = \frac{W/H}{U/L} = \frac{U^2}{N^2 H^2} = Fr^2 \Rightarrow \text{Froude n.} \Rightarrow Fr \lesssim 1$$

HORIZONTAL BLOCKING: $Fr \ll 1$



Alps act as an obstacle: in case of high stratification vertical displacements are strongly suppressed and air quality gets worse

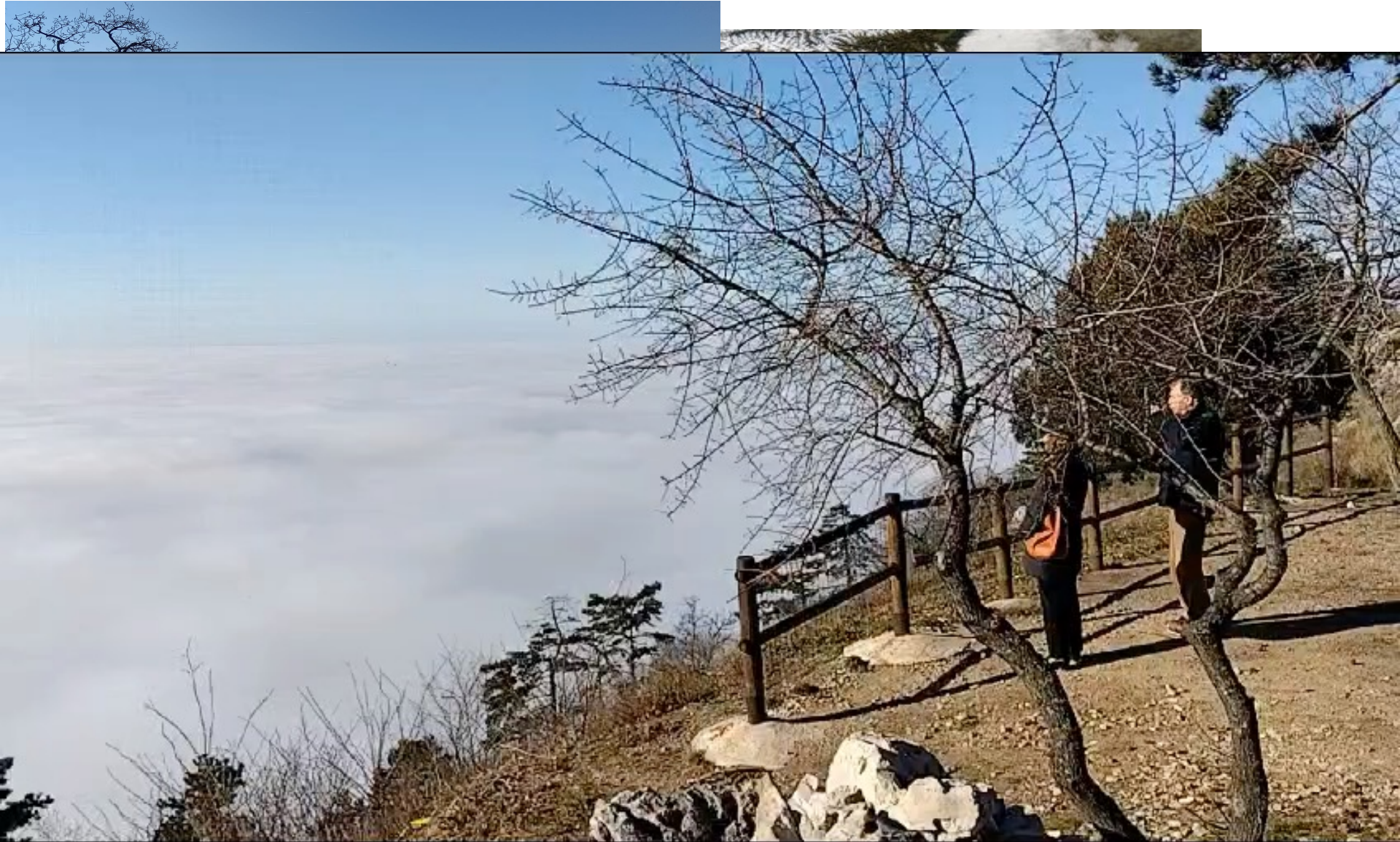
HORIZONTAL BLOCKING: $Fr \ll 1$



Also here
[28-1-2024]

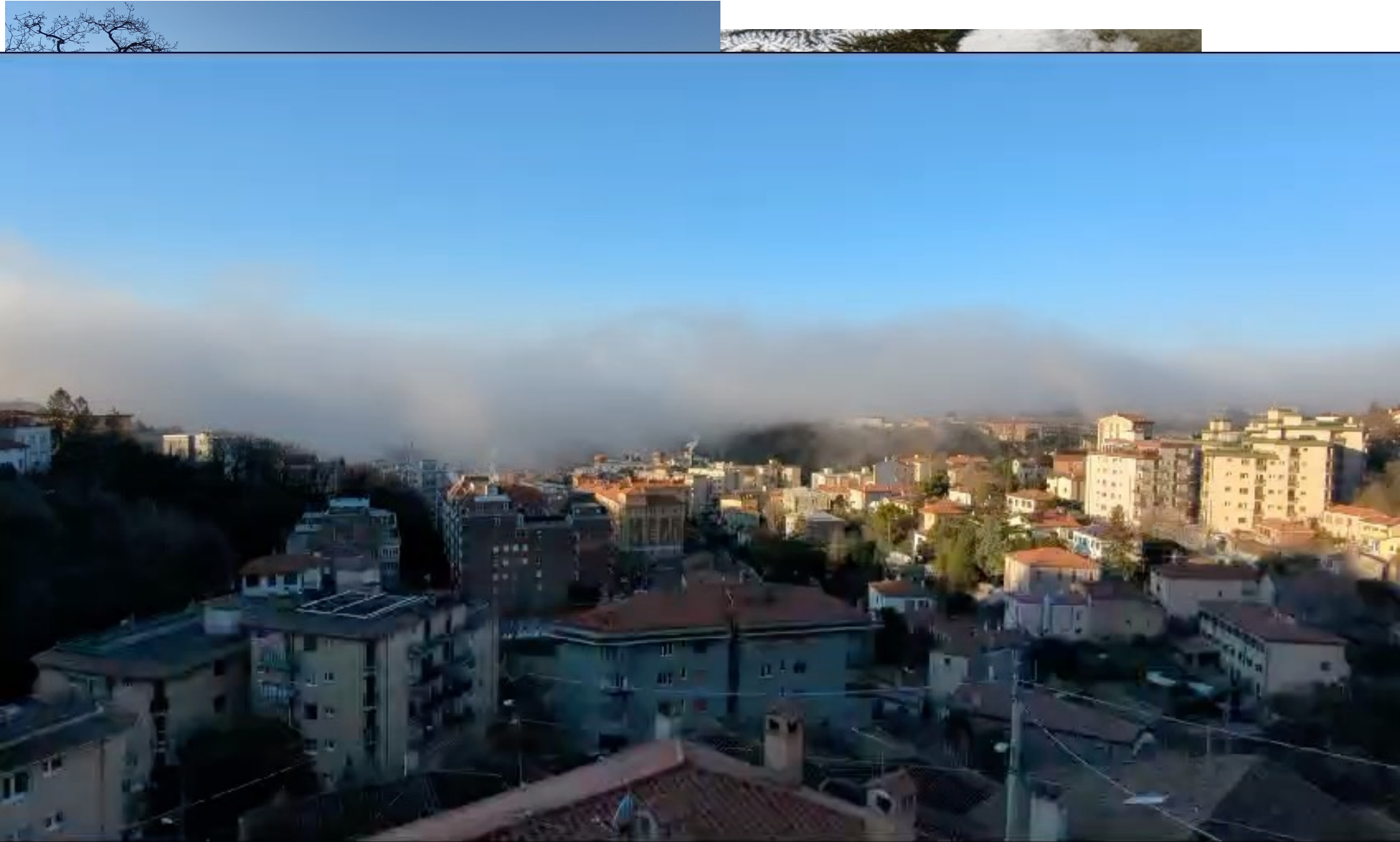
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HORIZONTAL BLOCKING: $Fr \ll 1$



Ridge acts as an obstacle in case of high stratification. Vertical displacements are strongly suppressed and air quality gets worse

HORIZONTAL BLOCKING: $Fr \ll 1$



RiPs act as an obstacle in case of high stratification vertical displacements are strongly suppressed and air quality gets worse

***WHAT controls
the vertical
motions?***

Rotation?

