

**Planck unveils the Cosmic Microwave Background**



planck

# L'eco del BigBang misurato da un satellite Italo/Francese, PLANCK

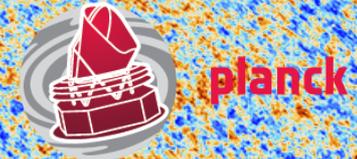
A. Zacchei

INAF-OATs

on behalf of the Planck Collaboration

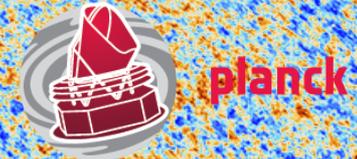


# Overview

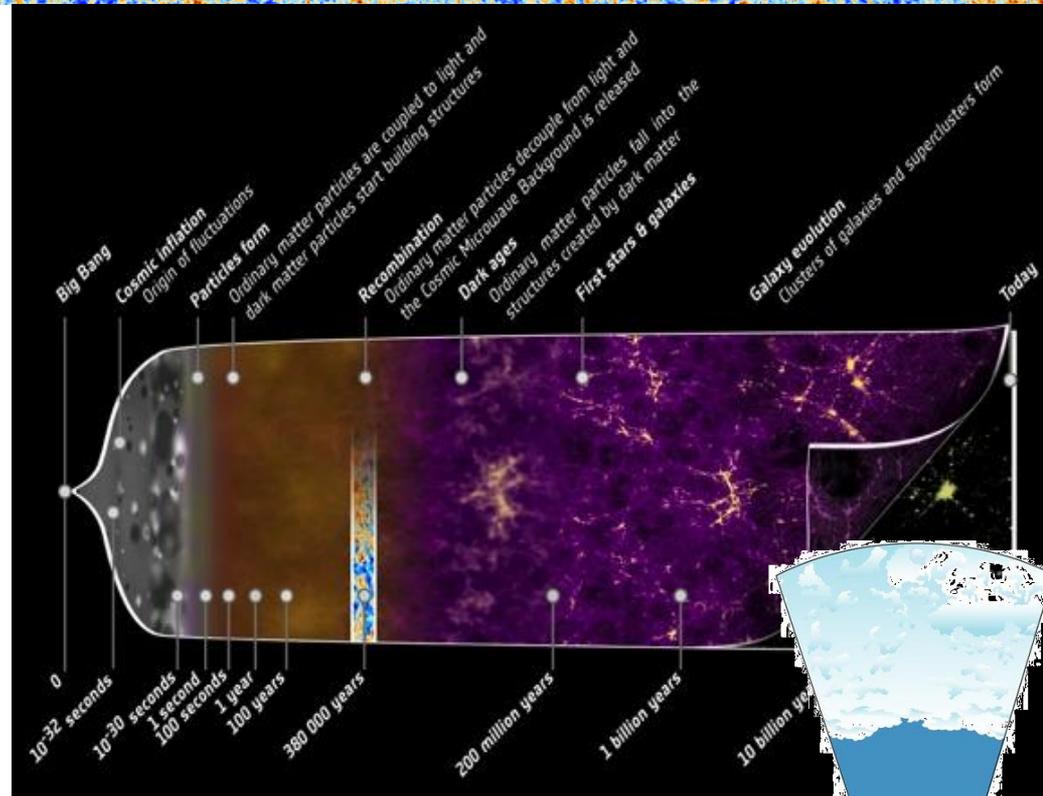


1. La CMB
2. Il Satellite Planck
3. L'acquisizione ed analisi dei dati
4. I primi risultati pubblici

# Il Big Bang

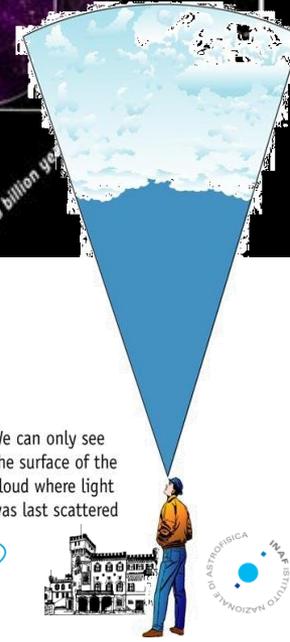


- Universo Primordiale: opaco, plasma caldo e ionizzato ad altissima densita'
  - Fotoni interagenti, elettroni, quarks, il tutto in equilibrio termico
- Universo in espansione
  - Scattering di Thomson
  - Il plasma si raffredda formando atomi di Idrogeno

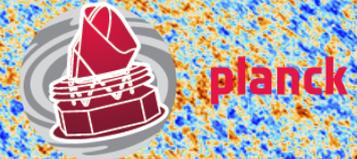


- $T \approx 3000 \text{ K}$  ricombinazione (disaccoppiamento)
  - $t \approx 380000$  anni dopo il Big Bang
- La materia diventa trasparente alla radiazione, la luce e' libera di viaggiare...

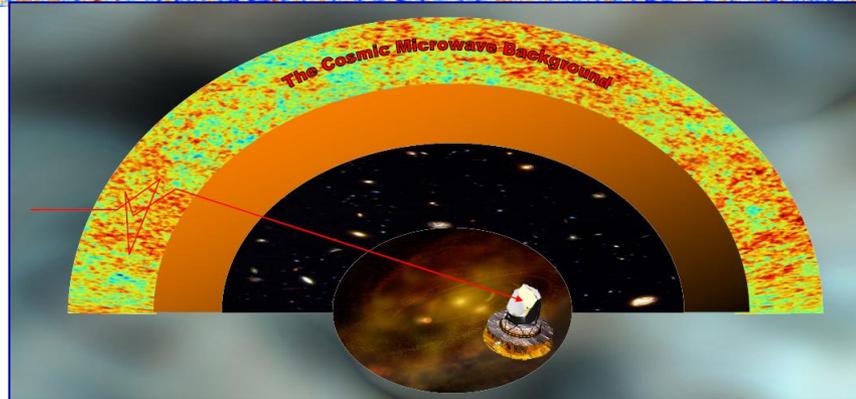
We can only see the surface of the cloud where light was last scattered



# Uno sguardo nel passato... da 3000 a 3 k



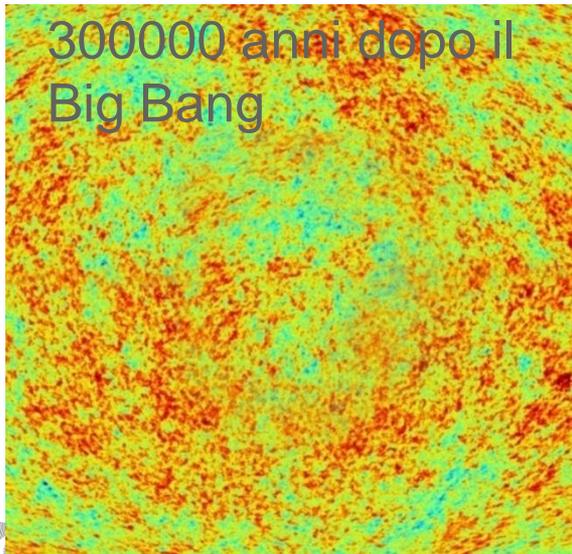
I primi 3 minuti



Strutture a piccola scala: risoluzione

Necessita' di rilevare radiazioni molto deboli: sensibilita'

300000 anni dopo il Big Bang



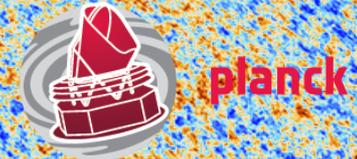
$5 \times 10^9$  anni dal Big Bang



$1 \times 10^{10}$  anni dal Big Bang



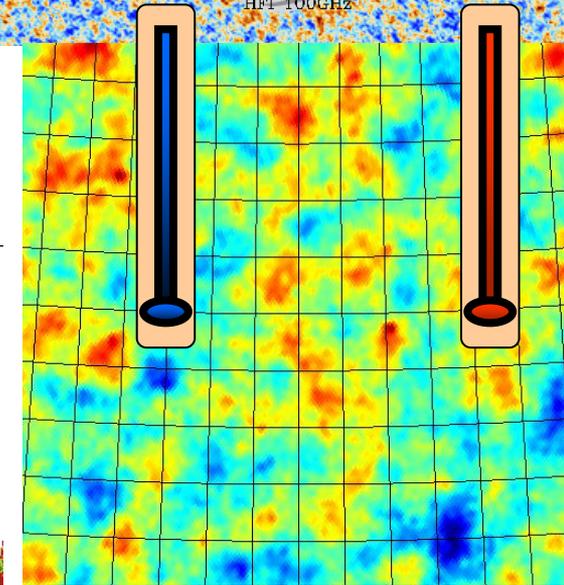
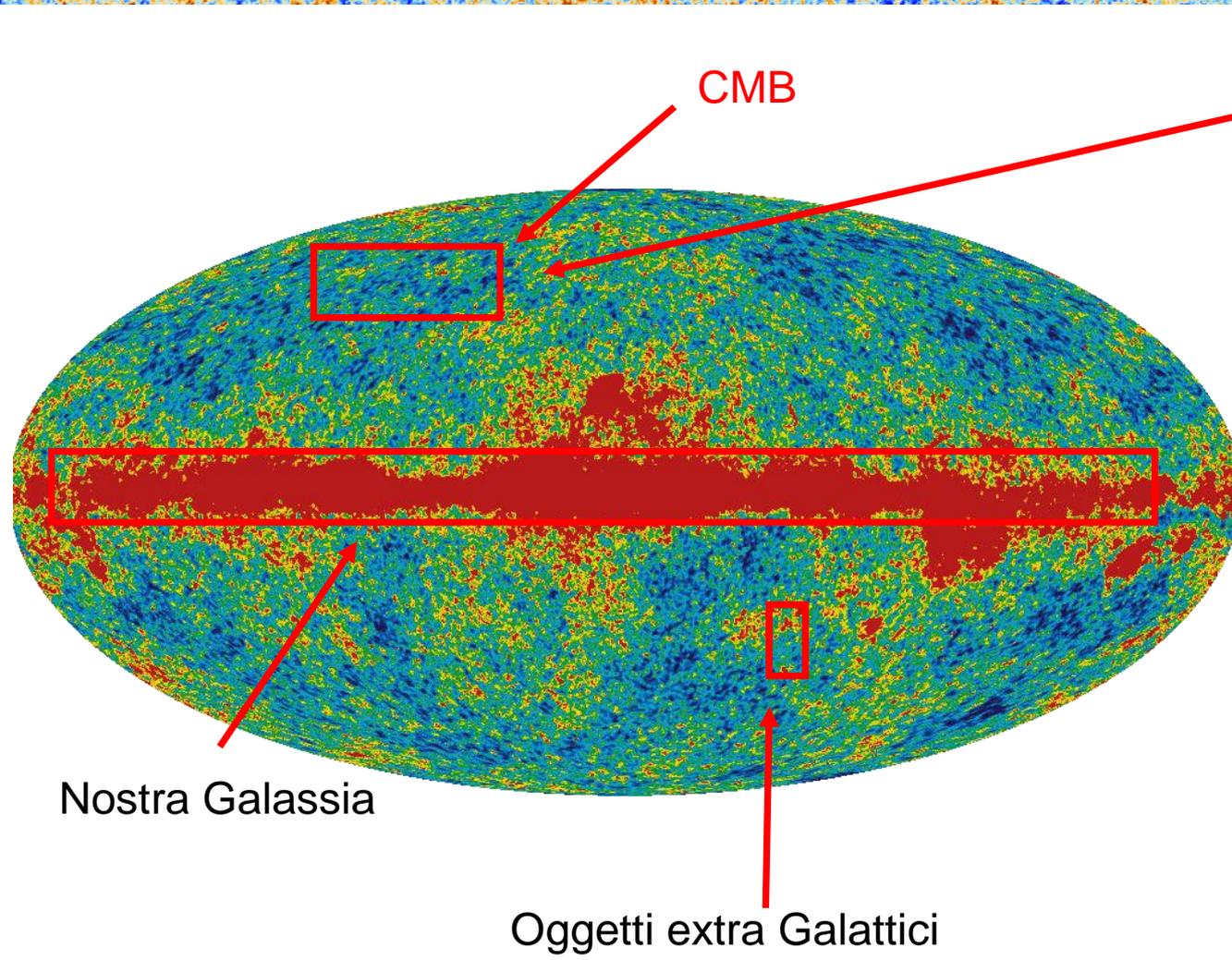
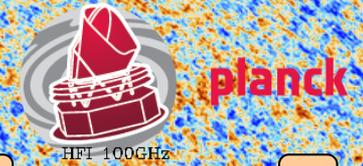
# The CMB