Stratification?

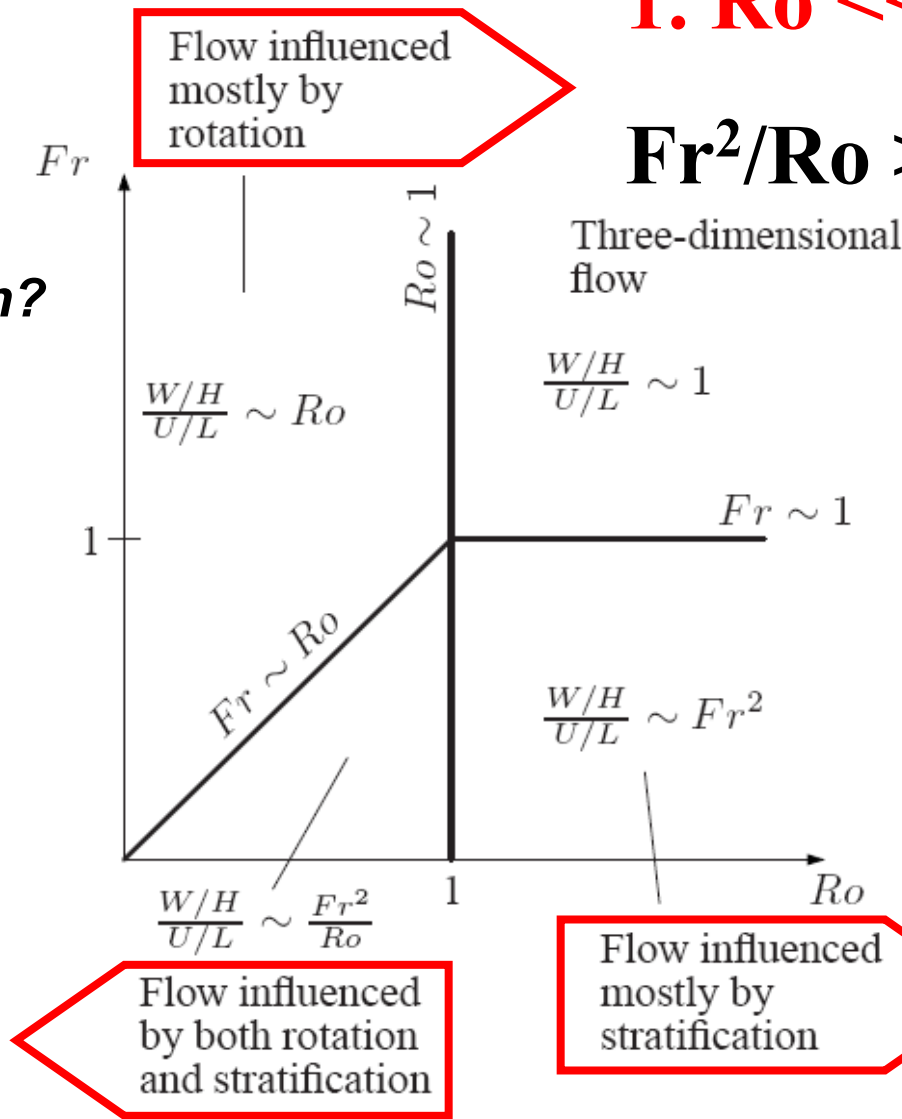
Both?

WHAT controls the vertical motions?

Rotation?

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Both?



1. $Ro \ll 1$

$Fr^2/Ro > Ro$

Figure 11-6 Recapitulation of the various scalings of the ratio of vertical convergence (divergence), W/H , to horizontal divergence (convergence), U/L , as a function of the Rossby number, $Ro = U/(\Omega L)$, and Froude number, $Fr = U/(NH)$.

2. $Fr \ll 1$

$Fr^2/Ro < Ro$

$$\frac{\text{vert. conv.}}{\text{horiz. div.}} = \frac{W/H}{U/L} = \begin{cases} Fr^2 \rightarrow \text{strat. dominates} \\ Fr^2/Ro, \rightarrow \text{strat. + rot.} \\ Ro \rightarrow \text{rot. dominates} \end{cases}$$

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Rotation?

Stratification?

Both?

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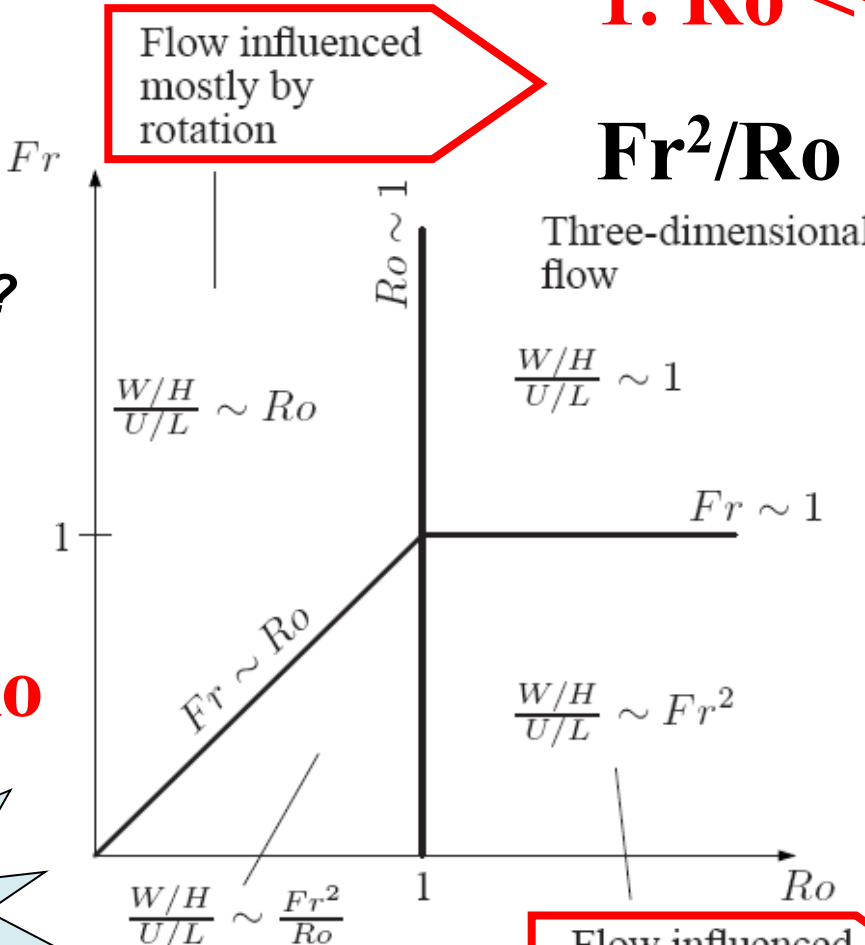


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3. $Fr^2 \lesssim Ro$

$L = NH/\Omega$

2. $Fr \ll 1$

$Fr^2/Ro < Ro$

$Bu = \left(\frac{NH}{\Omega L}\right)^2 = \left(\frac{Ro}{Fr}\right)^2$ **Burger n.: if $Bu=1 \Rightarrow L = NH/\Omega$**

2-layer stratified fluid + shear

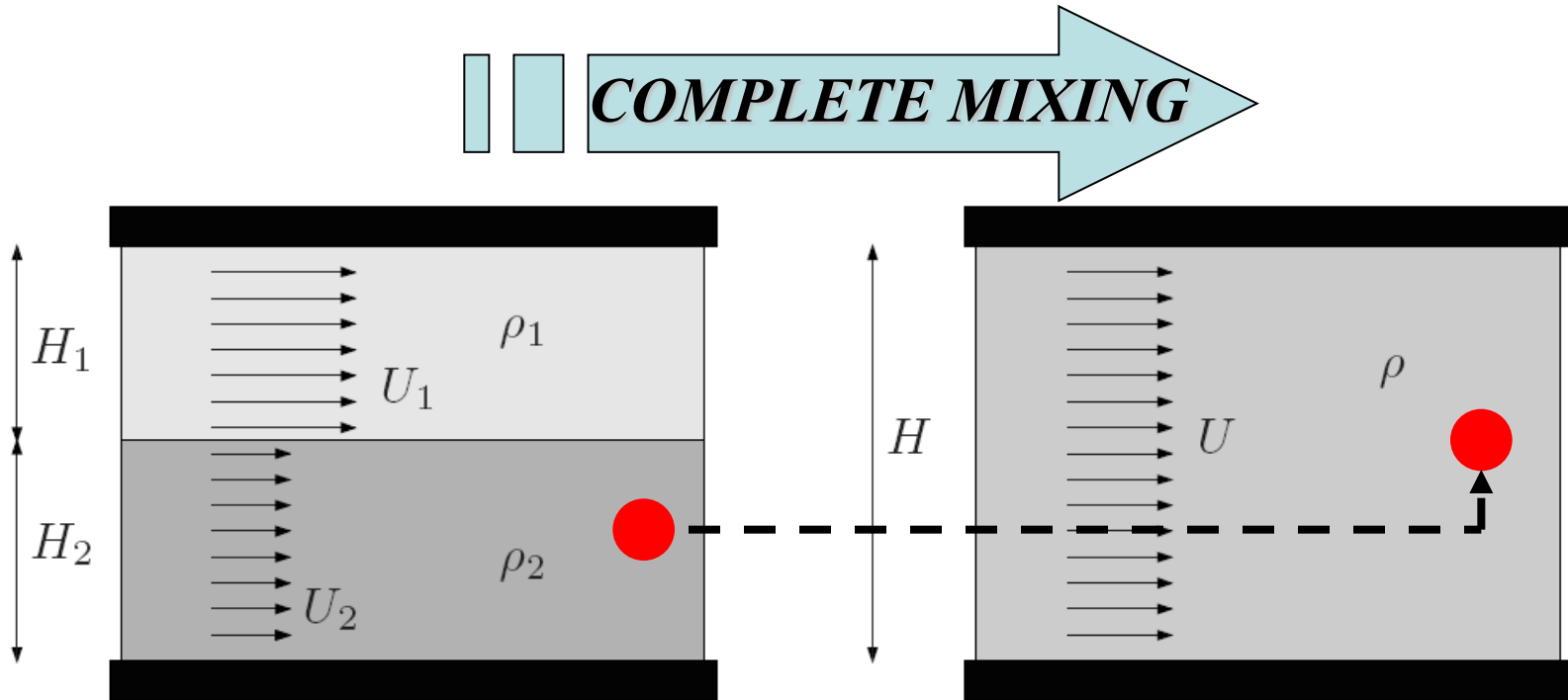


Figure 14-1 Mixing of a two-layer stratified fluid with velocity shear. Rising of dense fluid and lowering of light fluid both require work against buoyancy forces and thus lead to an increase in potential energy. Concomitantly, the kinetic energy of the system decreases during mixing. Only when the kinetic-energy drop exceeds the potential-energy rise can mixing proceed spontaneously.

● = center of gravity

2-layer stratified fluid + shear

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$$\begin{aligned} PE \text{ gain} &= \int_0^H \rho_{\text{final}} g z \, dz - \int_0^H \rho_{\text{initial}} g z \, dz \\ &= \frac{1}{2} \rho g H^2 - \left[\frac{1}{2} \rho_2 g \frac{H^2}{4} + \frac{1}{2} \rho_1 g \frac{3H^2}{4} \right] \\ &= \frac{1}{8} (\rho_2 - \rho_1) g H^2. \end{aligned}$$

$$\begin{aligned} KE \text{ loss} &= \int_0^H \frac{1}{2} \rho_0 u_{\text{initial}}^2 \, dz - \int_0^H \frac{1}{2} \rho_0 u_{\text{final}}^2 \, dz \\ &= \frac{1}{2} \rho_0 U_2^2 \frac{H}{2} + \frac{1}{2} \rho_0 U_1^2 \frac{H}{2} - \frac{1}{2} \rho_0 U^2 H \\ &= \frac{1}{8} \rho_0 (U_1 - U_2)^2 H. \end{aligned}$$

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COMPLETE VERTICAL MIXING is naturally possible when $KE_{\text{loss}} > PE_{\text{gain}}$:

$$\frac{(\rho_2 - \rho_1) g H}{\rho_0 (U_1 - U_2)^2} < 1$$

...meaning that the density variation barrier is not too strong or the shear is large enough to supply the necessary amount of energy to overcome the stratification ([...Ri](#))

In the opposite case ($KE_{\text{loss}} < PE_{\text{gain}}$) we have **LOCALIZED MIXING** occurring only near the initial interface, not extending to the entire system...

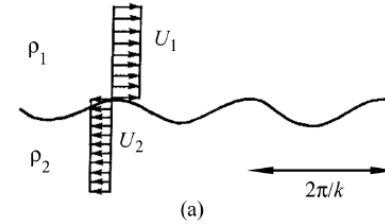
LOCALIZED MIXING

We know (Kundu, 1990) that a sinusoidal perturbation of wavenumber k is unstable if

$$(\rho_2^2 - \rho_1^2)g < \rho_1\rho_2k (U_1 - U_2)^2,$$

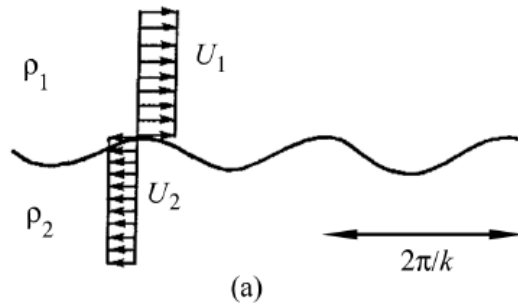
or for a Boussinesq fluid ($\rho_1 \simeq \rho_2 \simeq \rho_0$),

$$2(\rho_2 - \rho_1)g < \rho_0k (U_1 - U_2)^2.$$

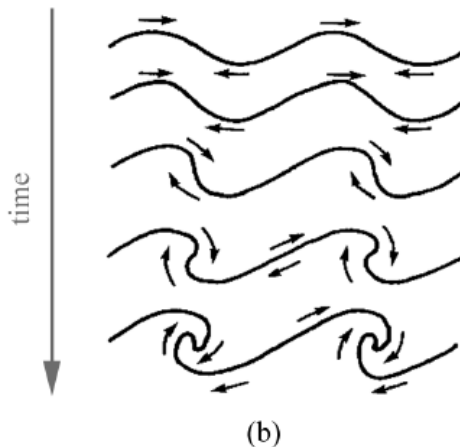


In a stability analysis, waves of all wavenumbers must be considered

$\Rightarrow \forall k \exists$ at least a \tilde{k} large enough to cause instability



a 2-layer shear flow is always unstable to short waves



(localized mixing)

Figure 14-2 Kelvin–Helmholtz instability: (a) initial perturbation of wavenumber k , (b) temporal evolution of an unstable perturbation. The system is always unstable to short waves, which steepen, overturn and ultimately cause mixing. As waves overturn, their vertical and lateral dimensions are comparable.