Abbiamo una formula per capire il nostro Universo;

1. L'Universo, subito dopo il big bang, e' un oggetto piuttosto semplice (assomiglia ad un corpo nero perfetto);
2. La CMB (Cosmological Microwave Background) ci da' l'immagine dell'Universo al tempo della ricombinazione
3. Proprieta' dell'Universo a larga scala non sono cambiate dall'epoca della ricombinazione e possono essere sintetizzate in pochi numeri;
4. Usando la CMB questi numeri descrittivi del nostro universo possono essere misurati con elevata accuratezza.

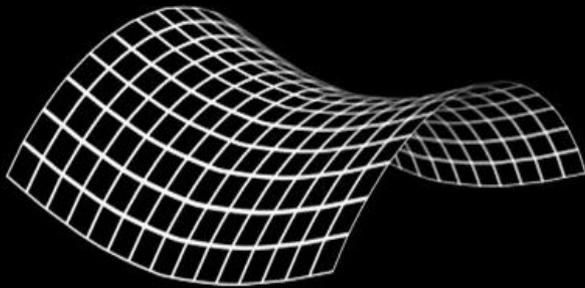
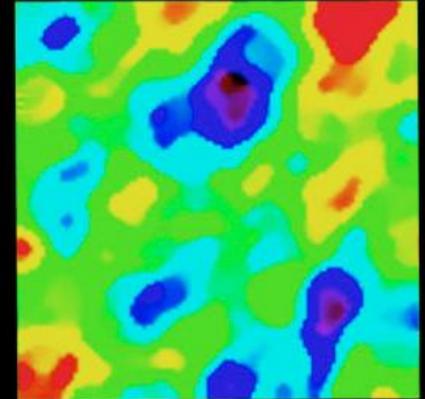
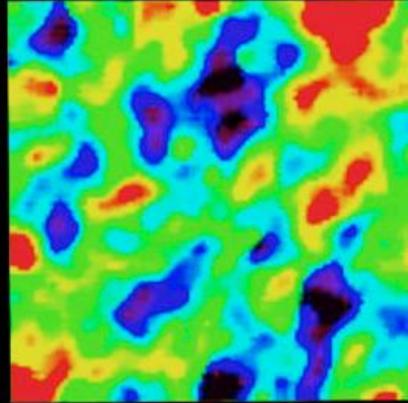
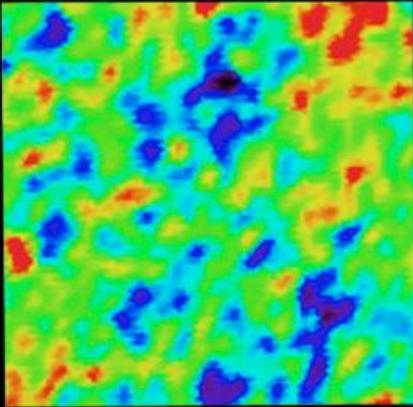
# Le anisotropie della CMB



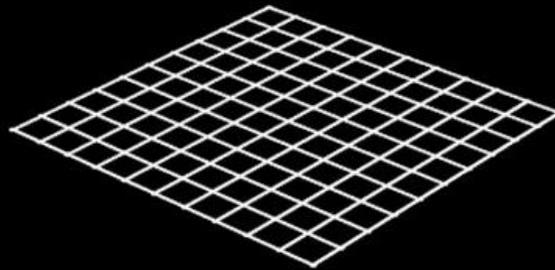
70  $\mu\text{K}$



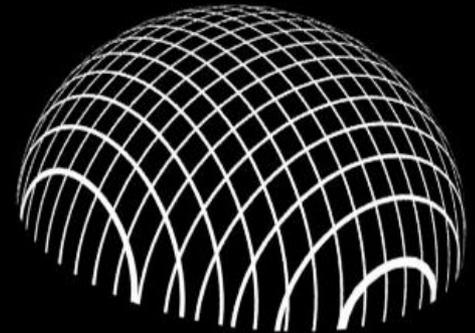
# La CMB e la Geometria dell'Universo



Open

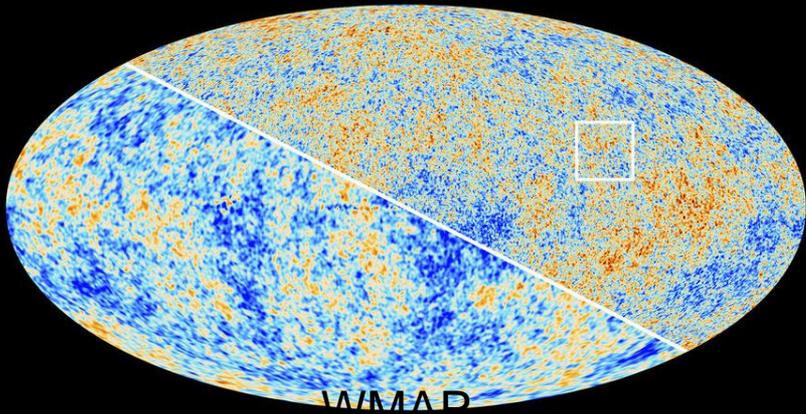


Flat

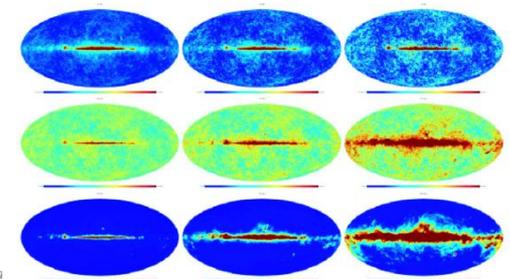
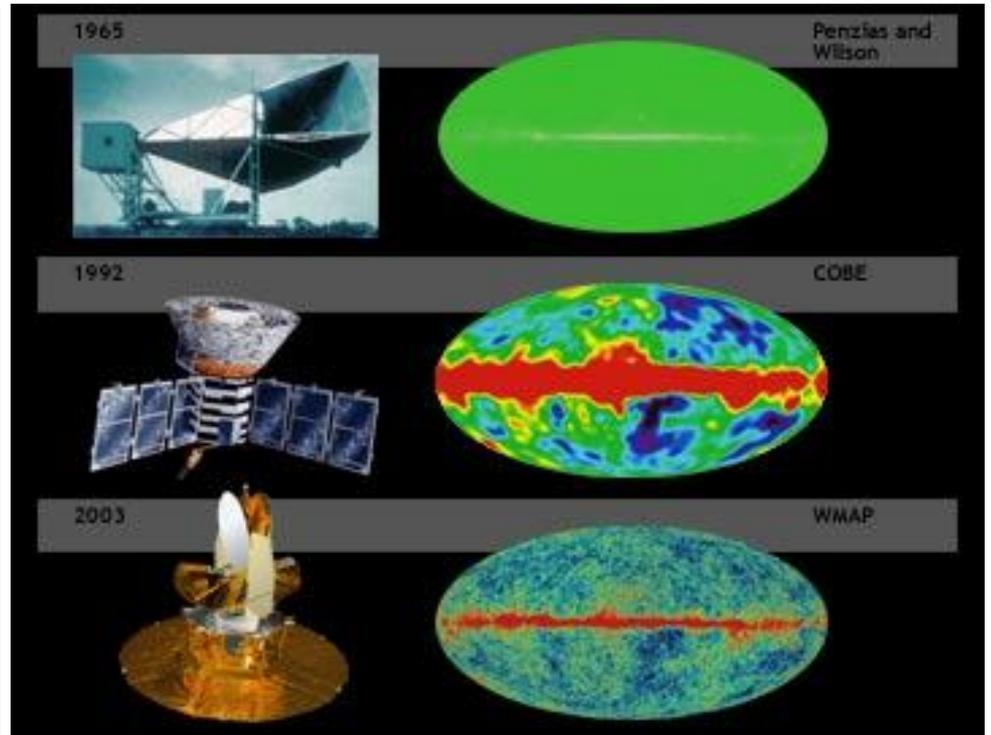
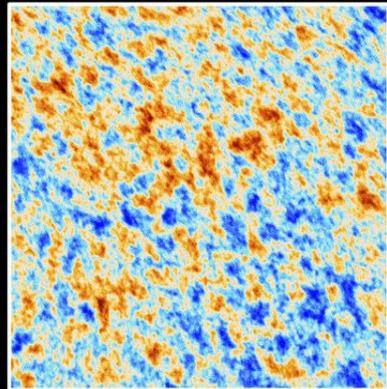
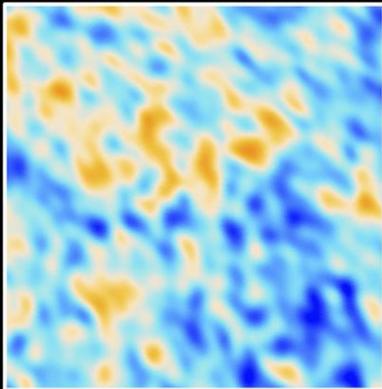


Closed

# Planck vs WMAP

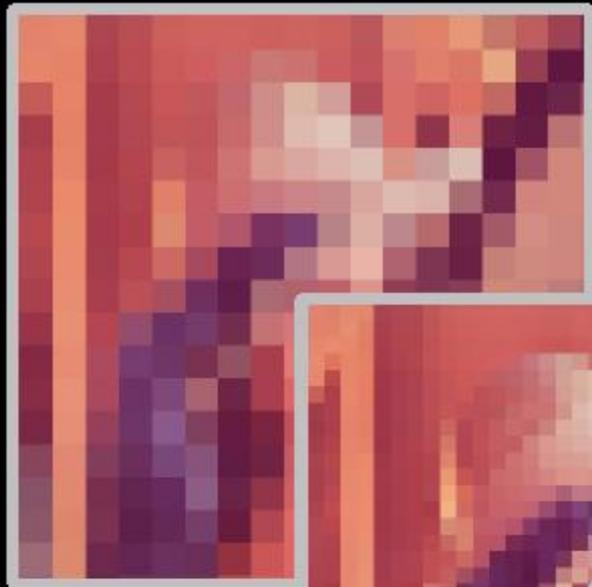
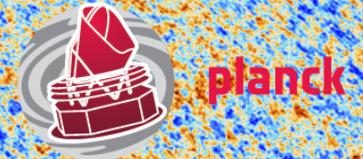


WMAP

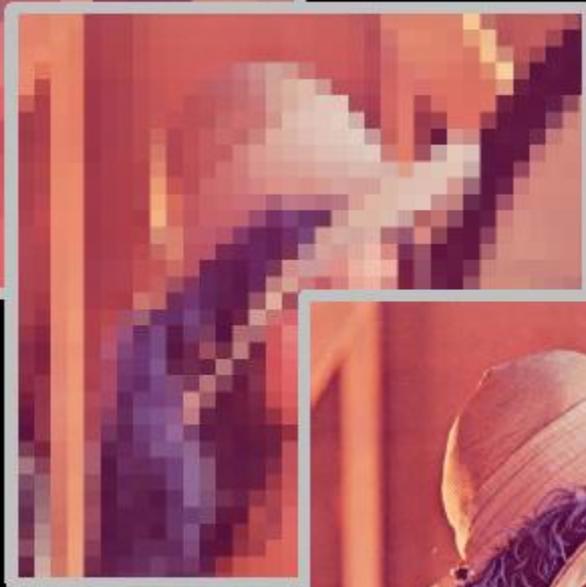


April 16th 2013

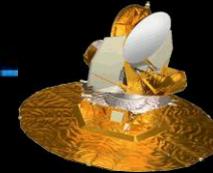
# Why Planck?