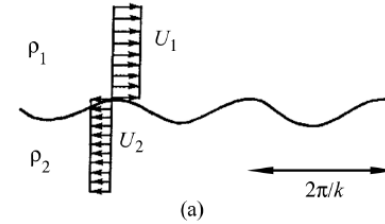
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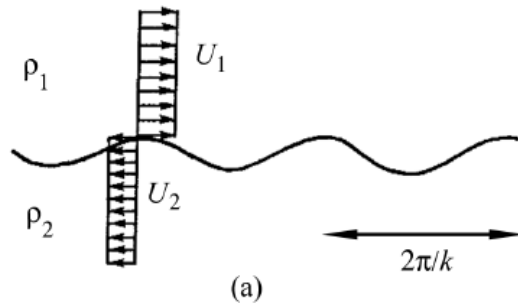
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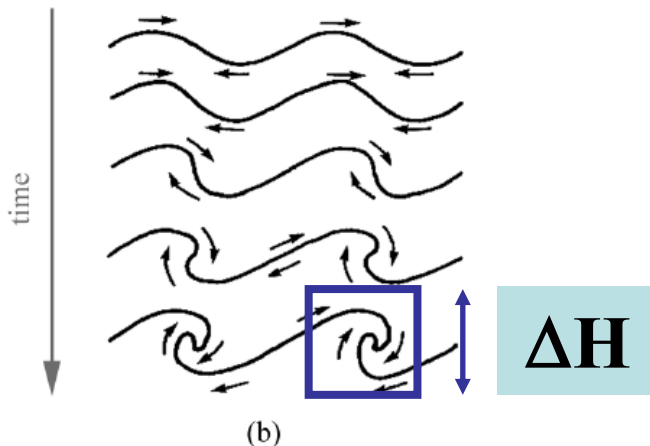


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Interfacial unstable waves grow and form ROLLS of height comparable to their width

$\Delta H = \text{mixing zone}$

$$\Delta H \propto \lambda_{max} = \frac{2\pi}{k_{min}} \sim \frac{\rho_0(U_1 - U_2)^2}{2(\rho_2 - \rho_1)g}$$

...link with H from $KE_{loss} < PE_{gain}$

$H > \Delta H$: *localized mixing*

$H < \Delta H$: *complete mixing*

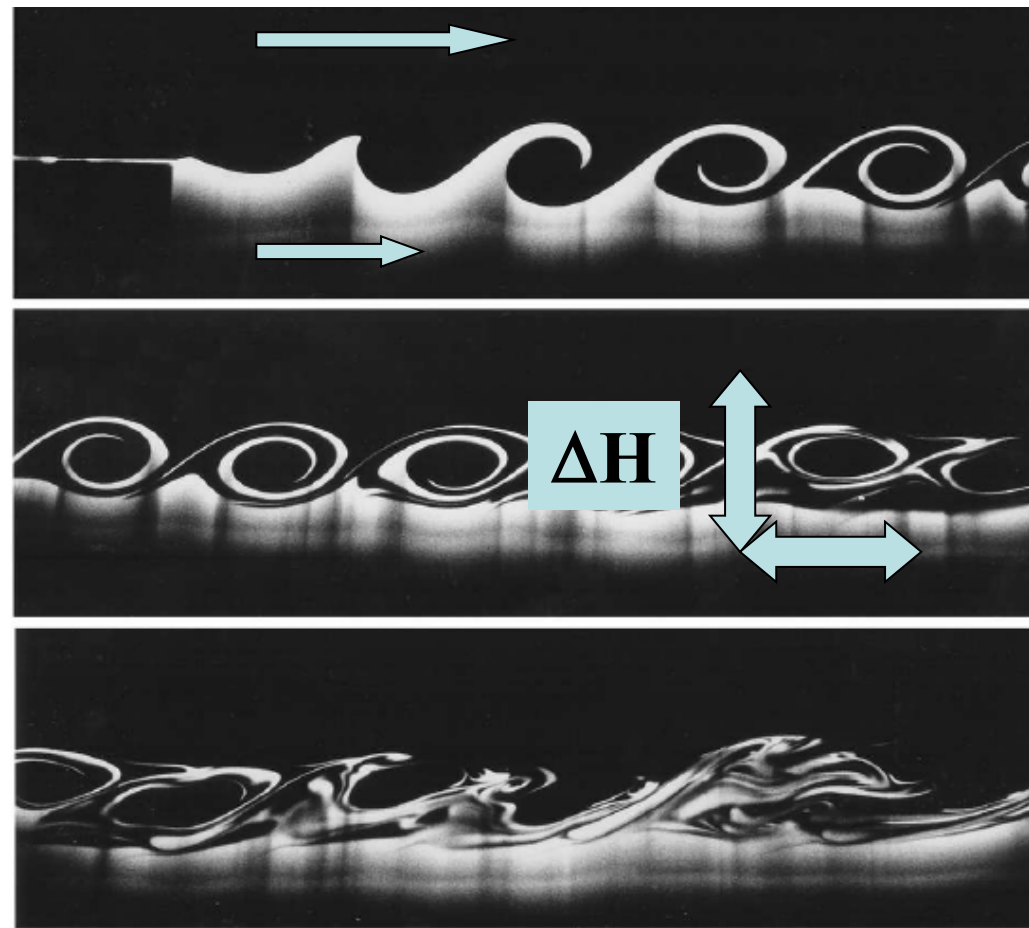


Figure 14-3 Development of a Kelvin–Helmholtz instability in the laboratory. Here, two layers flowing from left to right join downstream of a thin plate (visible on the left of the top photograph). The upper and faster moving layer is slightly less dense than the lower layer. Downstream distance (from left to right on each photograph and from top to bottom panel) plays the role of time. At first, waves form and overturn in a two-dimensional fashion (in the vertical plane of the photo) but, eventually, three-dimensional motions appear that lead to turbulence and complete the mixing. (Courtesy of Greg A. Lawrence. For more details on the laboratory experiment, see Lawrence *et al.*, 1991.)

ROLLING + BREAKING = TURBULENT MIXING

Kelvin-Helmoltz instability



Figure 14-5 Kelvin-Helmoltz instability in the Algerian sky. (Photo by the second author)

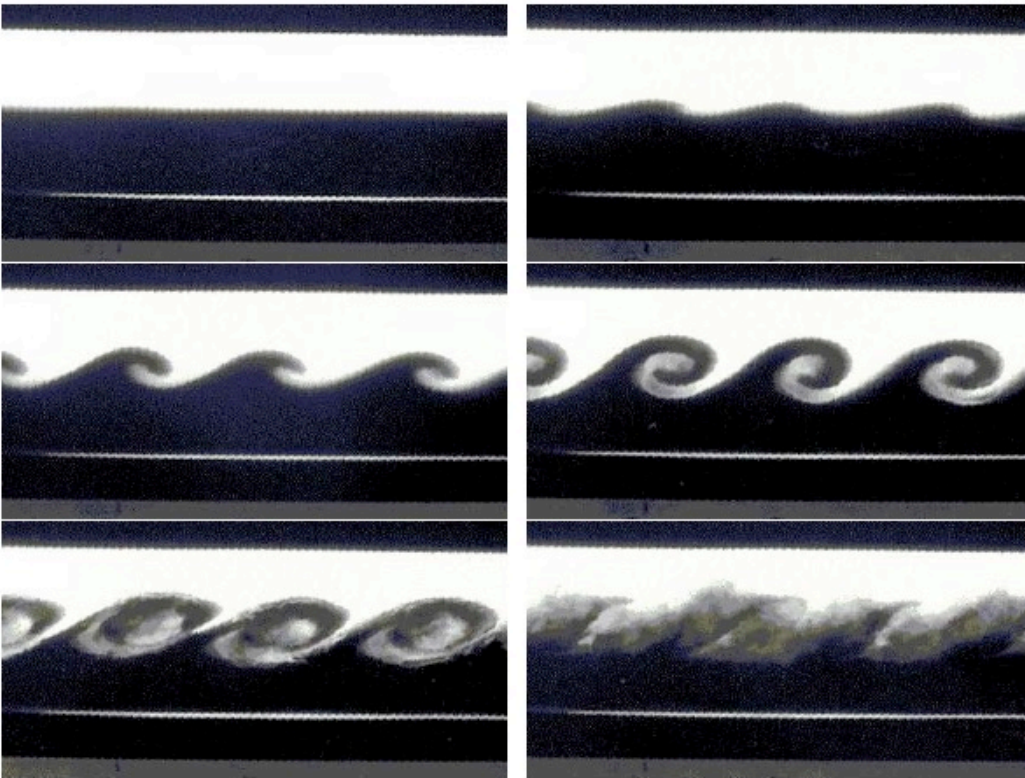


Figure 14-4 Kelvin-Helmoltz instability generated in a laboratory with fluids of two different densities and colours. (Adapted from GFD-online, Satoshi Sakai, Isawo Iizawa, Eiji Aramaki)

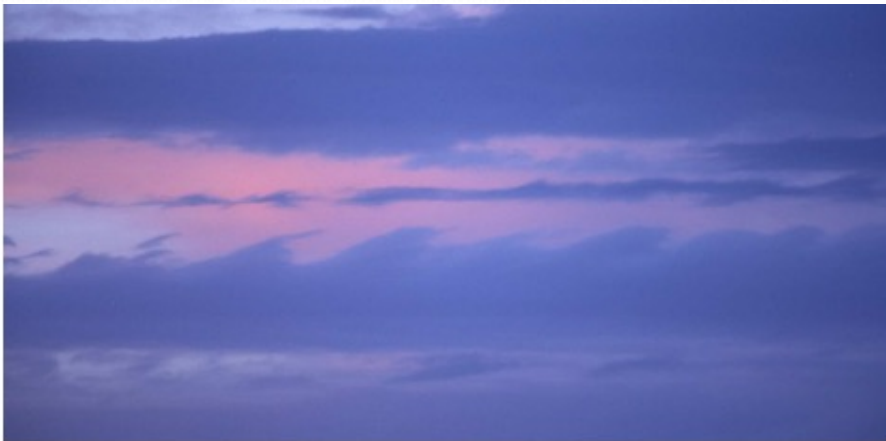


Figure 14-6 Kelvin-Helmoltz instability of the Sahara desert. (Photo by the second author)

Instability of a stratified shear flow

Q: For a given density stratification (N^2), what is the critical velocity shear for the instability \rightarrow mixing?

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Consider a 2-d (x,z) inviscid, non-rotating, non-diffusive fluid with velocity (u,w), dynamic pressure p and density anomaly ρ

$$\begin{aligned}\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + w \frac{\partial u}{\partial z} &= -\frac{1}{\rho_0} \frac{\partial p}{\partial x} \\ \frac{\partial w}{\partial t} + u \frac{\partial w}{\partial x} + w \frac{\partial w}{\partial z} &= -\frac{1}{\rho_0} \frac{\partial p}{\partial z} - \frac{\rho g}{\rho_0} \\ \frac{\partial u}{\partial x} + \frac{\partial w}{\partial z} &= 0 \\ \frac{\partial \rho}{\partial t} + u \frac{\partial \rho}{\partial x} + w \frac{\partial \rho}{\partial z} &= 0.\end{aligned}$$

Basic state + small Perturbation and Linearization

Our basic state consists of a steady, sheared horizontal flow [$u = \bar{u}(z)$, $w = 0$] in a vertical density stratification [$\rho = \bar{\rho}(z)$]. The accompanying pressure field $\bar{p}(z)$ obeys $d\bar{p}/dz = -g\bar{\rho}(z)$. The addition of an infinitesimally small perturbation ($u = \bar{u} + u'$, $w = w'$, $p = \bar{p} + p'$, $\rho = \bar{\rho} + \rho'$) and a subsequent **linearization** of the equations yield: $\mathbf{O}(2) \rightarrow \mathbf{0}$

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$$1) \quad \frac{\partial u'}{\partial t} + \bar{u} \frac{\partial u'}{\partial x} + w' \frac{d\bar{u}}{dz} = - \frac{1}{\rho_0} \frac{\partial p'}{\partial x}$$

$$2) \quad \frac{\partial w'}{\partial t} + \bar{u} \frac{\partial w'}{\partial x} = - \frac{1}{\rho_0} \frac{\partial p'}{\partial z} - \frac{\rho' g}{\rho_0}$$

$$3) \quad \frac{\partial u'}{\partial x} + \frac{\partial w'}{\partial z} = 0$$

$$4) \quad \frac{\partial \rho'}{\partial t} + \bar{u} \frac{\partial \rho'}{\partial x} + w' \frac{d\bar{\rho}}{dz} = 0.$$

def. Perturbation streamfunction ψ : $u' = +\partial\psi/\partial z$, $w' = -\partial\psi/\partial x$

Basic state + small Perturbation and Linearization

Hypotheses:

- a linear density vertical profile: $N^2 = -\frac{g}{\rho_0} \frac{d\rho}{dz} = \text{const}$
- the horizontal and temporal evolution of the perturbations $u' w' \rho' \psi'$ are formally **harmonic functions** in (x, t) , propagating with $\lambda_x = 2\pi/k$ and having a Fourier-like wave structure $\sim e^{ik(x-ct)}$...**BUT** $f(z) = ?$
- substituting in Eq. 4 we obtain the **density perturbation** $\rho' = \frac{-N^2 \rho_0}{g(\bar{u}-c)} \psi$
- [after some maths] we obtain the **Taylor-Goldstein equation** governing the vertical structure of a perturbation $\psi' = \psi(z) e^{il(x-ct)}$ in a stratified shear flow:

$$(\bar{u} - c) \left(\frac{d^2 \psi}{dz^2} - k^2 \psi \right) + \left(\frac{N^2}{\bar{u} - c} - \frac{d^2 \bar{u}}{dz^2} \right) \psi = 0$$

- with the **boundary conditions** $w'(0) = w'(H) = 0 \Rightarrow \psi(0) = \psi(H) = 0$ in a domain vertically bounded by two horizontal planes at $z = 0, H$ we obtain an **eigenvalue problem** which in general may have complex eigenvalues $c = c_r + ic_i$ and $c^* = c_r - ic_i$

Basic state + small Perturbation and Linearization

- from $c = c_r + ic_i$ and $c^* = c_r - ic_i \Rightarrow \psi' \sim e^{ik[x-(c_r \pm ic_i)t]} \sim e^{ikx} e^{-ikc_r t} e^{\mp kc_i t}$
- real exponential: the presence of $c_i \neq 0 \Rightarrow \exists$ at least one unstable mode
- the flow is **stable** $\Leftrightarrow c_i = 0$

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- the flow is **stable** $\Leftrightarrow c_i = 0$
- using integral constraints we can analyze the T.-G. eq.: with $\psi = \sqrt{\bar{u} - c} \phi$.

$$\frac{d}{dz} \left[(\bar{u} - c) \frac{d\phi}{dz} \right] - \left[k^2(\bar{u} - c) + \frac{1}{2} \frac{d^2 \bar{u}}{dz^2} + \frac{1}{\bar{u} - c} \left(\frac{1}{4} \left(\frac{d\bar{u}}{dz} \right)^2 - N^2 \right) \right] \phi = 0$$

$$\phi(0) = \phi(H) = 0.$$

- [after some math and using the BCs] we obtain a **complex equation**:

$$\int_0^H \left[N^2 - \frac{1}{4} \left(\frac{d\bar{u}}{dz} \right)^2 \right] \frac{|\phi|^2}{\bar{u} - c} dz = \int_0^H (\bar{u} - c) \left(\left| \frac{d\phi}{dz} \right|^2 + k^2 |\phi|^2 \right) dz + \frac{1}{2} \int_0^H \frac{d^2 \bar{u}}{dz^2} |\phi|^2 dz$$