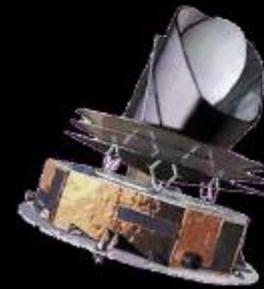
COBE - 1989



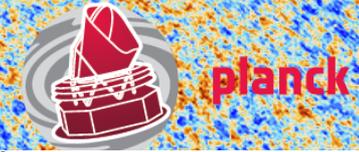
WMAP -



PLANCK

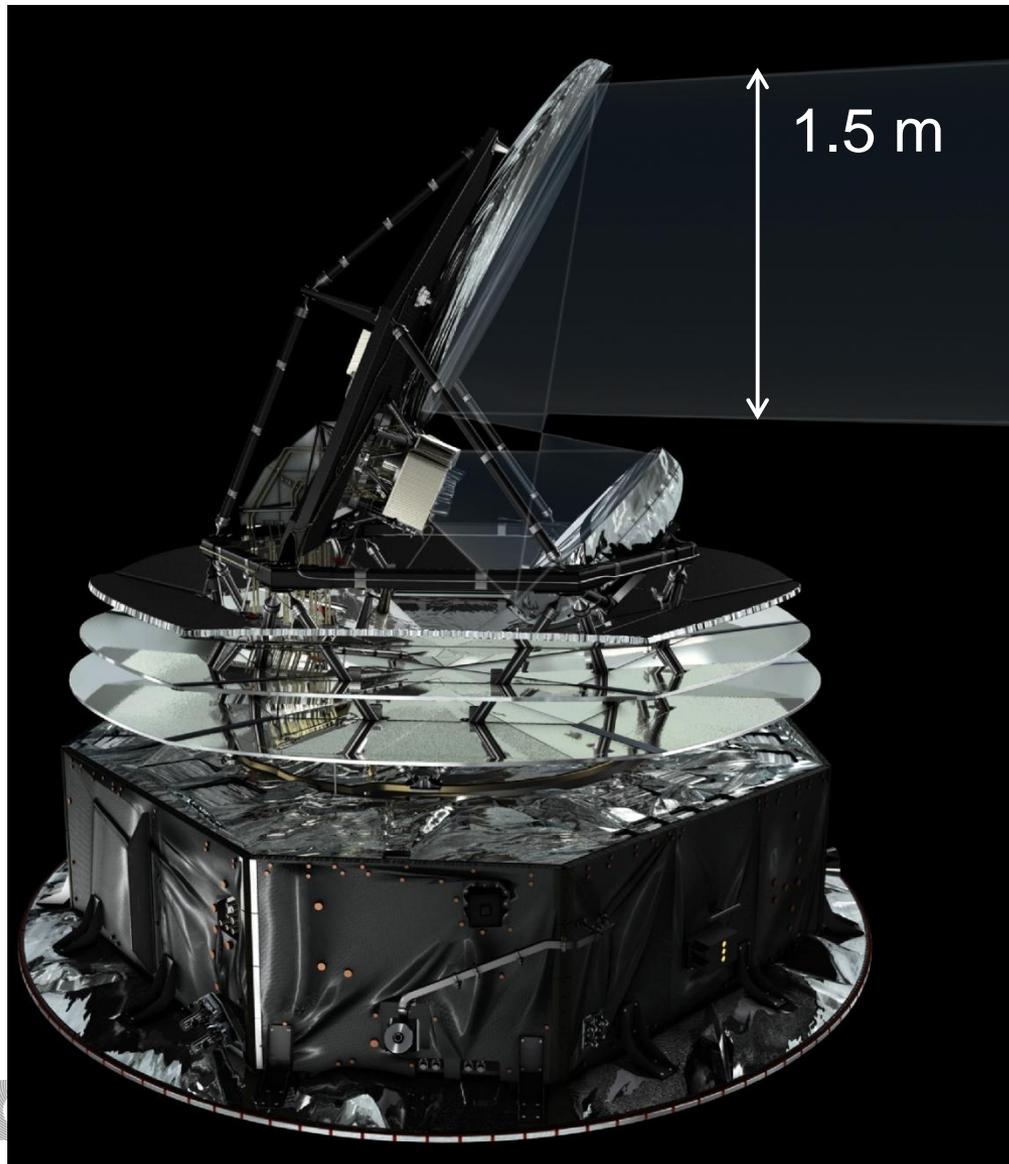
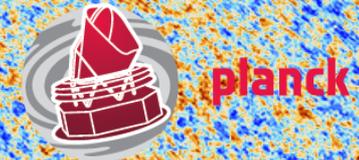


# Planck in pillole



1. Missione ESA, è stata lanciata il 14 Maggio 2009 (con Herschel).
2. Oltre a ESA, ci sono due Consorzi internazionali ognuno con la responsabilità della costruzione degli strumenti e l'analisi dei dati
  - a. LFI, 22 radiometri raffreddati a 20 K, PI: N.Mandolesi - INAF/IASF Bologna.
  - b. HFI, matrice di bolometri raffreddati a 0.1 K, PI: J.L.Puget - IAS Orsay.
  - c. Sistema di Raffreddamento (Sorption Cooler), PI: C. Lawrence - NASA/JPL.
3. La vita nominale (14 mesi dopo il raggiungimento del punto L2) è stata estesa di tre anni  $\Rightarrow$  cinque survey complete del cielo; tre survey "LFI-only". Durata complessiva: 51 mesi.
4. Costo COMPLESSIVO della missione circa 800 milioni di Euro.
5. Scopo scientifico: ricavare le mappe a tutto cielo di tutte le maggiori sorgenti che emettono dalle microonde sino al lontano infrarosso, mappe delle anisotropie nella RADIAZIONE DI FONDO COSMICO  $\Rightarrow$  **valutare i principali parametri cosmologici.**

# Il satellite Planck



3<sup>rd</sup> CMB space mission - 1<sup>st</sup>  
ESA in collaboration with  
European, US and Canadian  
scientific community

Mass 2'000 kg

Power 1'600 W

Size 4.2 × 4.2 m

Cost 800×10<sup>6</sup> €

50'000 Electronic components

36'000 I <sup>4</sup>He

12'000 I <sup>3</sup>He

11'400 Documents

20 yrs between project & results

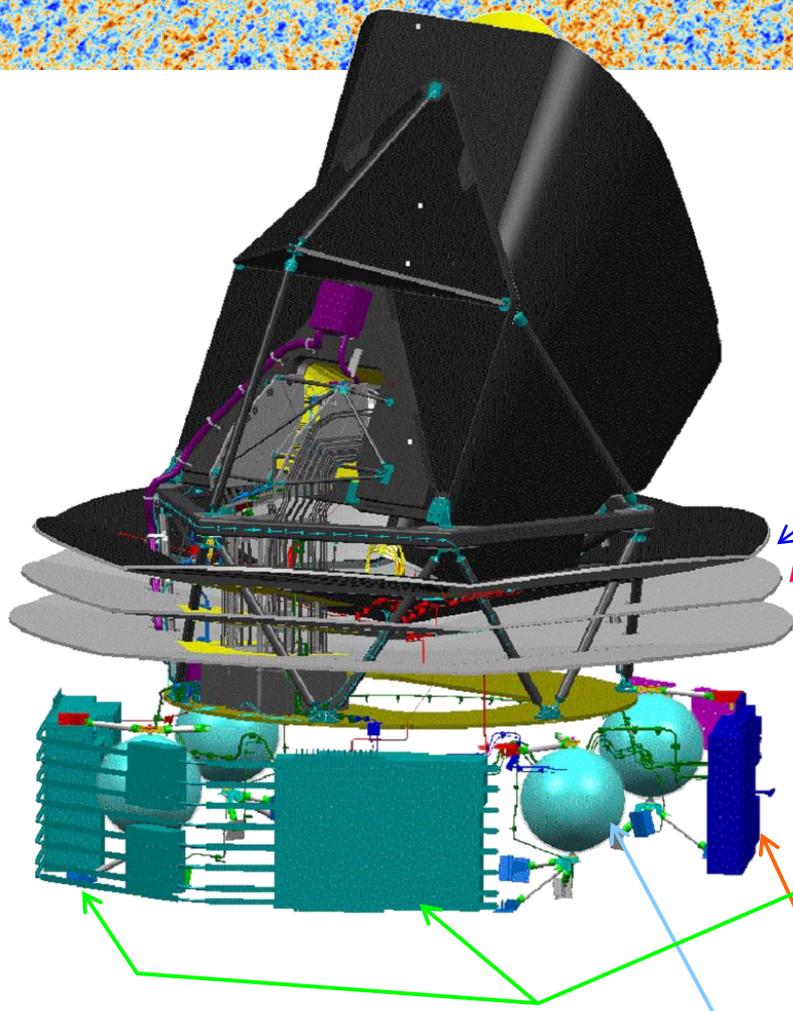
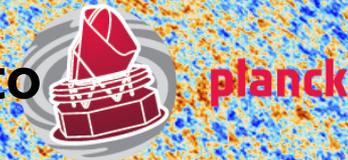
2 instruments & consortia