- whose imaginary part is $c_i \int_0^H \left[N^2 - \frac{1}{4} \left(\frac{d\bar{u}}{dz} \right)^2 \right] \frac{|\phi|^2}{|\bar{u} - c|^2} dz = - c_i \int_0^H \left(\left| \frac{d\phi}{dz} \right|^2 + k^2 |\phi|^2 \right) dz,$

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- whose imaginary part is $c_i \int_0^H \left[N^2 - \frac{1}{4} \left(\frac{d\bar{u}}{dz} \right)^2 \right] \frac{|\phi|^2}{|\bar{u} - c|^2} dz = -c_i \int_0^H \left(\left| \frac{d\phi}{dz} \right|^2 + k^2 |\phi|^2 \right) dz,$

- IF $N^2 > \frac{1}{4} \left(\frac{d\bar{u}}{dz} \right)^2 \Rightarrow c_i \cdot (> 0) = -c_i \cdot (> 0) \Rightarrow c_i = 0 \Rightarrow$ **STABLE**
- IF $N^2 < \frac{1}{4} \left(\frac{d\bar{u}}{dz} \right)^2 \Rightarrow c_i \cdot (< 0) = -c_i \cdot (> 0) \Rightarrow \forall c_i \Rightarrow$ **STABLE or UNSTABLE**

Q: For a given density stratification (N^2), what is the critical velocity shear for the instability \rightarrow mixing?

- [from $c = c_r + ic_i$ and $c^* = c_r - ic_i \Rightarrow \psi' \sim e^{ik[x-(c_r \pm ic_i)t]} \sim e^{ikx} e^{-ikt} e^{\mp k c_i t}$]

- def. Richardson number $Ri = \frac{N^2}{\left(\frac{d\bar{u}}{dz}\right)^2} \rightarrow \begin{cases} \text{IF } Ri > \frac{1}{4} \Rightarrow \text{STABLE} \\ \text{IF } Ri < \frac{1}{4} \Rightarrow \text{STABLE or UNSTABLE} \end{cases}$

- sufficient condition for stability is $N^2 > \frac{1}{4} \left(\frac{d\bar{u}}{dz}\right)^2 : Ri > \frac{1}{4} \dots (\Rightarrow \text{STABLE})$

- necessary condition for instability is $N^2 < \frac{1}{4} \left(\frac{d\bar{u}}{dz}\right)^2 : Ri < \frac{1}{4} \dots (\Leftarrow \text{UNSTABLE})$

- But measurements in atmosphere / ocean / laboratory indicate that $Ri < 1/4$ is a reliable condition of instability $\Rightarrow \left(\frac{d\bar{u}}{dz}\right)^2 > 4N^2$

- if the shear flow has linear variations of velocity and density we may refer to the 2-layer flow case, finding a similarity with $KE_{loss} > PE_{gain}$ (complete vertical mixing) $\Rightarrow Ri$ is the ratio between PE and KE !!!

- $Ri = \frac{\text{potential energy barrier that mixing - if occurring - must overcome}}{\text{kinetic energy available in the shear flow}}$

Q: can we say something more about the properties of the growing perturbation?

- if we introduce the **vertical displacement a** caused by the small wave perturbation: $\frac{\partial a}{\partial t} + \bar{u} \frac{\partial a}{\partial x} = w$ which corresponds to $(\bar{u} - c) a = -\psi$.
 $a \sim a(z) e^{ik(x-ct)}$ $w' = -\partial\psi/\partial x$
- we can rewrite the T.-G. eq. obtaining an eigenvalue problem for a :

$$\frac{d}{dz} \left[(\bar{u} - c)^2 \frac{da}{dz} \right] + [N^2 - k^2(\bar{u} - c)^2] a = 0$$

$$a(0) = a(H) = 0.$$

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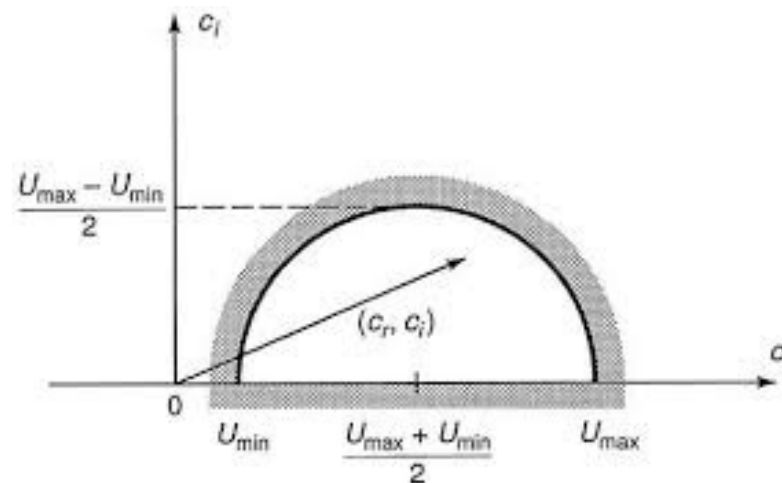
$$\frac{d}{dz} \left[(\bar{u} - c)^2 \frac{da}{dz} \right] + [N^2 - k^2(\bar{u} - c)^2] a = 0$$