16 countries

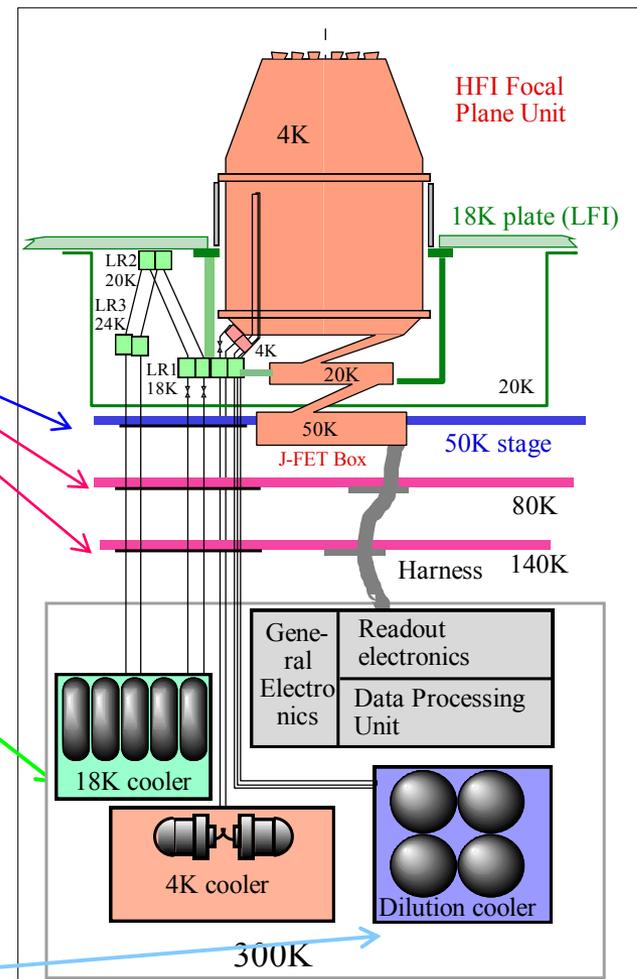
400 researchers



# Il frigorifero .... L'oggetto piu' freddo mai costruito



V-grooves

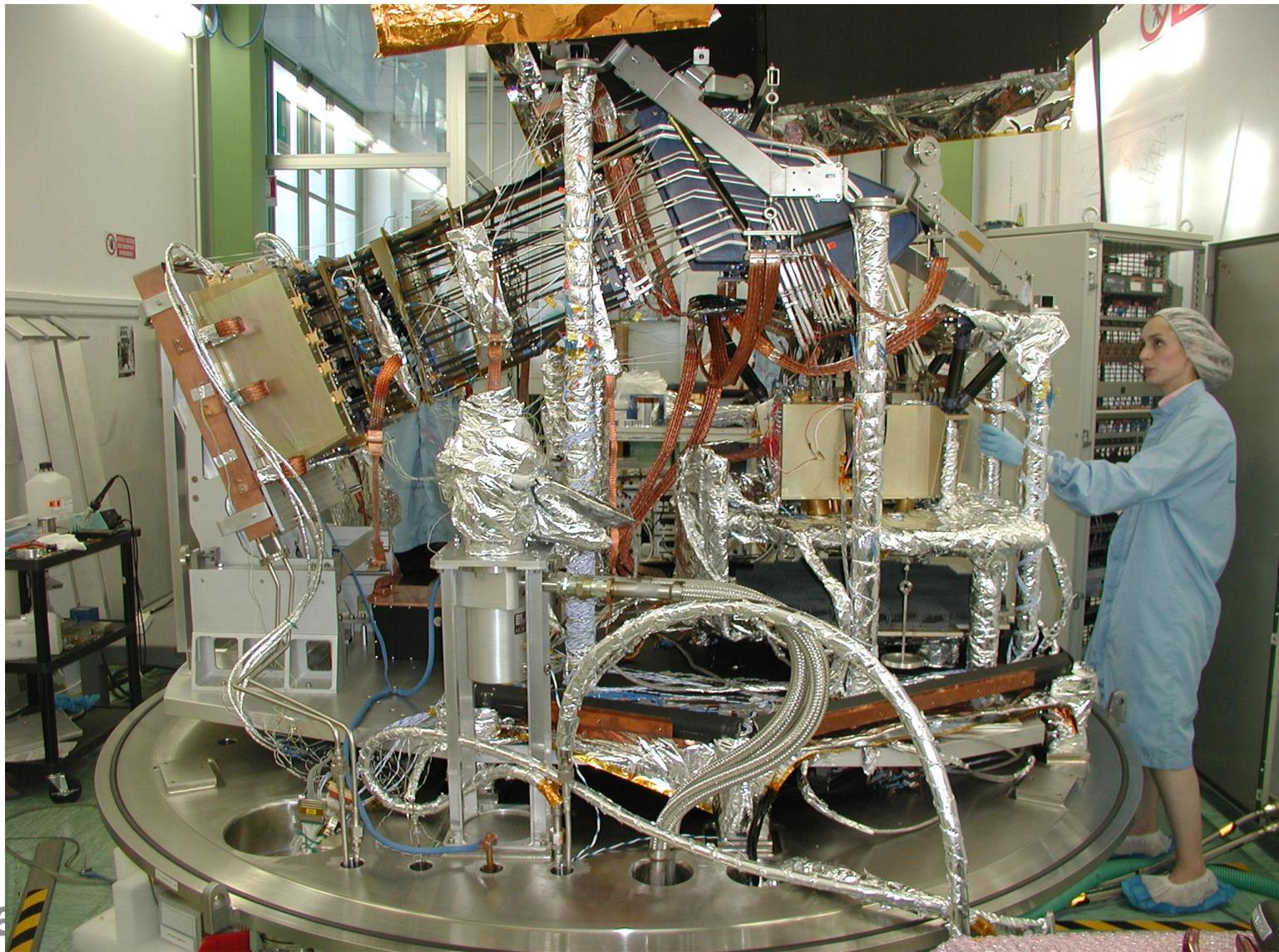
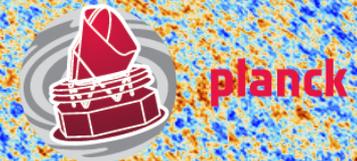


18-20 K H<sub>2</sub> sorption coolers (JPL)

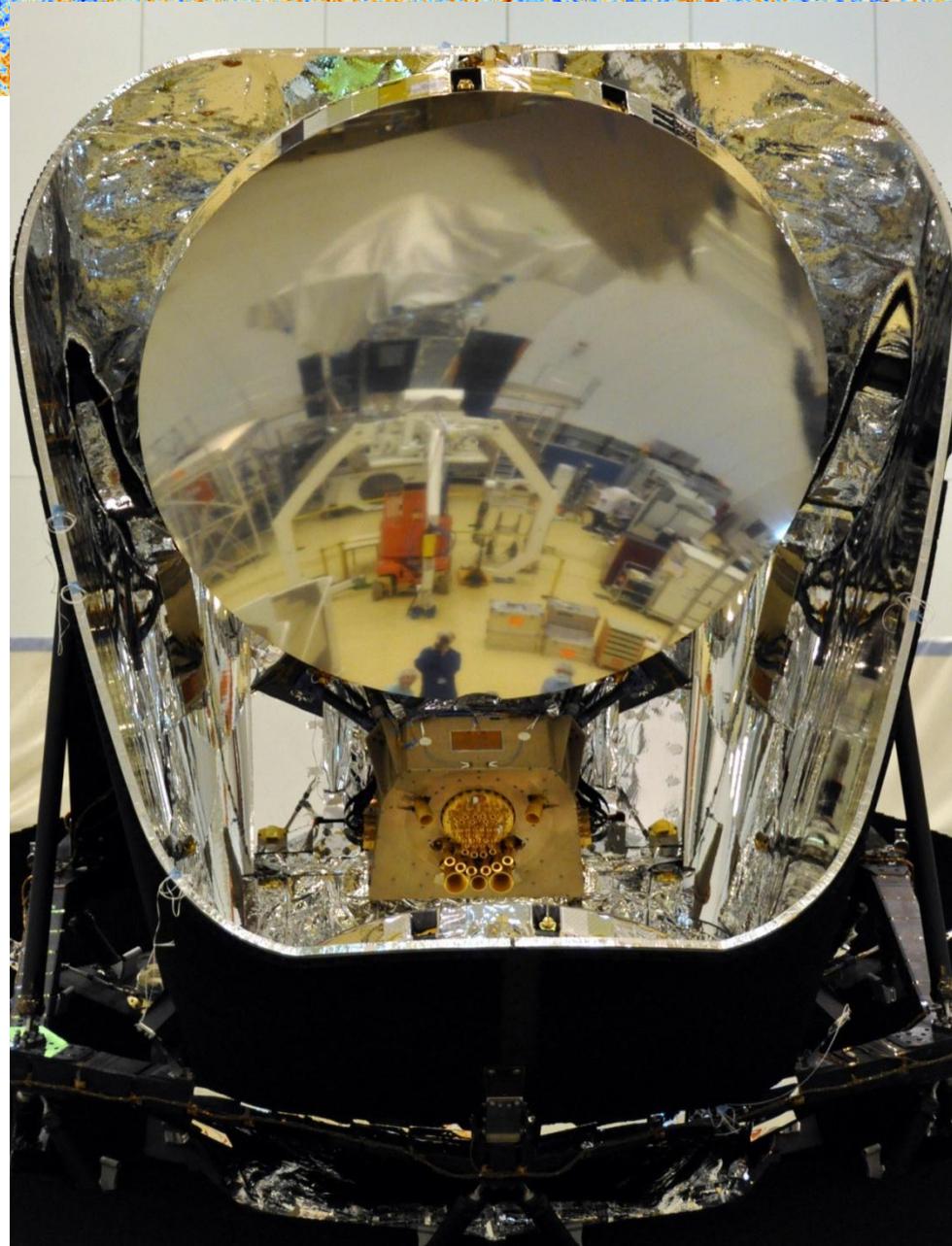
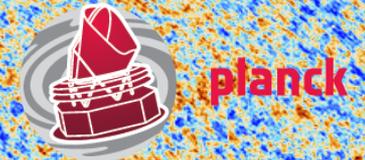
4 K Stirling cooler (RAL)

0.1 K <sup>3</sup>He/<sup>4</sup>He dilution cooler (CRTBT)

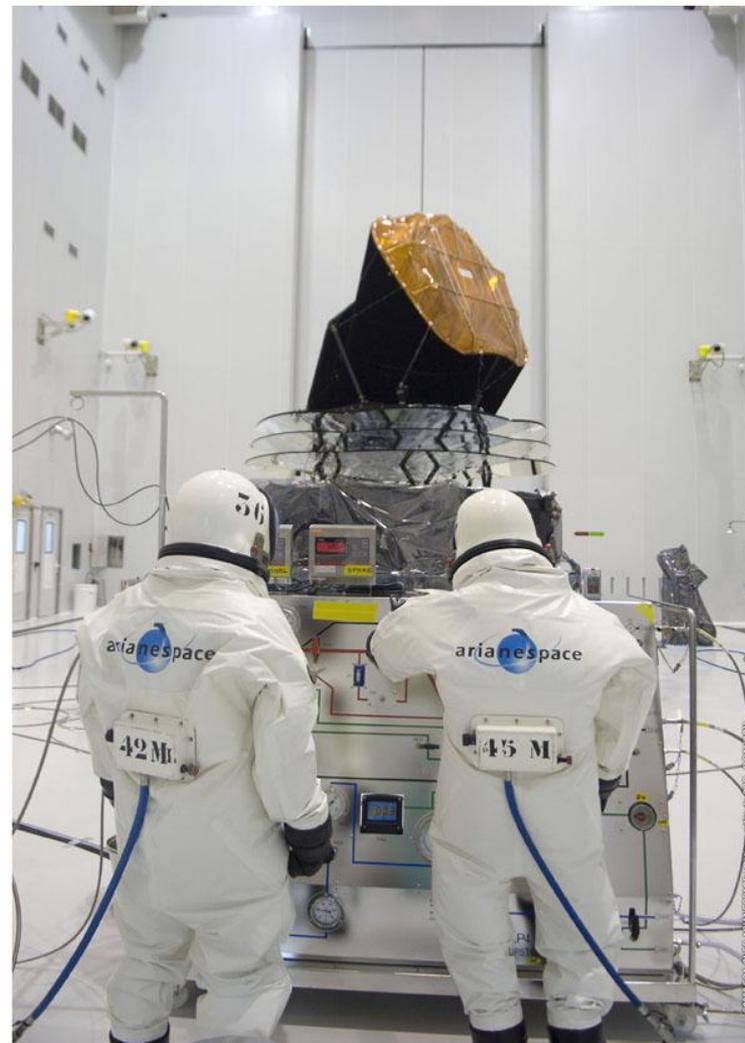
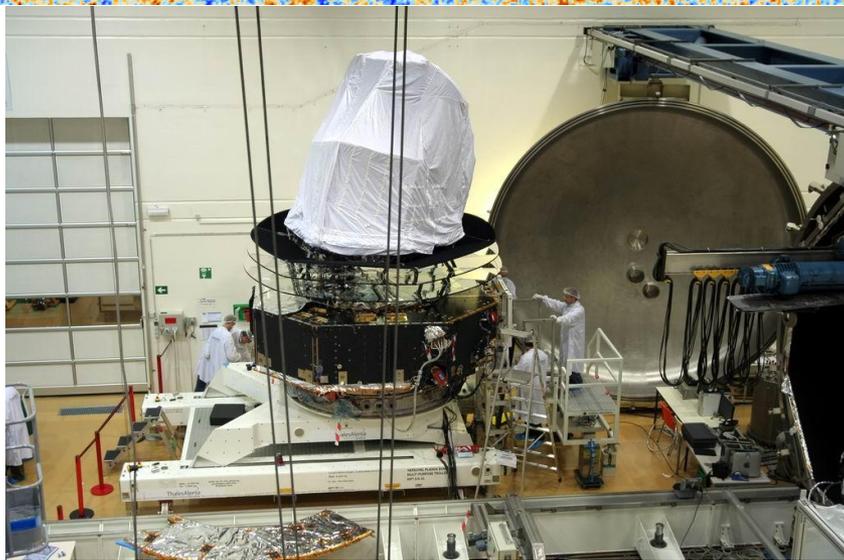
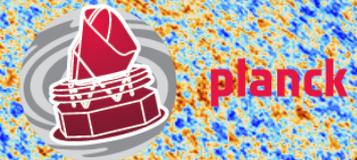
# Lo strumento Italiano e' stato integrato a Milano