$$a(0) = a(H) = 0.$$

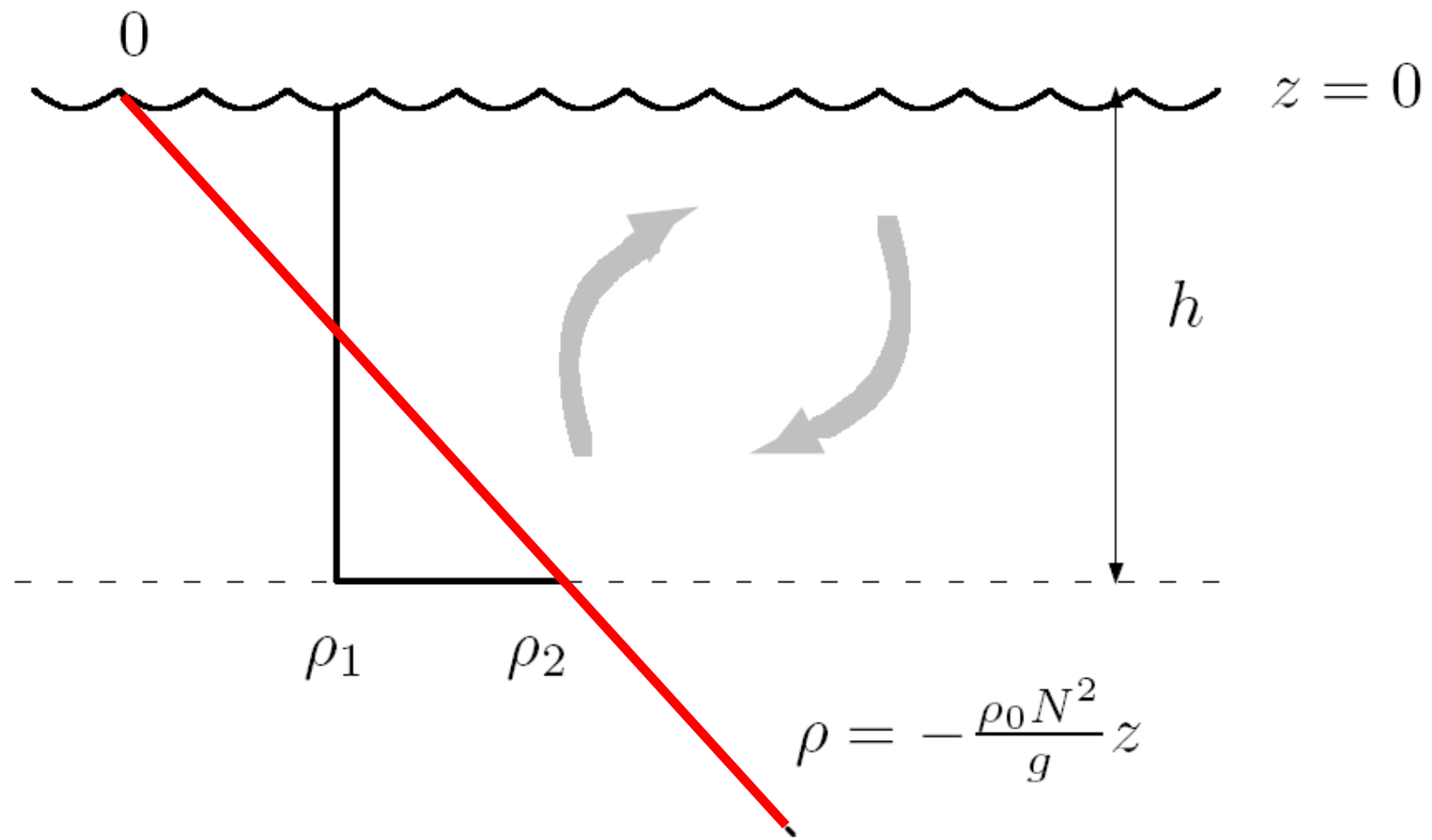
- [after some math and using the BCs]: $\int_0^H (\bar{u} - c)^2 P dz = \int_0^H N^2 |a|^2 dz$
- through the analysis of the real and the imaginary parts of the integral and requiring the instability condition $c_i \neq 0$:
 - $U_{\min} < c_r < U_{\max}$ the growing perturbation travels with the flow at some intermediate speed
 - $\left(c_r - \frac{U_{\min} + U_{\max}}{2} \right)^2 + c_i^2 \leq \left(\frac{U_{\max} - U_{\min}}{2} \right)^2$ in the complex plane, $c = c_r + ic_i$ must lie within the circle with the range of \bar{u} as the diameter on the real axis

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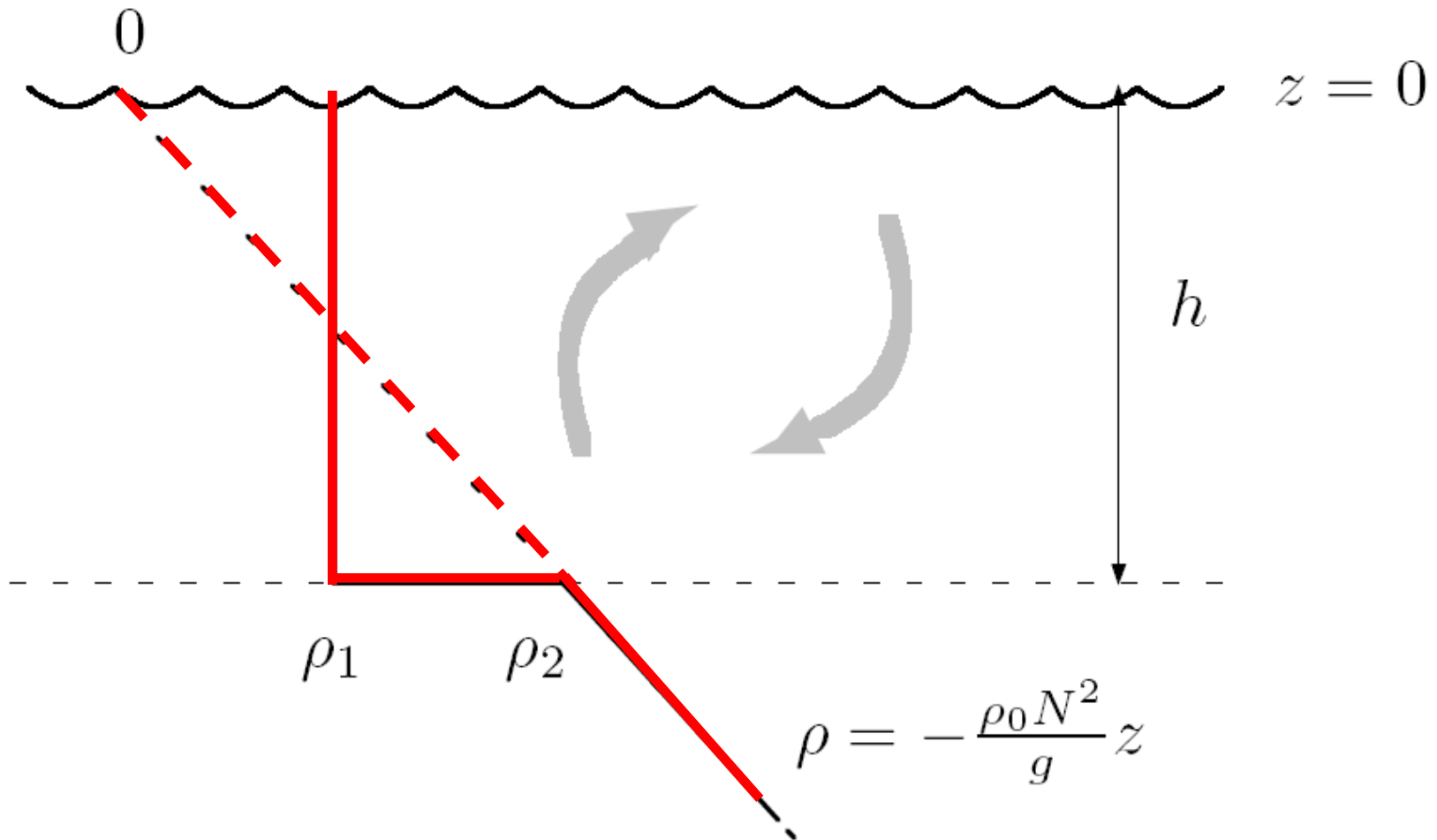
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 - $U_{\min} < c_r < U_{\max}$ the growing perturbation travels with the flow at some intermediate speed
 - $\left(c_r - \frac{U_{\min} + U_{\max}}{2}\right)^2 + c_i^2 \leq \left(\frac{U_{\max} - U_{\min}}{2}\right)^2$ in the complex plane (c_r, c_i) : $c = c_r + ic_i$ must lie within the circle with the range of \bar{u} as the diameter on the real axis
- since instability requires $c_i > 0 \Rightarrow \psi' \sim e^{ikx} e^{-ikt} e^{kc_i t}$ we are interested in the upper part of the circle \Rightarrow Howard semicircle theorem
- $c_i \leq \frac{U_{\max} - U_{\min}}{2} \Rightarrow kc_i \leq \frac{k}{2}(U_{\max} - U_{\min}) \Rightarrow$ the perturbation does not grow to infinite



Wind stress: $\tau = \rho_0 u_*^2$

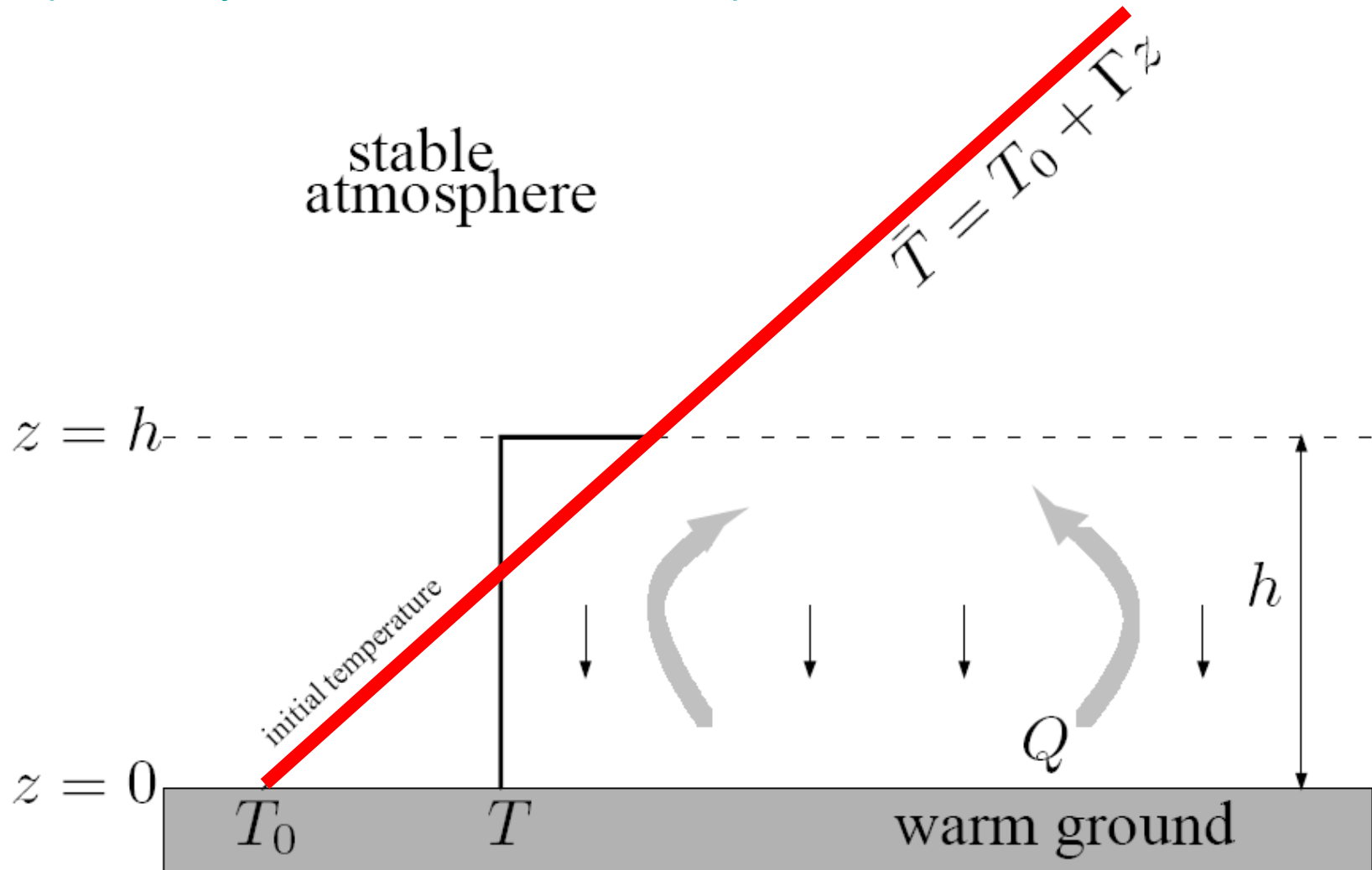


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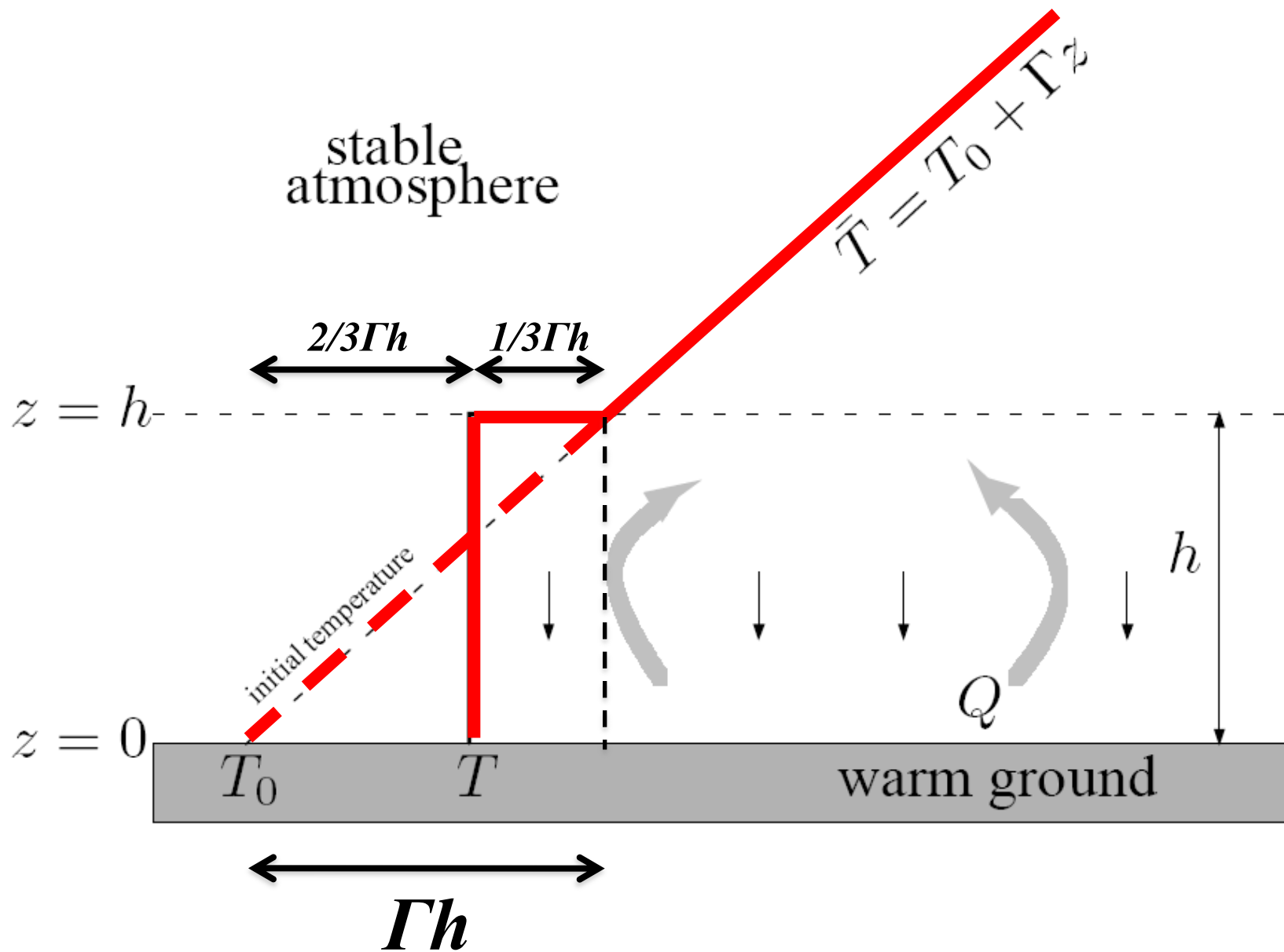
Rayleigh Benard Thermal Convection

<https://www.youtube.com/watch?v=buskqZIPdvl>



<https://www.youtube.com/watch?v=aLvK2cbsvzM>

<https://www.youtube.com/watch?v=0YatiDf9A8A>



THANKS

A dramatic sky scene featuring a sunburst breaking through a layer of clouds. The sun is positioned centrally, creating a bright glow and a lens flare effect. The clouds are dark and textured, with some lighter patches where the sun's light hits. The overall color palette is dominated by deep blues, greys, and whites. The word "THANKS" is overlaid in a large, bold, white sans-serif font across the upper portion of the image.