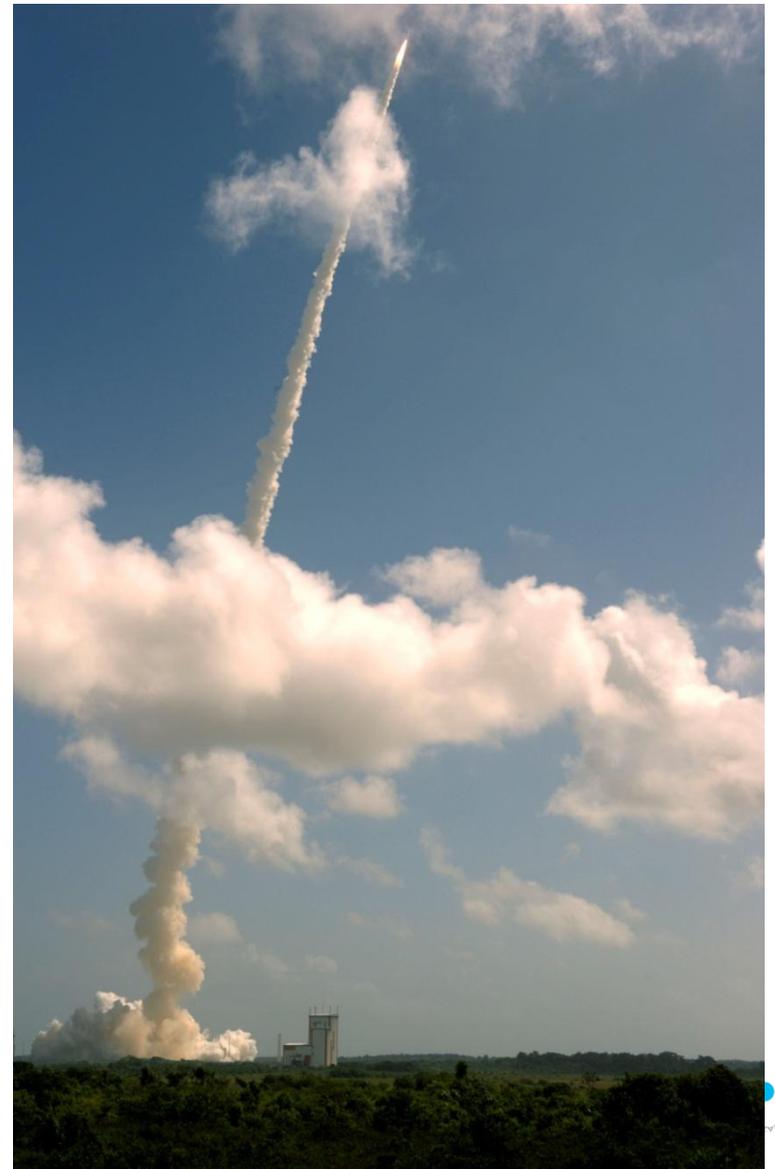
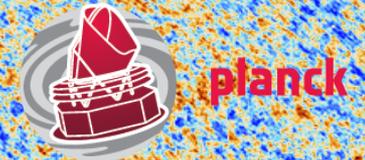
# L'Occhio di Planck



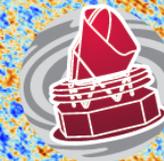
# Ultimi test e Planck fa il pieno ...



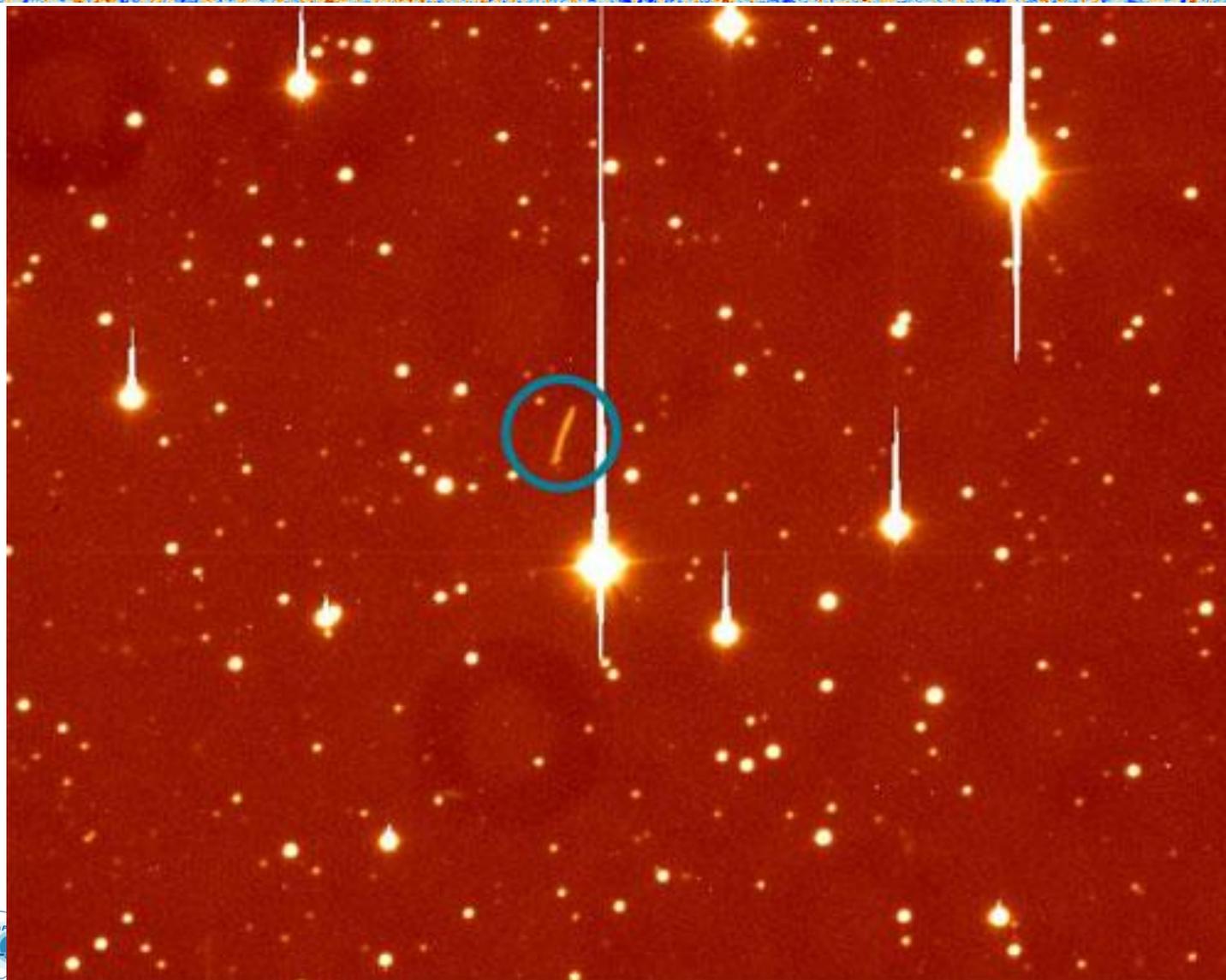
# Il Lancio (14 Maggio 2009)



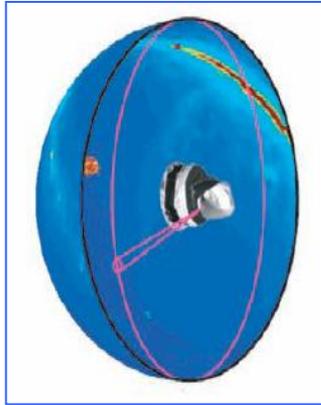
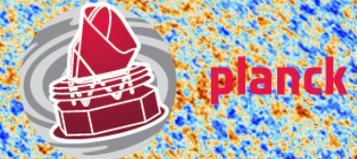
Ma siamo sicuri ???



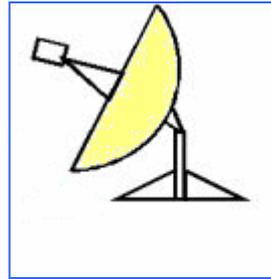
planck



# Acquisizione dei dati



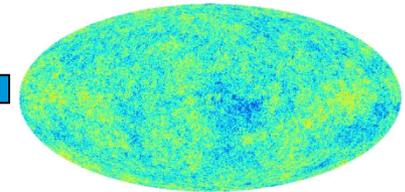
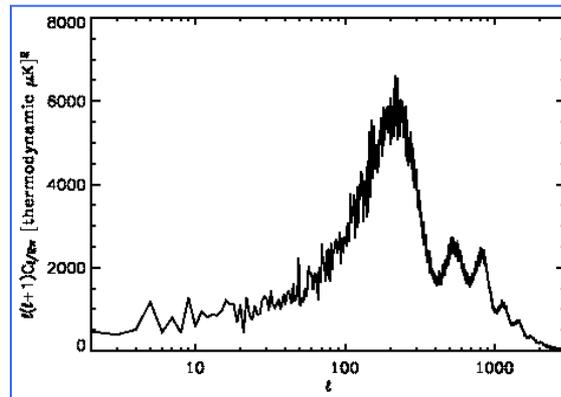
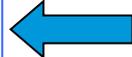
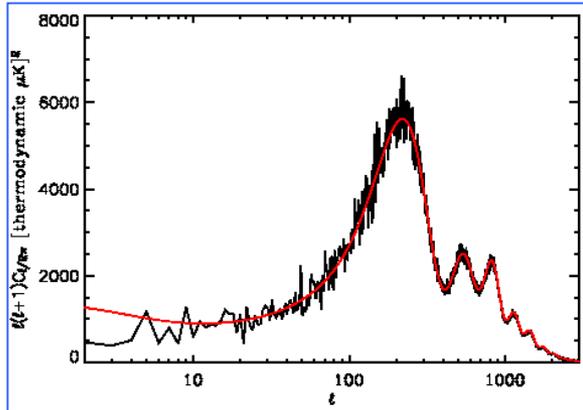
0110101  
1110110



MOC



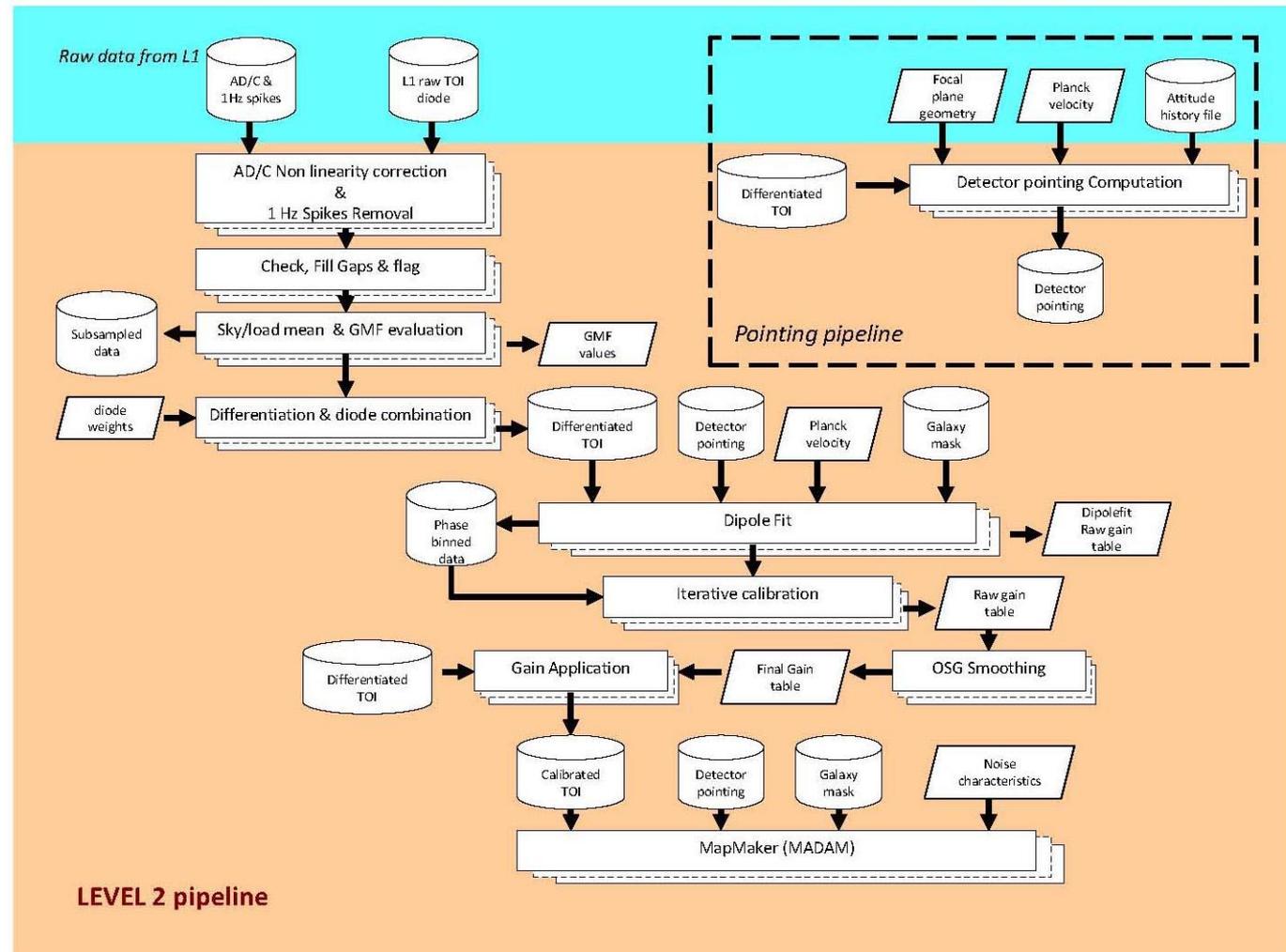
PARIGI E TRIESTE



# LFI Pipeline

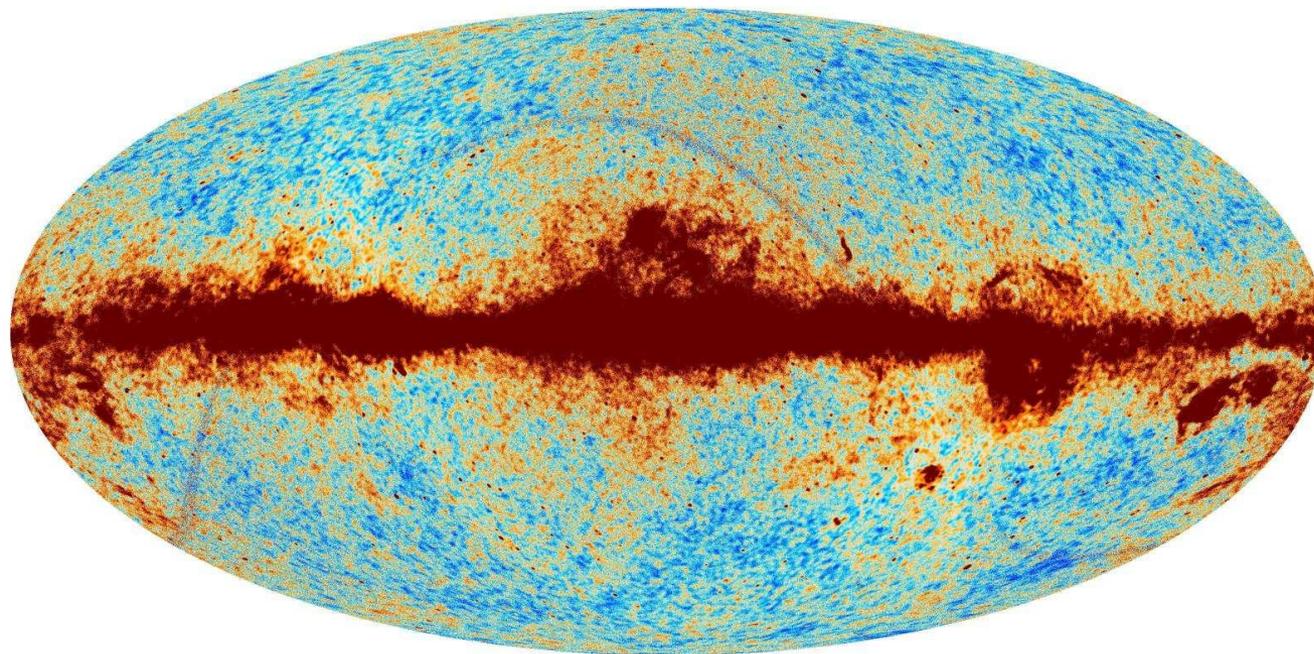
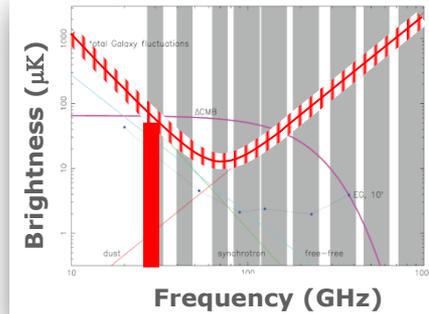
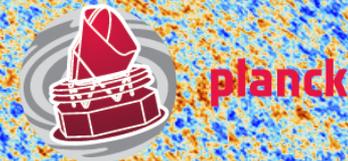
Tutto il software usato e' stato realizzato all'interno della collaborazione Planck.

Si tratta di centinaia di moduli di analisi scritte in C++ che lavorano su cluster(250/10000 CPU) in ambiente linux.



LEVEL 2 pipeline

# Planck-LFI – 30 GHz frequency map

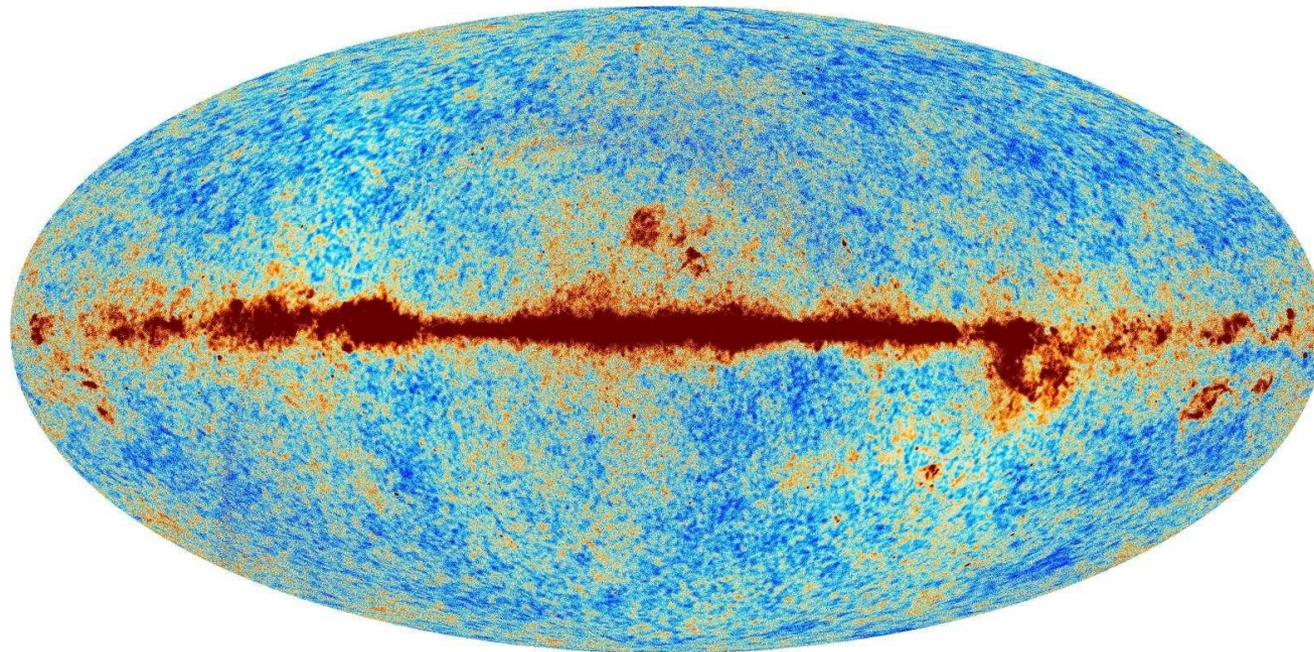
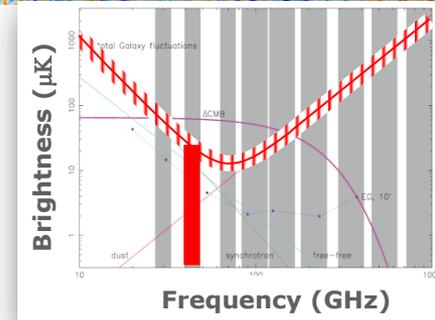
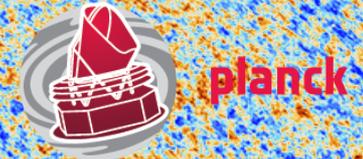


-250 500  $\mu\text{K}_{\text{CMB}}$

A. Zacchei "Frequency maps generation"



# Planck-LFI – 44 GHz frequency map

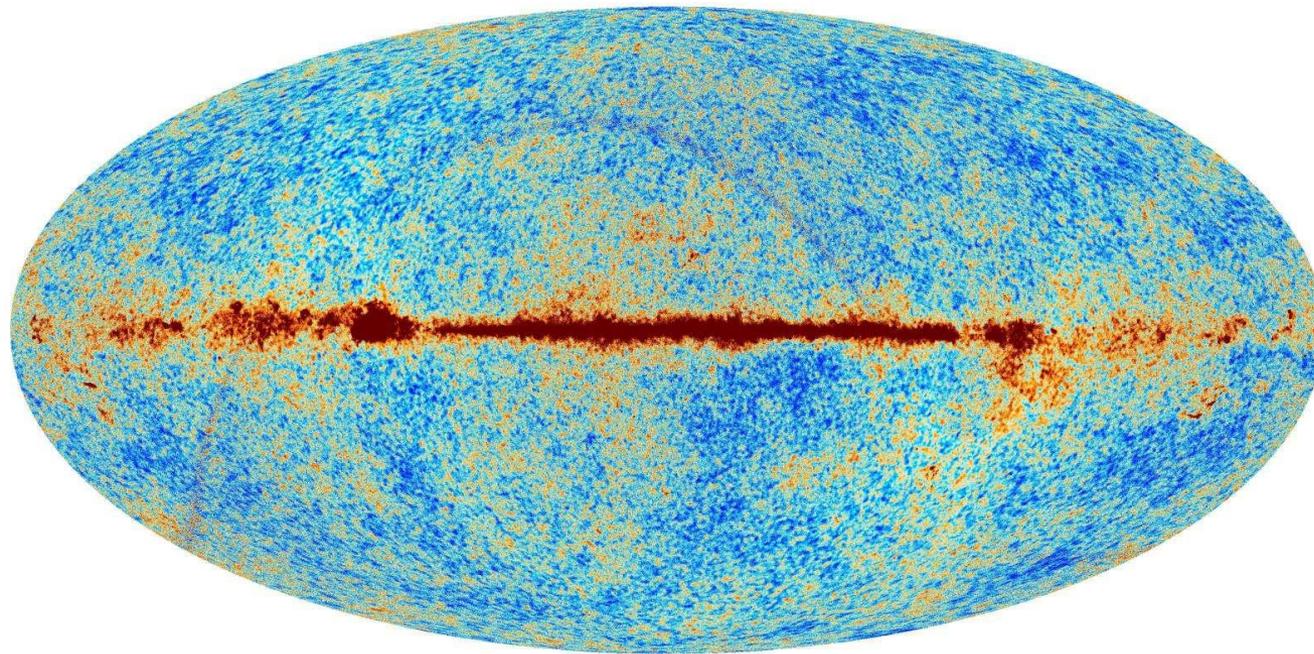
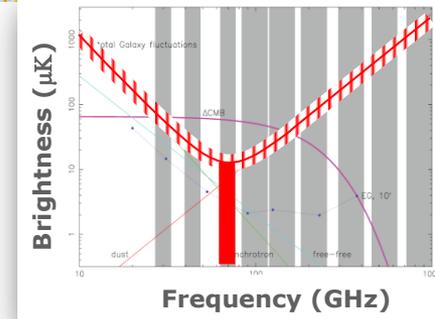
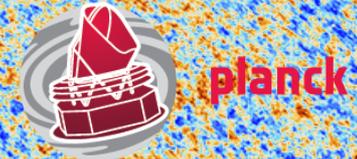


-250 500  $\mu\text{K}_{\text{CMB}}$

A. Zacchei "Frequency maps generation"



# Planck-LFI – 70 GHz frequency map

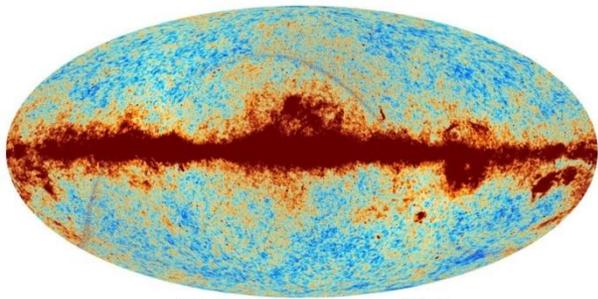
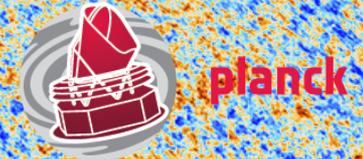


-250  500  $\mu\text{K}_{\text{CMB}}$

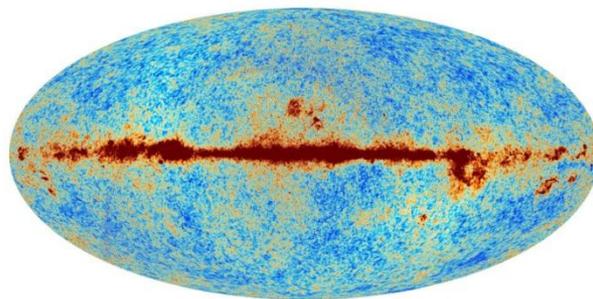
A. Zacchei "Frequency maps generation"



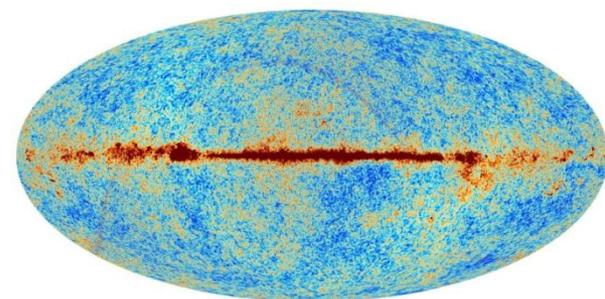
# Planck le Mappe in frequenza



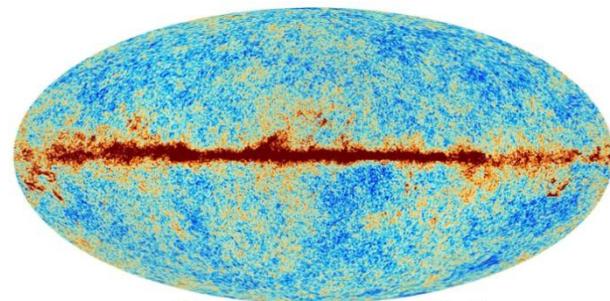
-250 500  $\mu\text{K}_{\text{CMB}}$



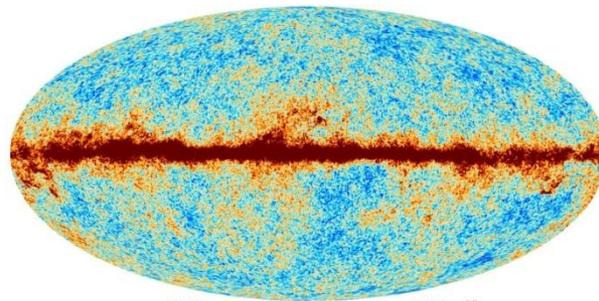
-250 500  $\mu\text{K}_{\text{CMB}}$



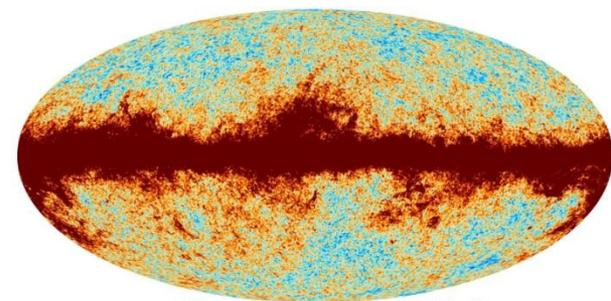
-250 500  $\mu\text{K}_{\text{CMB}}$



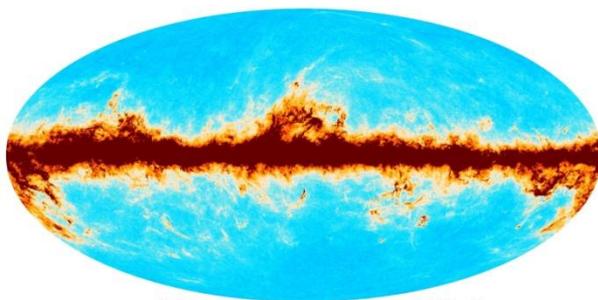
-250 500  $\mu\text{K}_{\text{CMB}}$



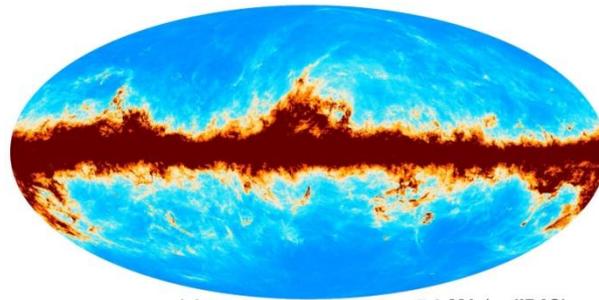
-250 500  $\mu\text{K}_{\text{CMB}}$



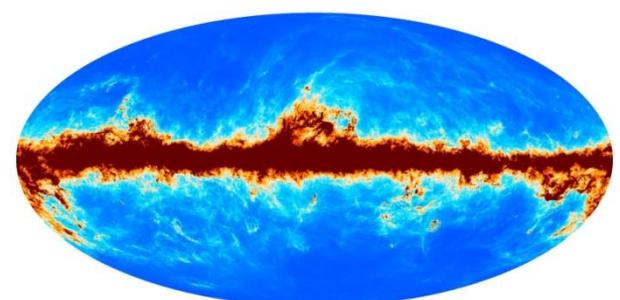
-250 500  $\mu\text{K}_{\text{CMB}}$



-2500 7500  $\mu\text{K}_{\text{CMB}}$

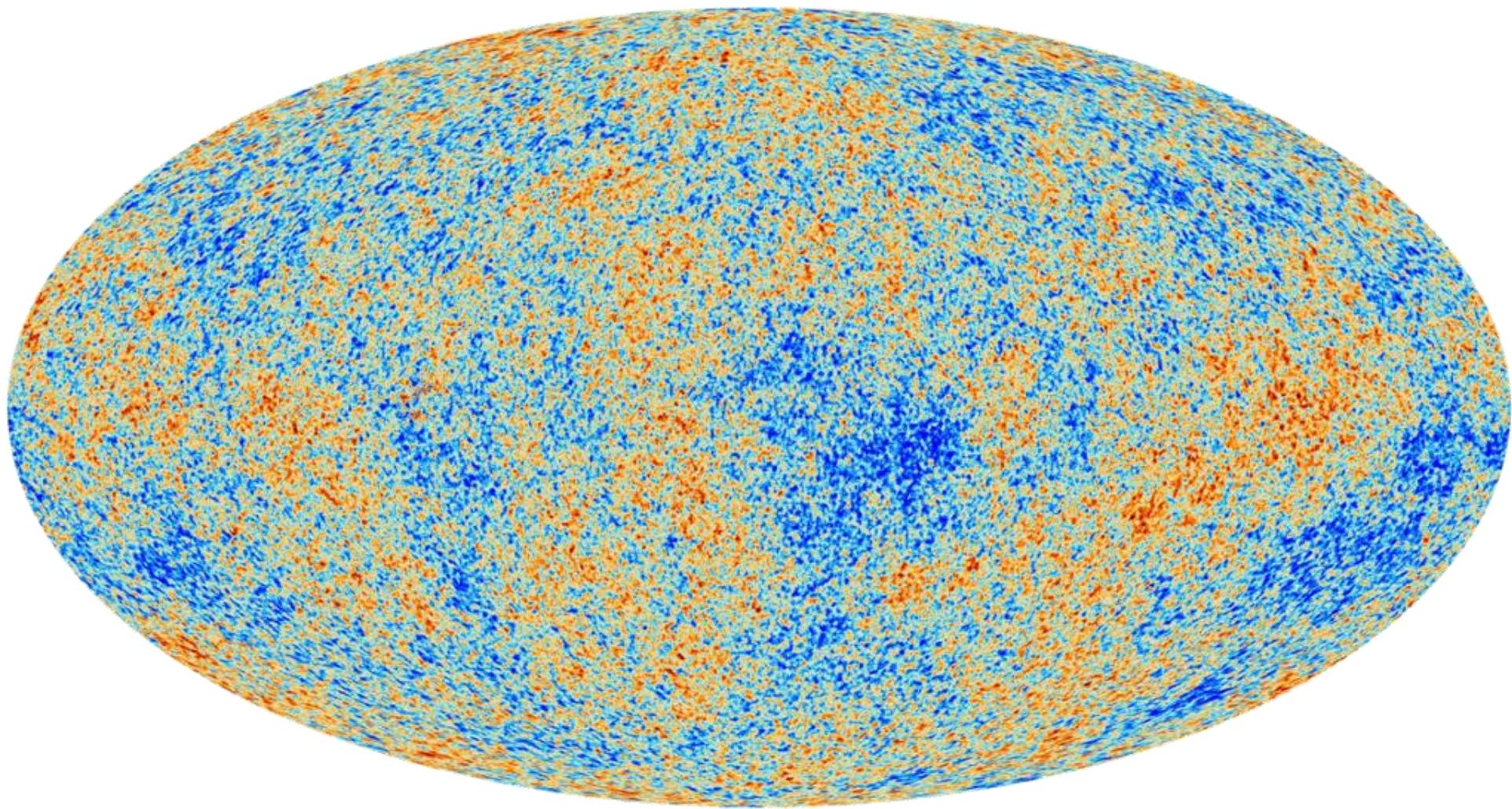
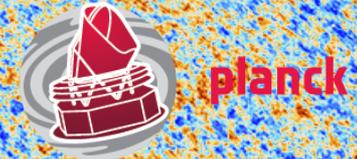


-1.0 5.0 MJy/sr (IRAS)

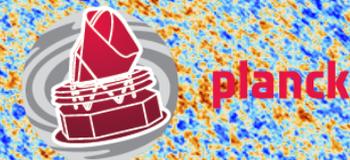


-2.0 20.0 MJy/sr (IRAS)

# Planck CMB Map



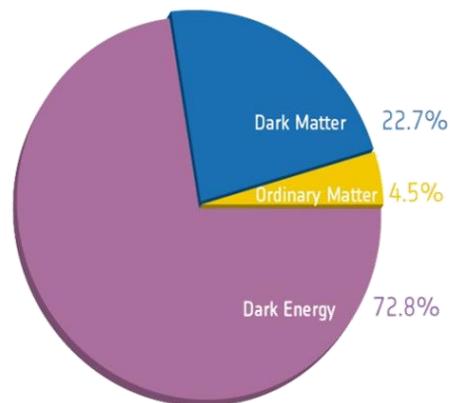
# I primi risultati: di cosa e' fatto l'universo ?



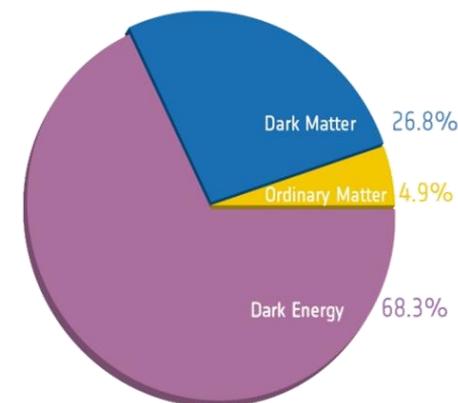
➤ Eta' dell'universo:  $13.82 \times 10^9$  anni

## ➤ Composizione

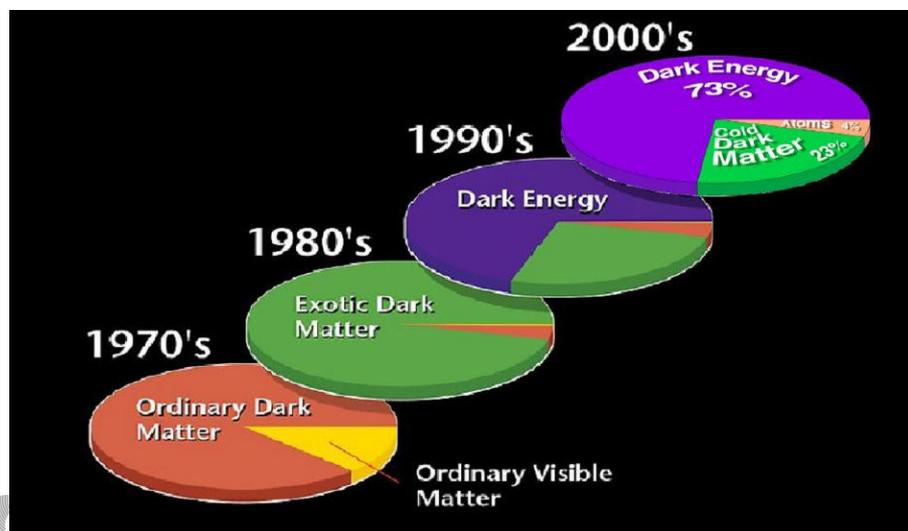
- 4.9% Materia ordinaria
- 26.8% Materia Oscura
- 58.3% Energia Oscura



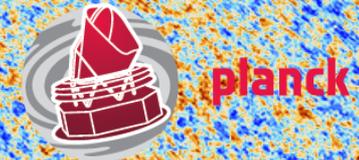
Before Planck



After Planck



# I primi risultati: di cosa e' fatto l'universo ?

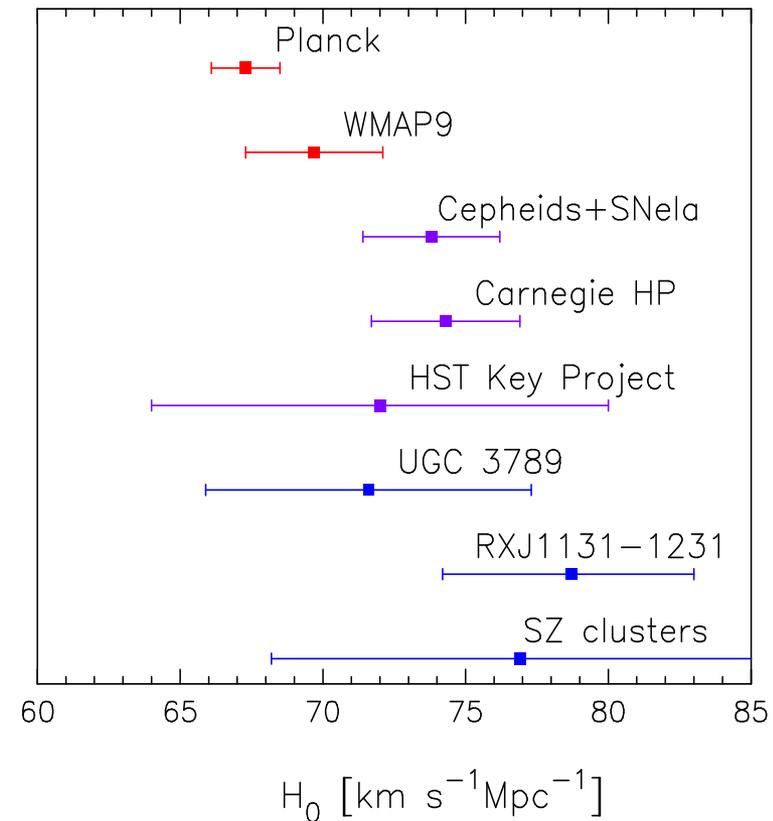


## Geometria dell'Universo

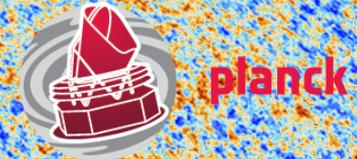
- $\Omega_0=1$ : Universe e' piatto, in continua espansione

## Hubble constant (velocita' di espansione):

- $H_0 = 67.3 \pm 1.2 \text{ km s}^{-1}\text{Mpc}^{-1}$
- Prima di Planck era  $72 \pm 8 \text{ km s}^{-1}\text{Mpc}^{-1}$



# Domande ancora aperte....

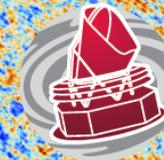
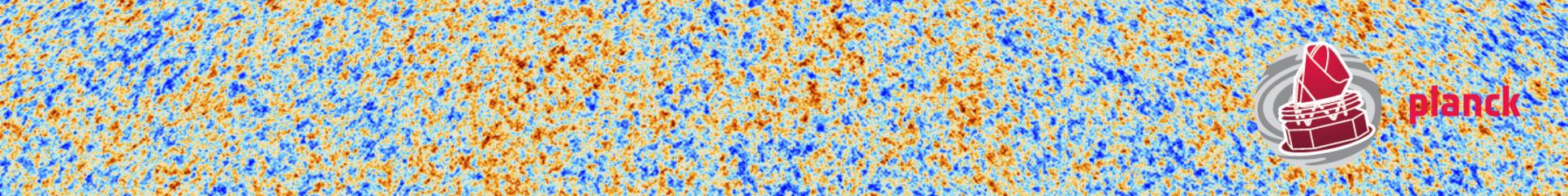


- Che cosa e' la Materia Oscura?
- Che cosa e' l'energia Oscura
- Perche' proprio nel momento in cui viviamo l'accelerazione dell'Universo sta aumentando ?
- Quanto pesano i neutrini ?

# The scientific results that we present today are a product of the Planck Collaboration, including individuals from more than 100 scientific institutes in Europe, the USA and Canada



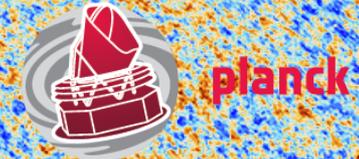
Planck is a project of the European Space Agency, with instruments provided by two scientific Consortia funded by ESA member states (in particular the lead countries: France and Italy) with contributions from NASA (USA), and telescope reflectors provided in a collaboration between ESA and a scientific Consortium led and funded by Denmark.



planck



# Domandone Finale ...



## Filosofico:

- Perche' l'ente Spaziale Europeo (ESA) e quindi NOI tutti spendiamo 800 Milioni di Euro per conoscere l'Universo ?

## Tecnico

- Perche' quasi tutti I siti di Lancio dei satelliti si trovano vicino all'Equatore e in zone costiere ?